

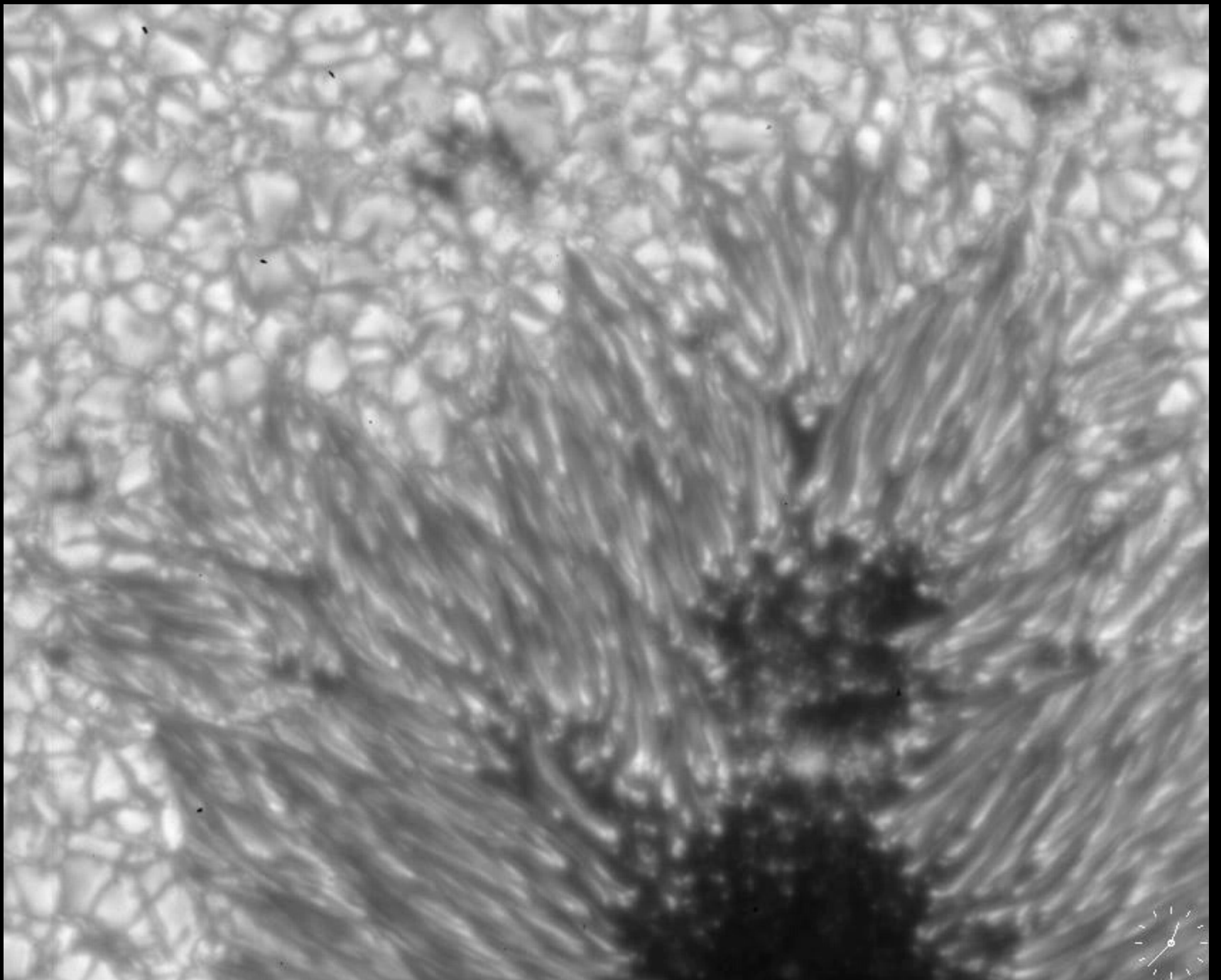
# **Convective nature of the Evershed Effect observed by SOT/Hinode**

Kiyoshi Ichimoto (Kyoto Univ.)

and

Hinode/SOT team

Hinode 2 in Boulder, 2008.9.30 – 10.4





*Filamentary structure of penumbra*

*Evershed flow = horizontal outflow along penumbral filaments*

Flow starts in bright filaments in inner penumbra,  
tends to flow in dark filaments in outer penumbra  
(Ichimoto et al, 2007, PASJ, 59, 593)



Issues addressed in this talk are,,

- what is the nature of the Evershed flow?
- what is the origin of the filamentary structure of the penumbra?

*Answers (?)..... 'flux tube model' or 'gap model'*

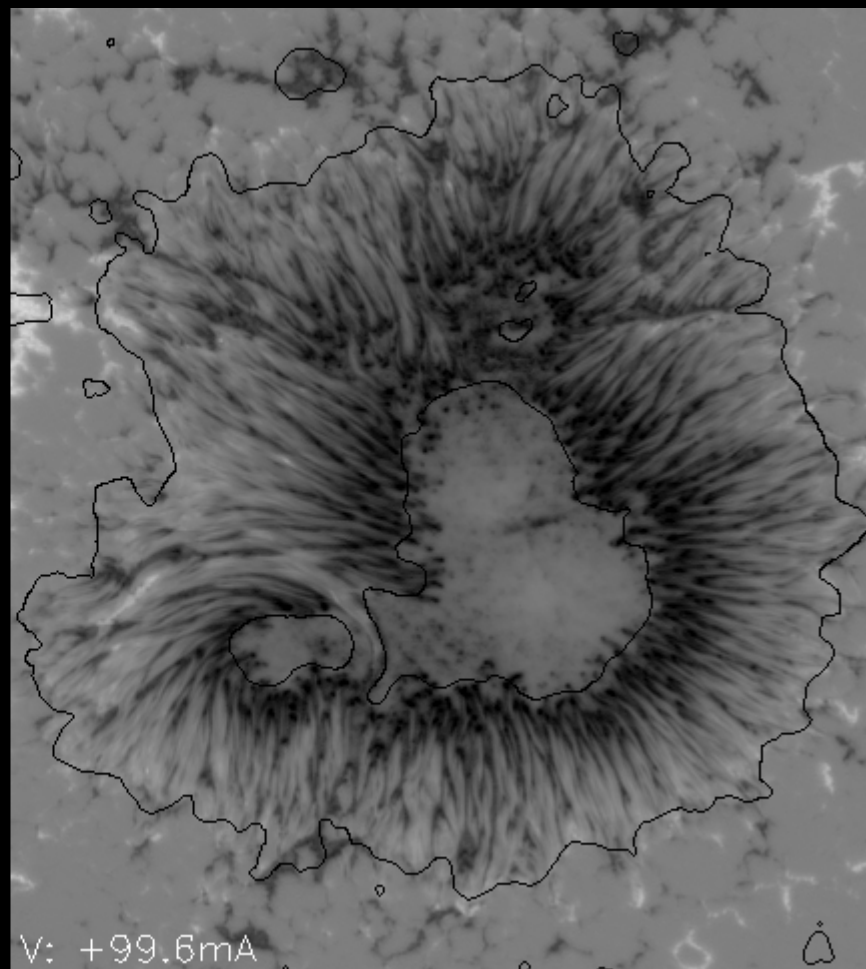
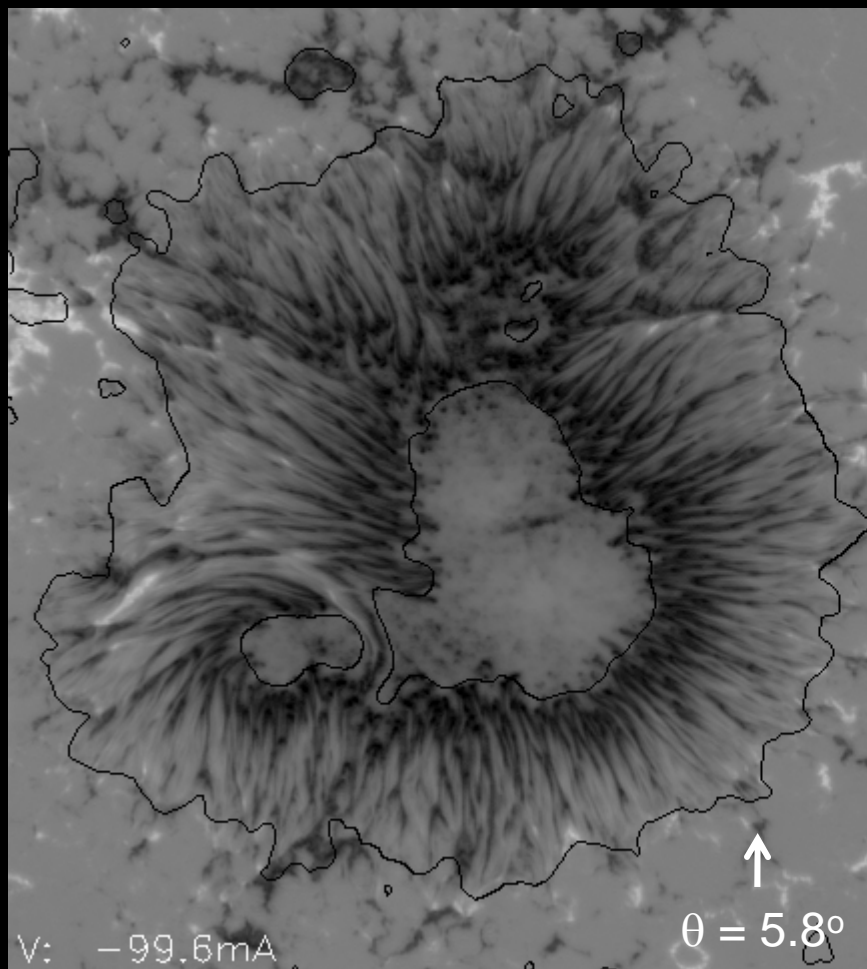
Let's start with a magnetogram of a sunspot at DC to see the 3D structure of Evershed flow.



2007.5.1

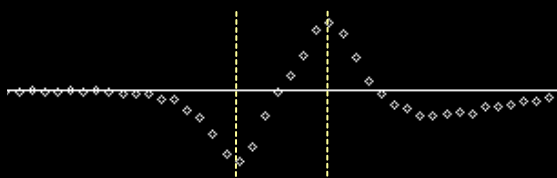
# Stokes-V at 6302.5A $\pm 100$ mA

(sign reversed)

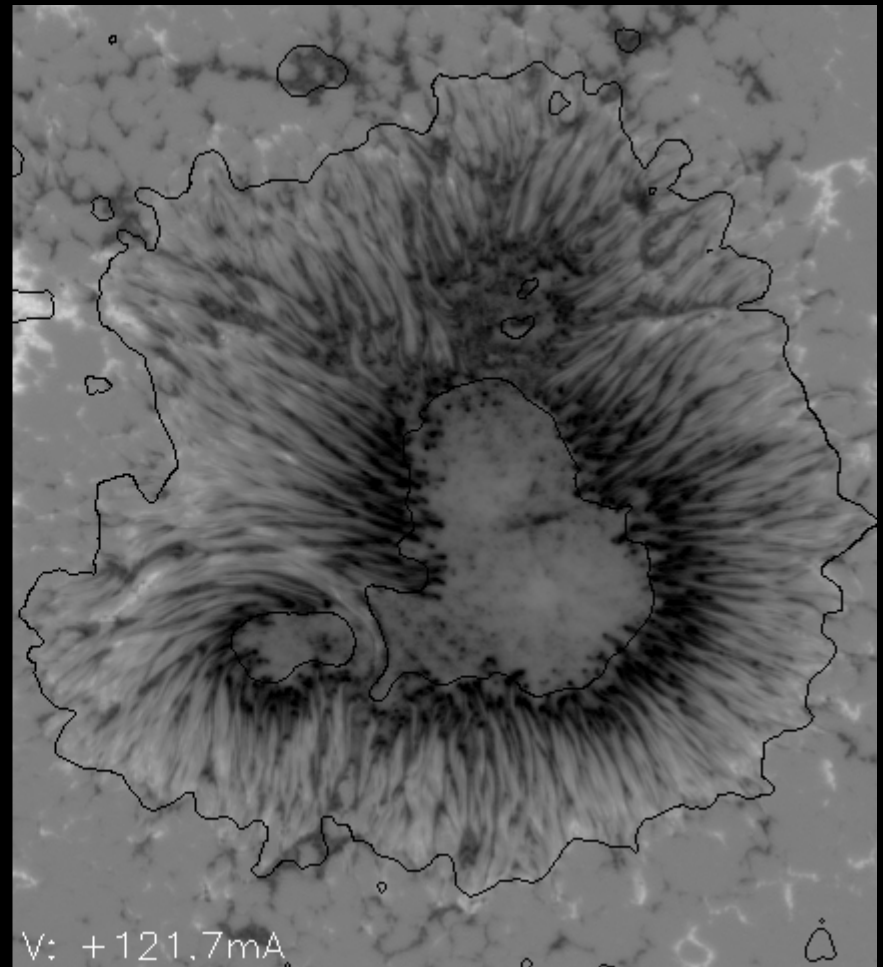
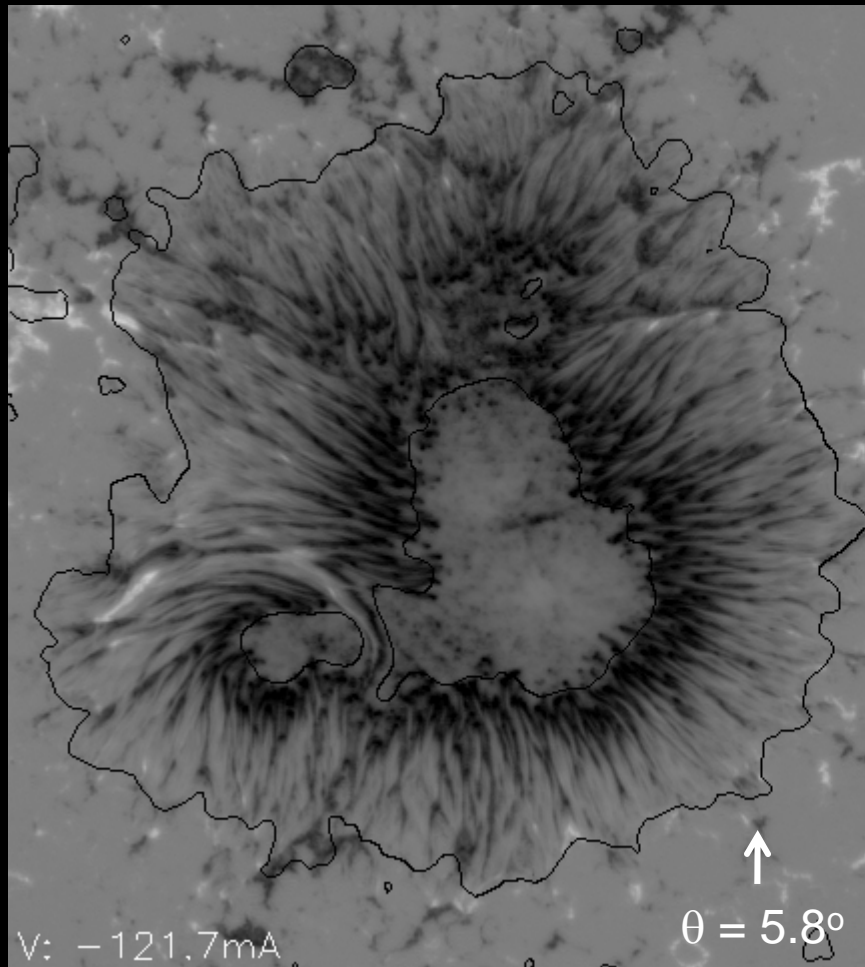


-100mA

+100mA

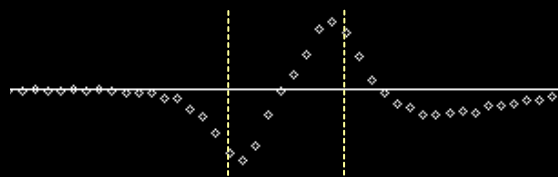


# Stokes-V at 6302.5Å $\pm 120\text{mA}$

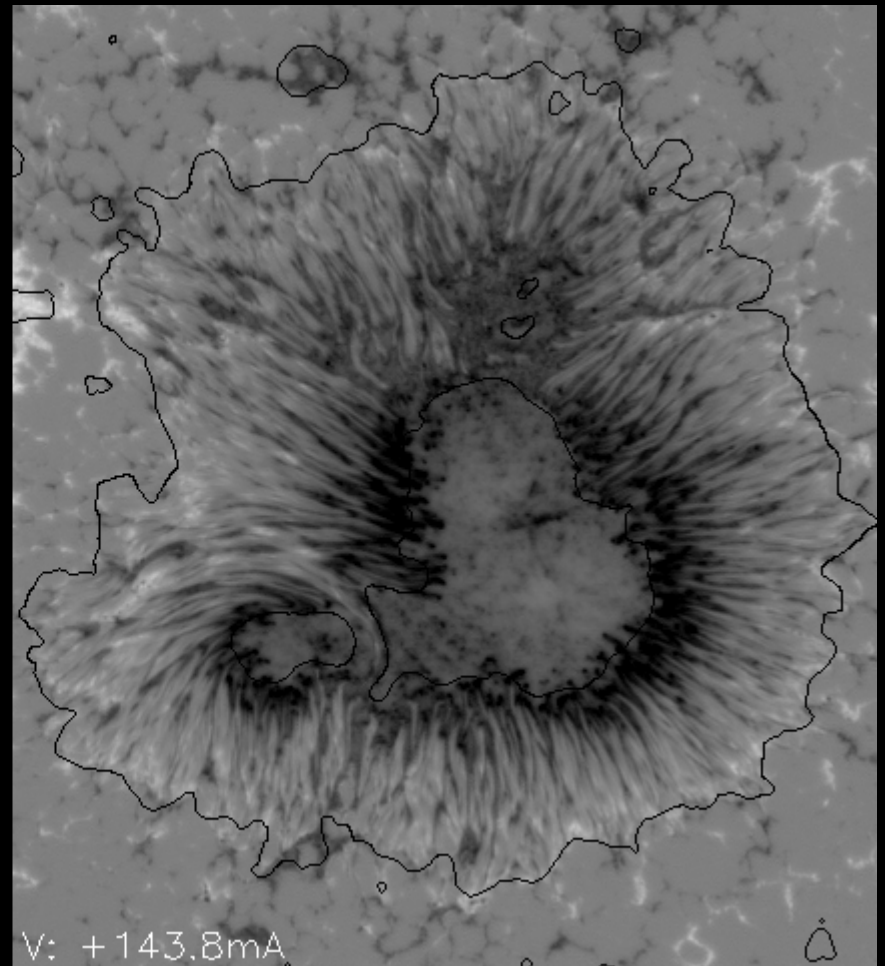
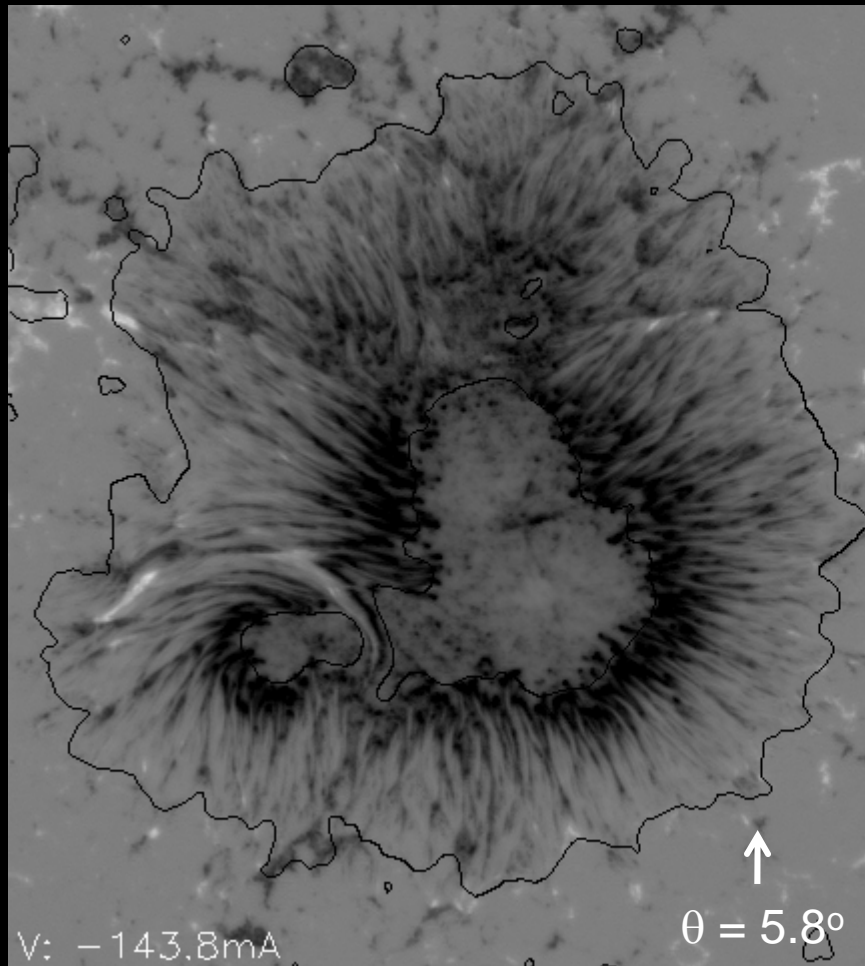


-120mA

+120mA

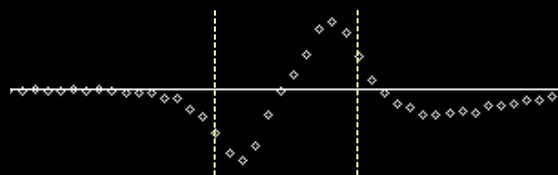


# Stokes-V at 6302.5A $\pm 144\text{mA}$



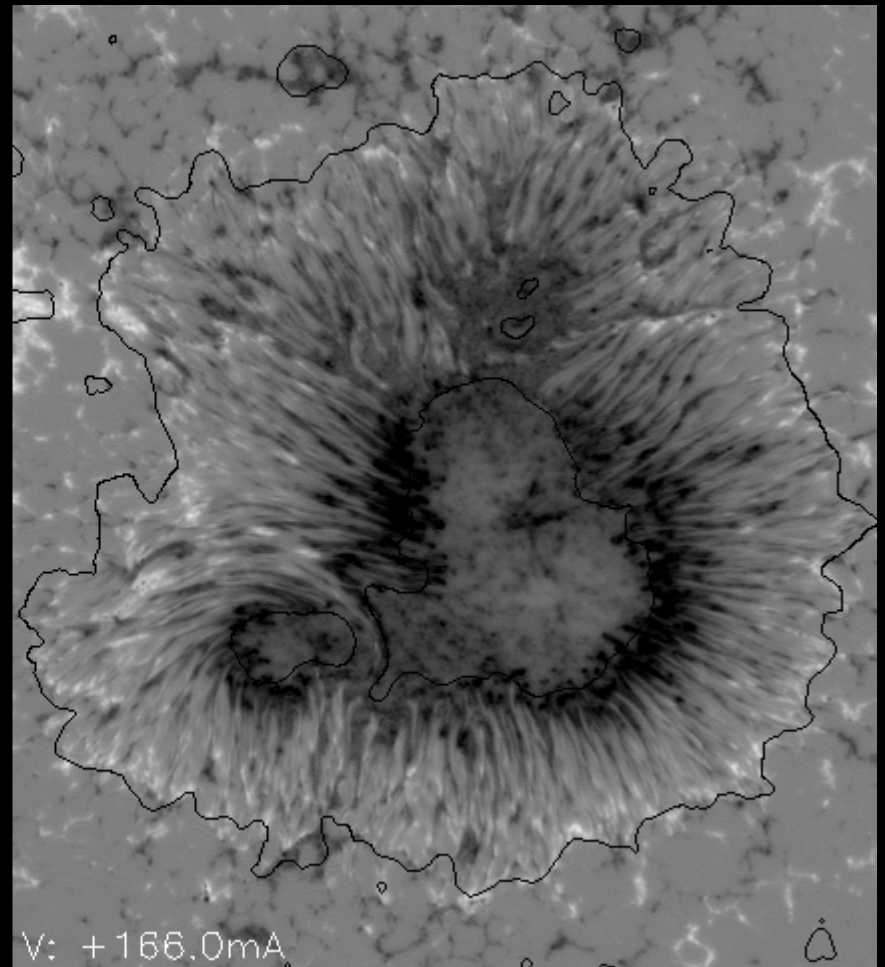
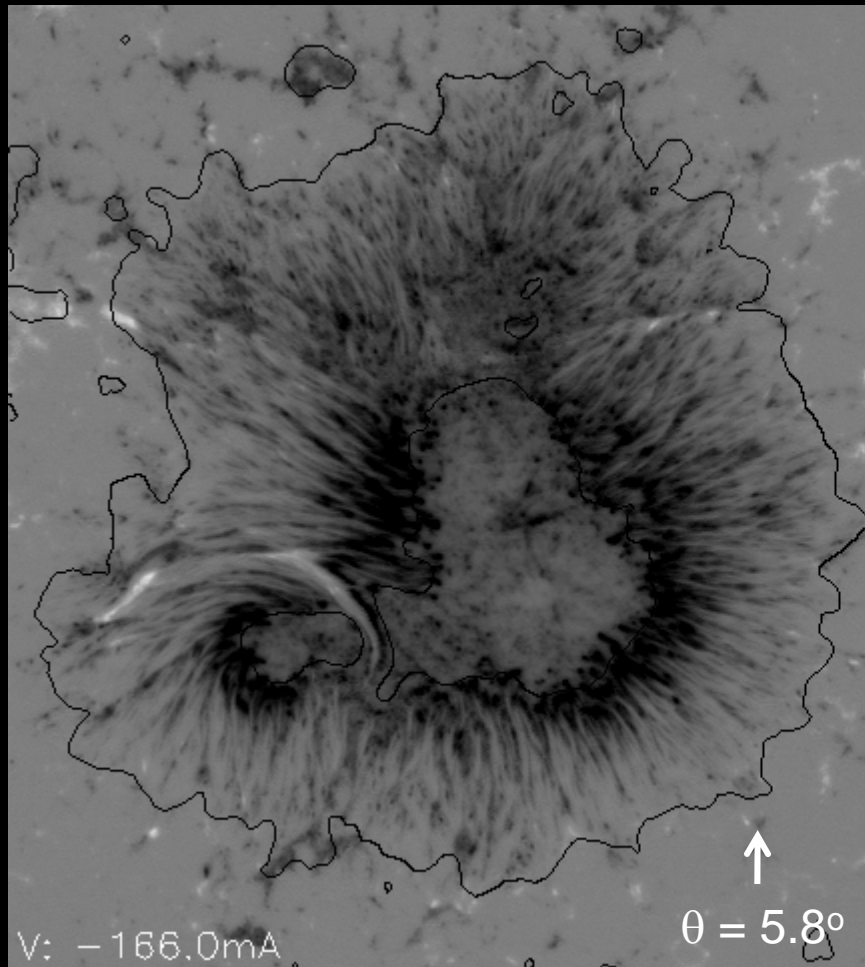
-144mA

+144mA



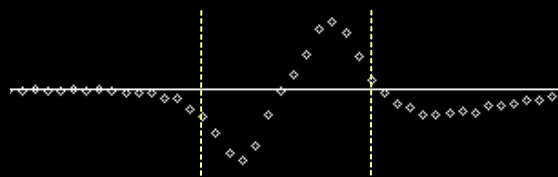


# Stokes-V at 6302.5A $\pm$ 166mA

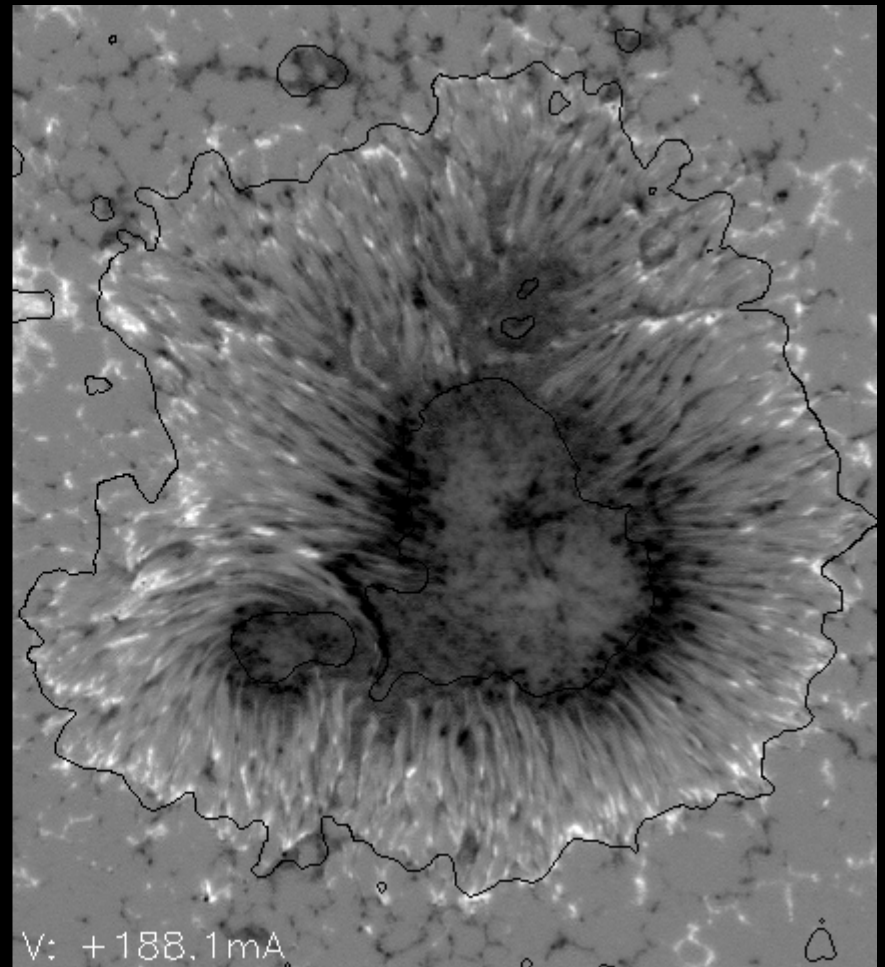
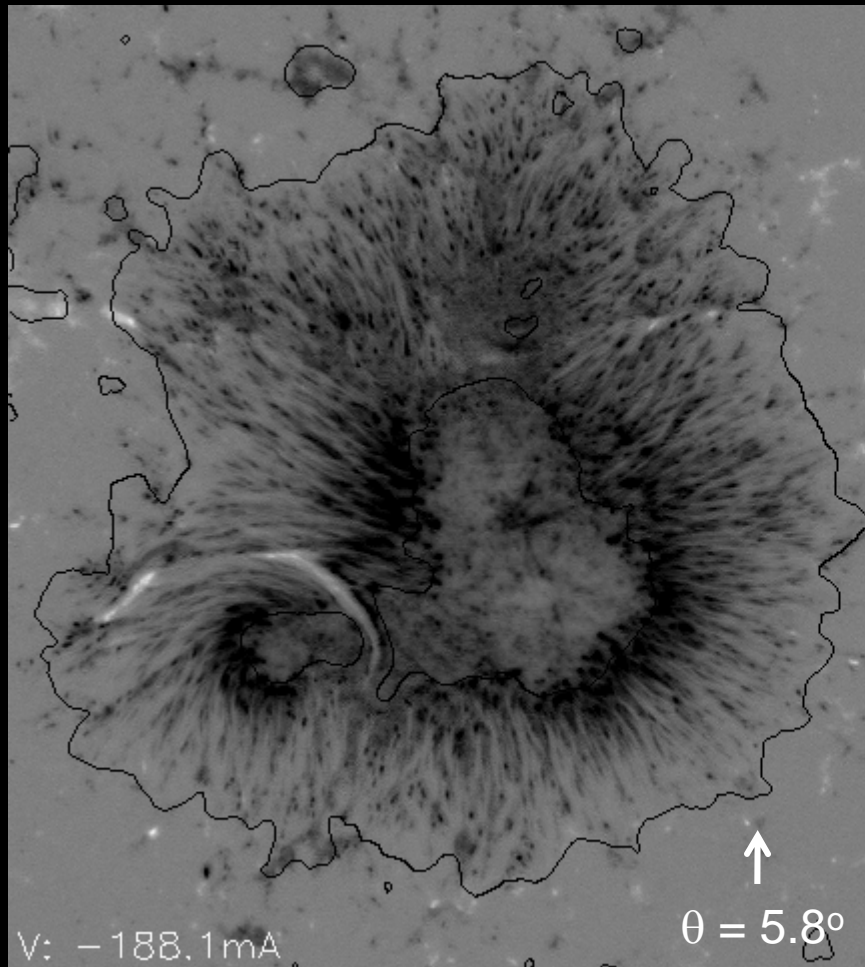


-166mA

+166mA

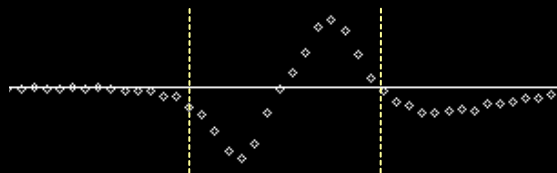


# Stokes-V at 6302.5Å $\pm$ 188mA

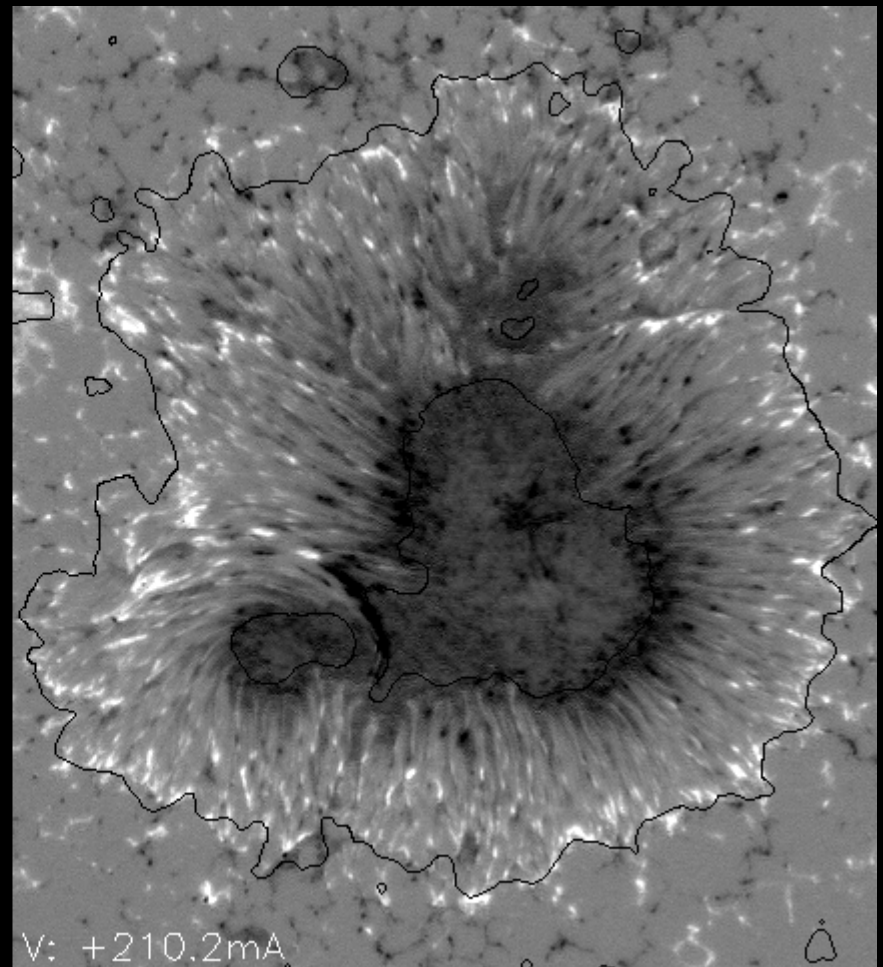
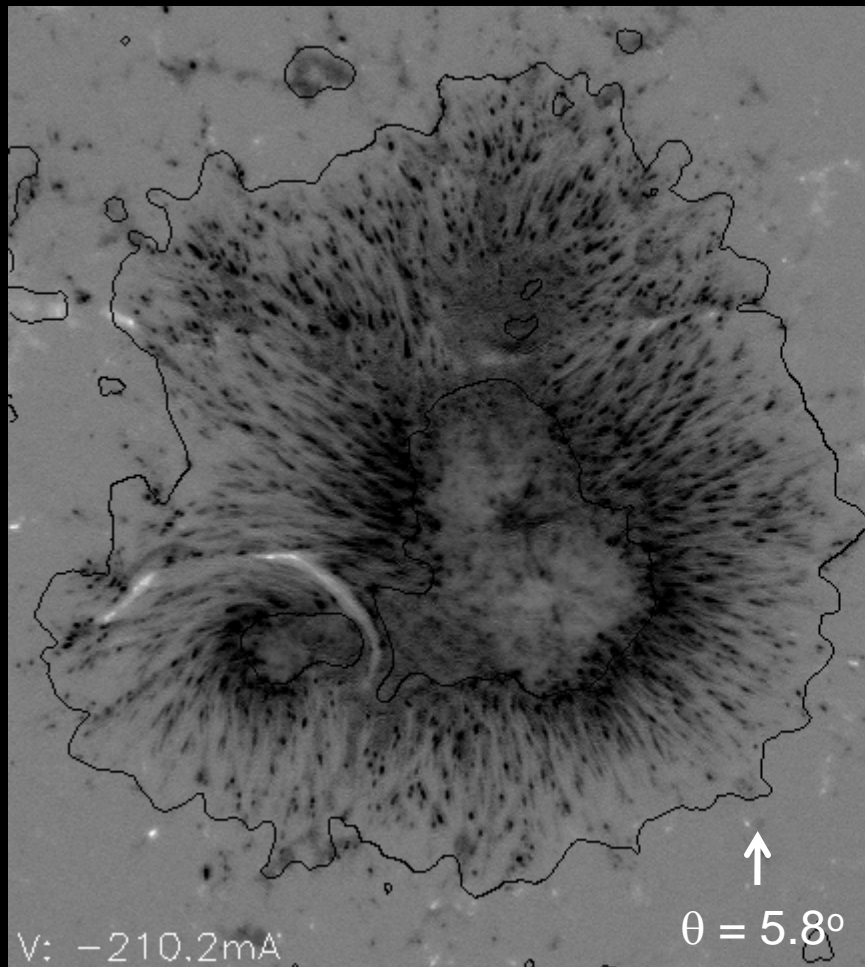


-188mA

+188mA

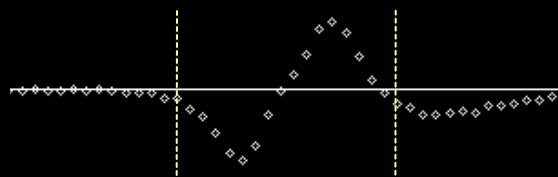


# Stokes-V at 6302.5Å ±210mA



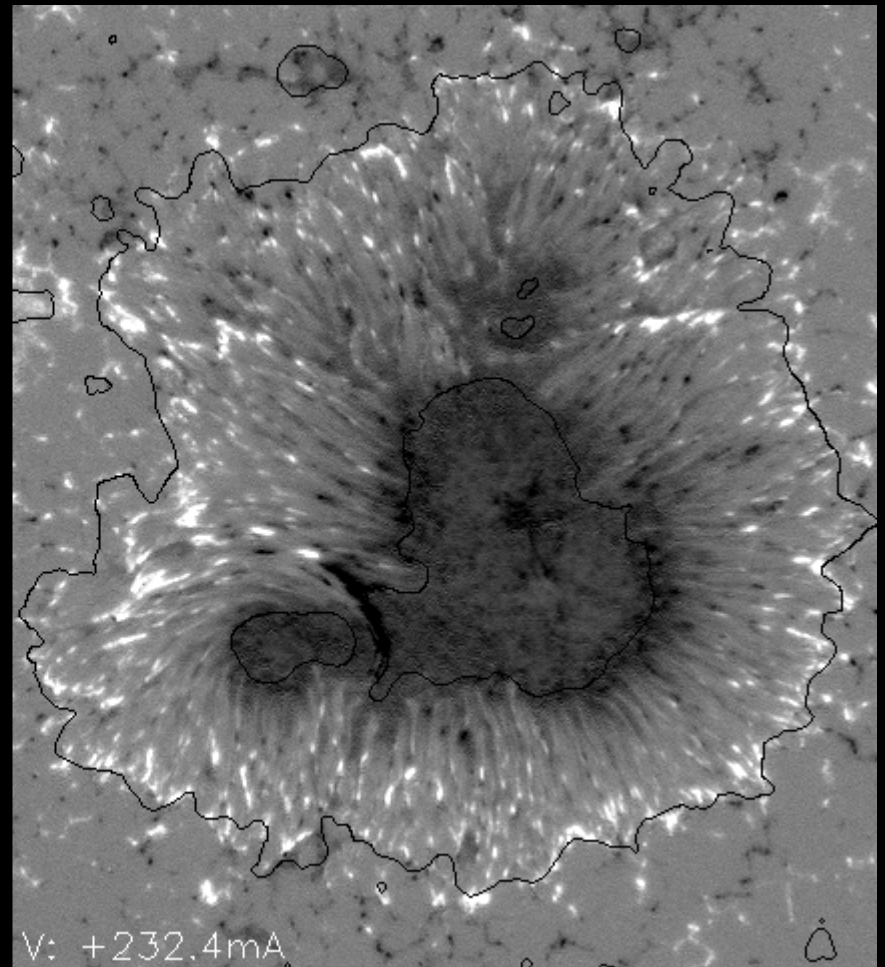
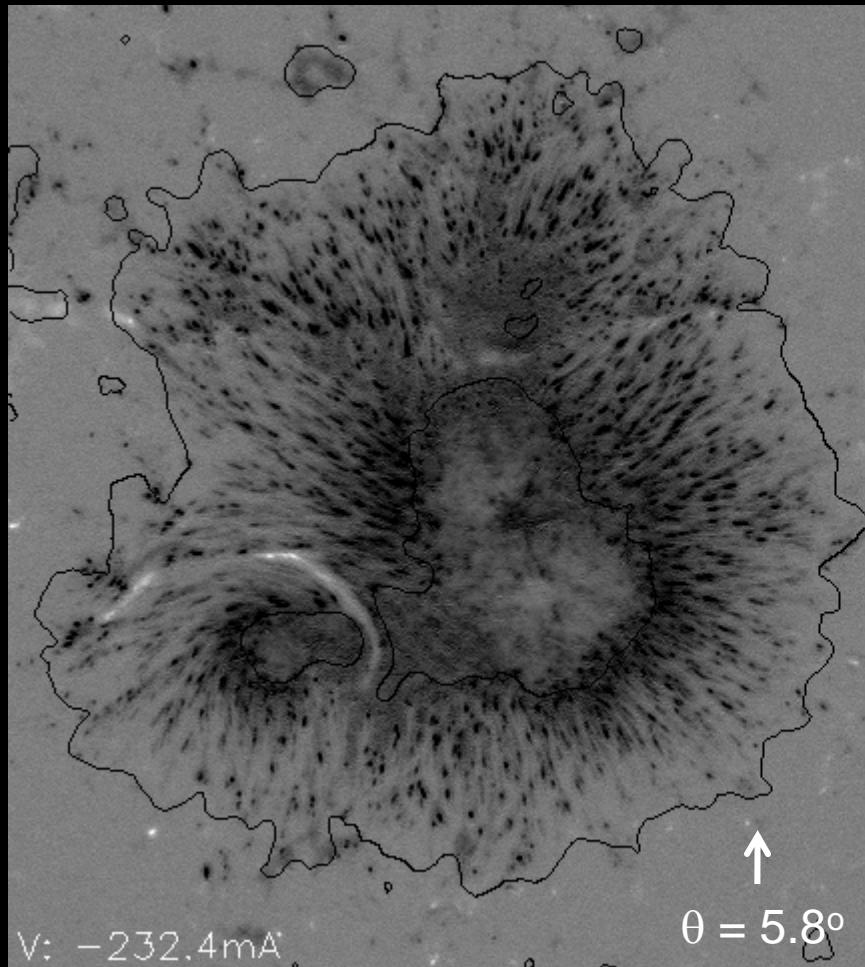
-210mA

+210mA

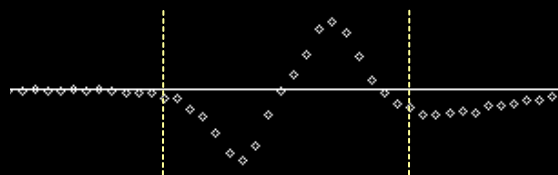




# Stokes-V at 6302.5A ±232mA



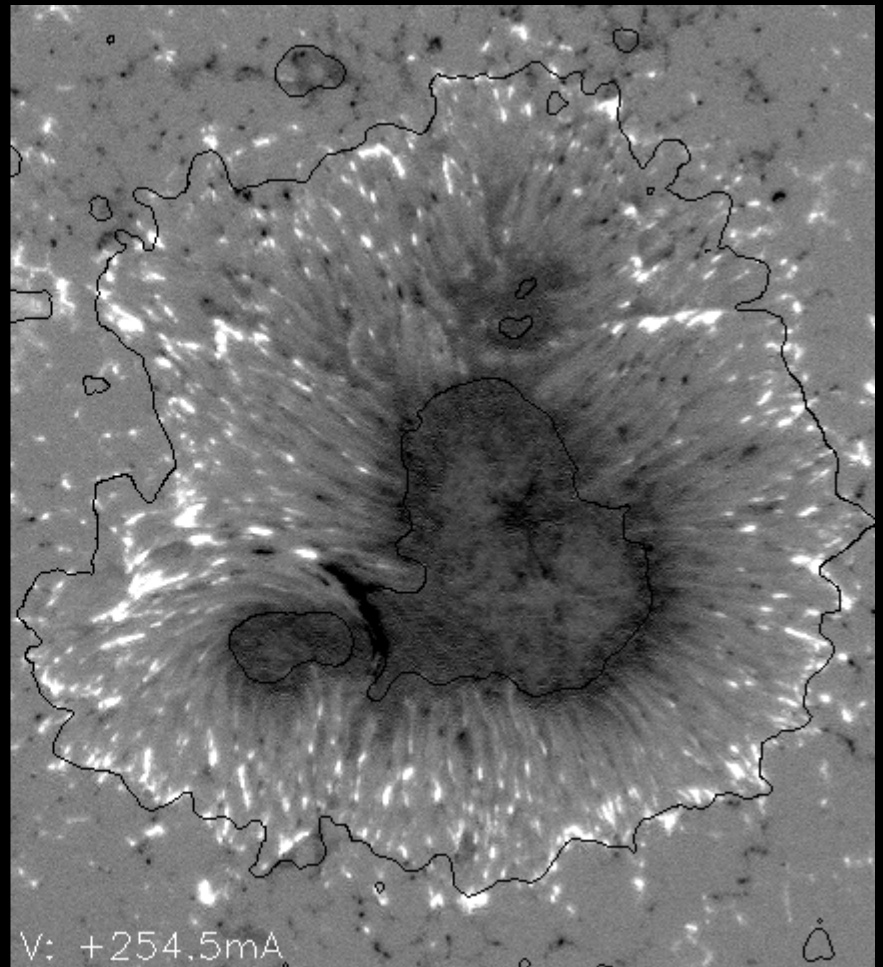
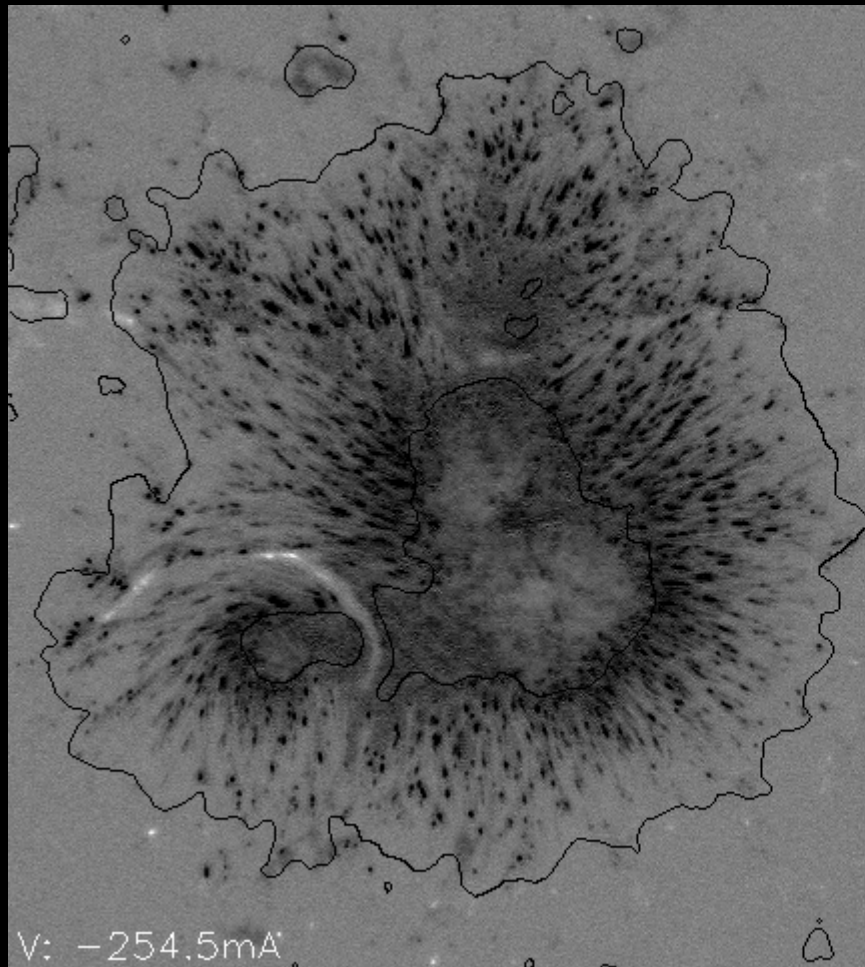
-232mA



+232mA

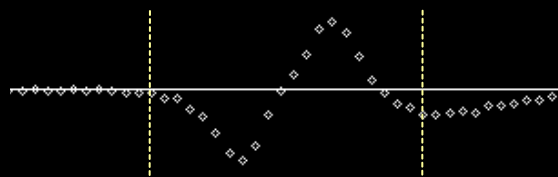


# Stokes-V at 6302.5A $\pm$ 254mA

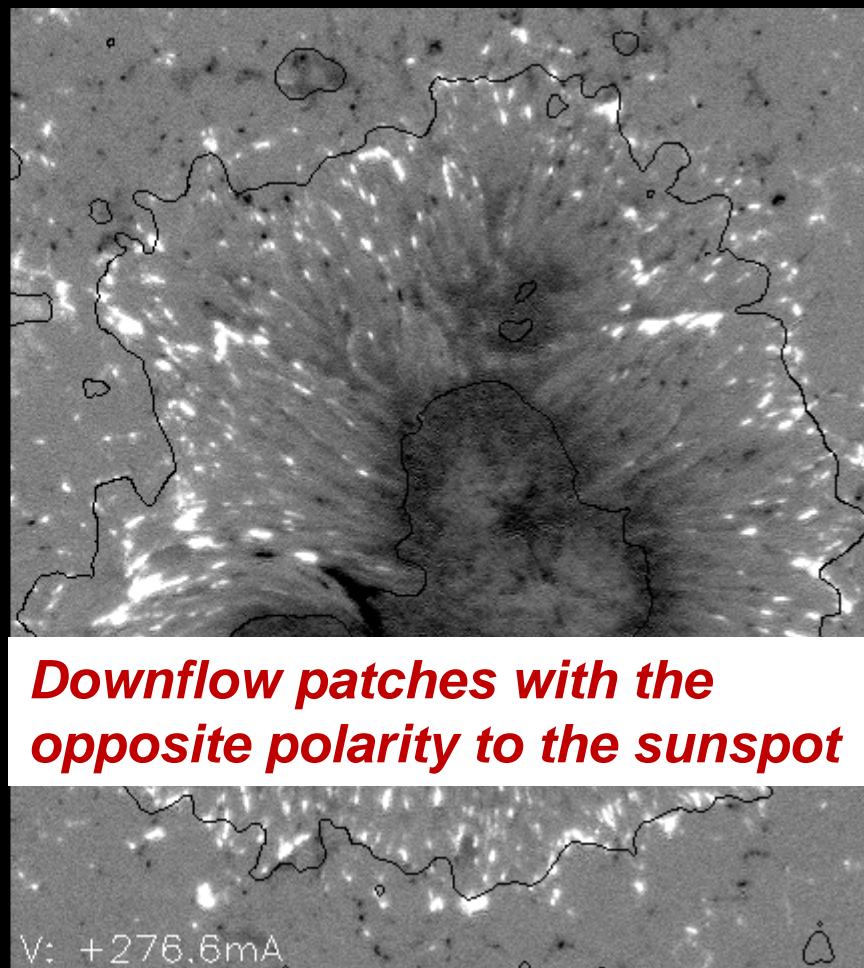
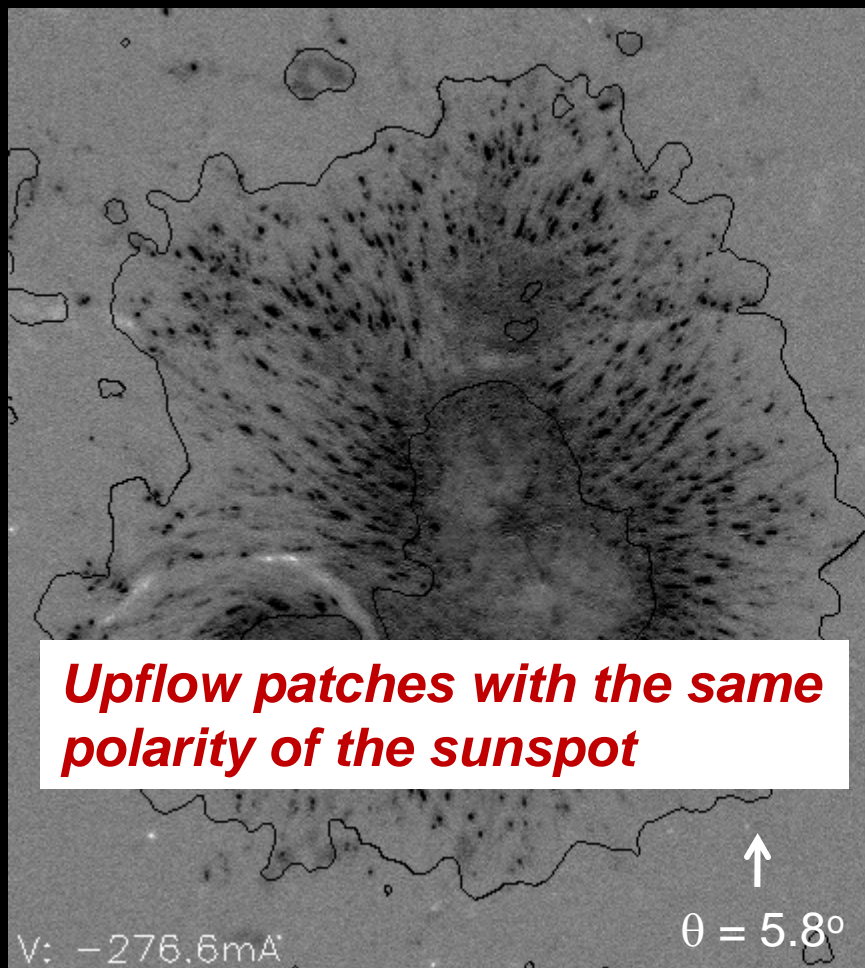


-254mA

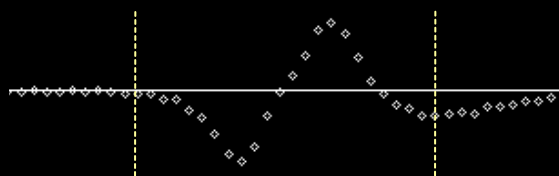
+254mA



# Stokes-V at 6302.5Å $\pm 277$ mA



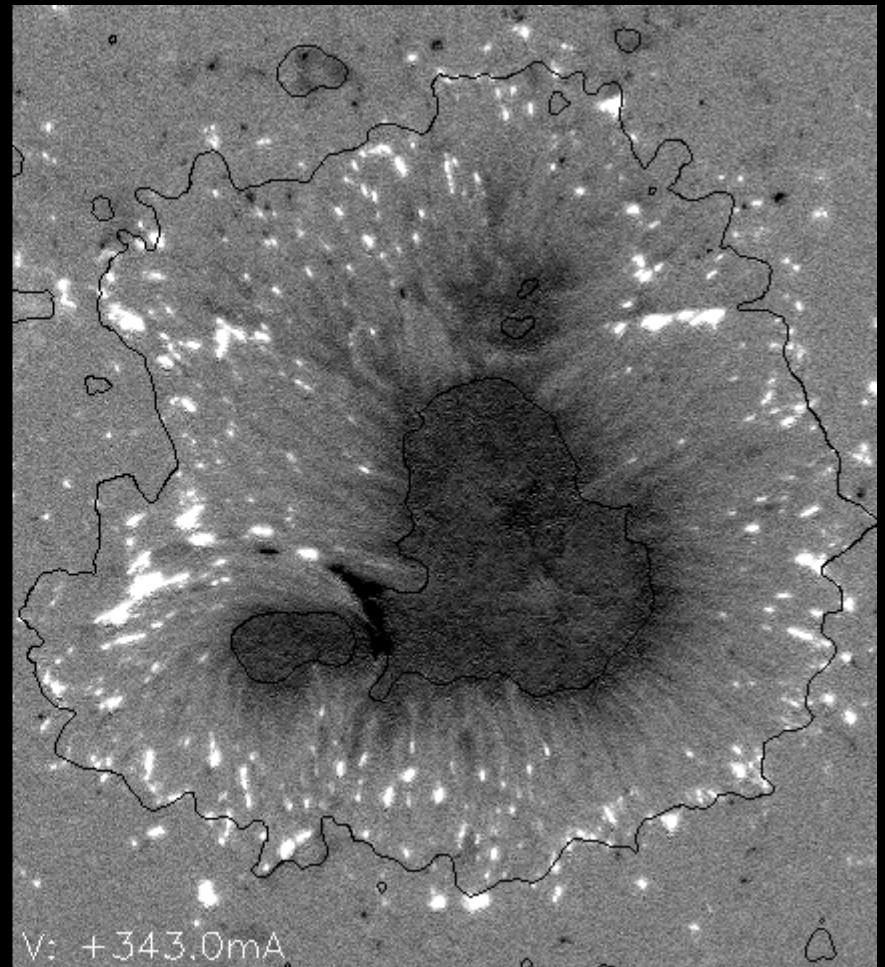
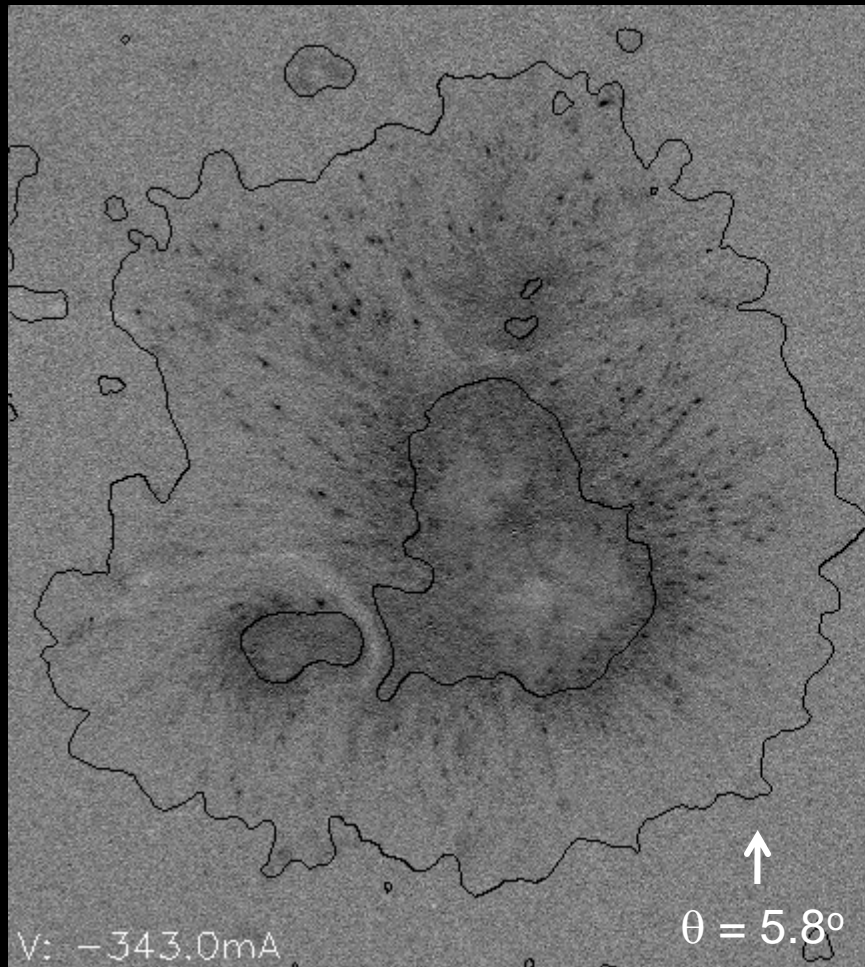
-277mA



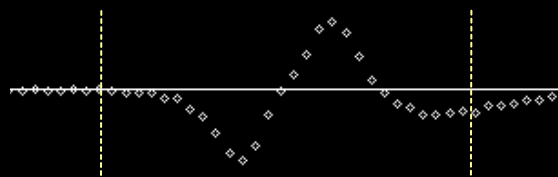
+277mA



# Stokes-V at 6302.5A $\pm$ 343mA



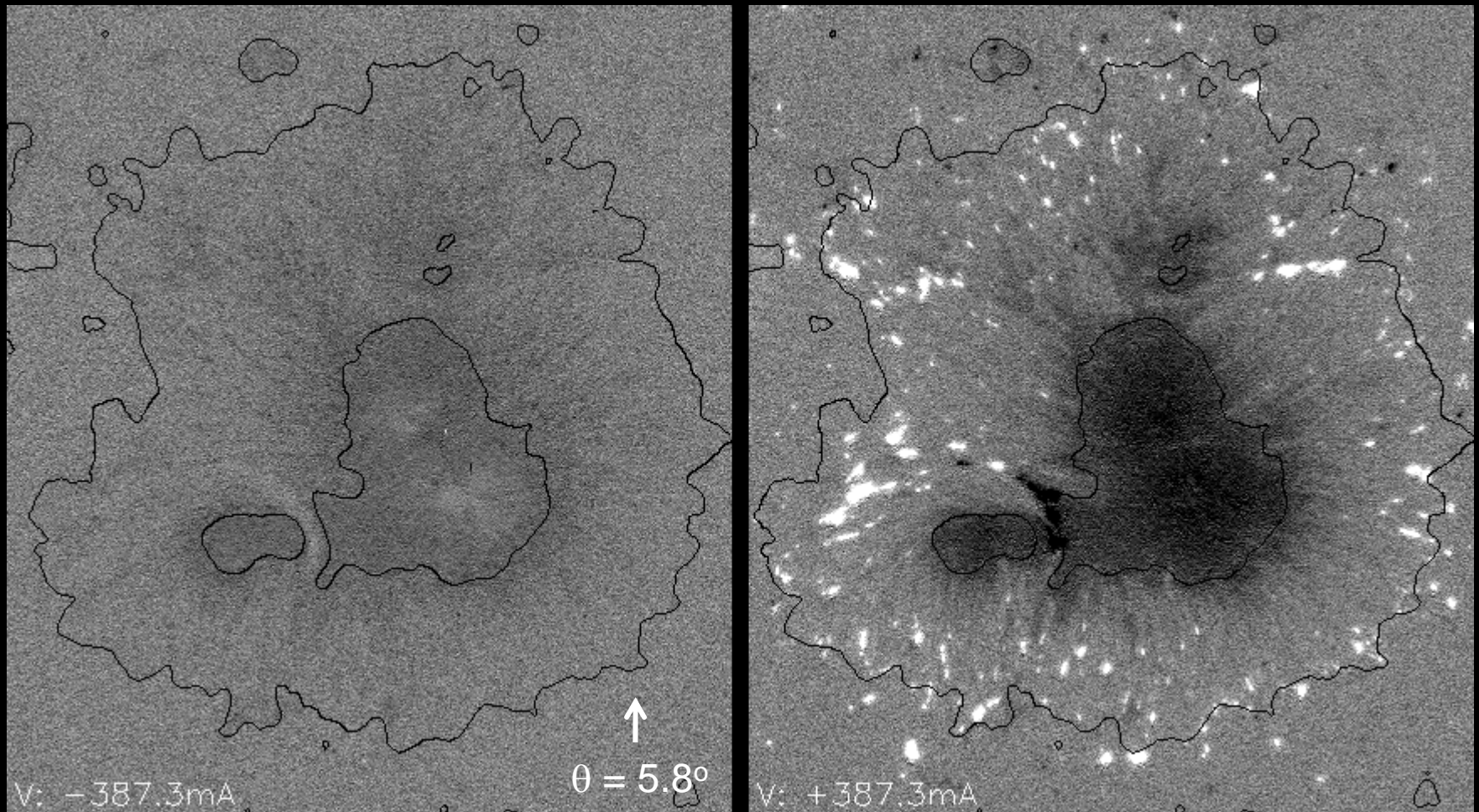
-343mA



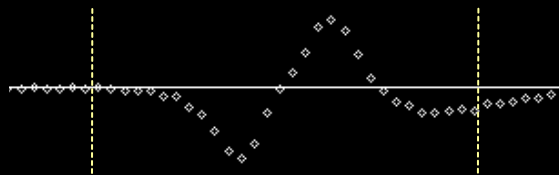
+343mA



# Stokes-V at 6302.5A $\pm$ 343mÅ



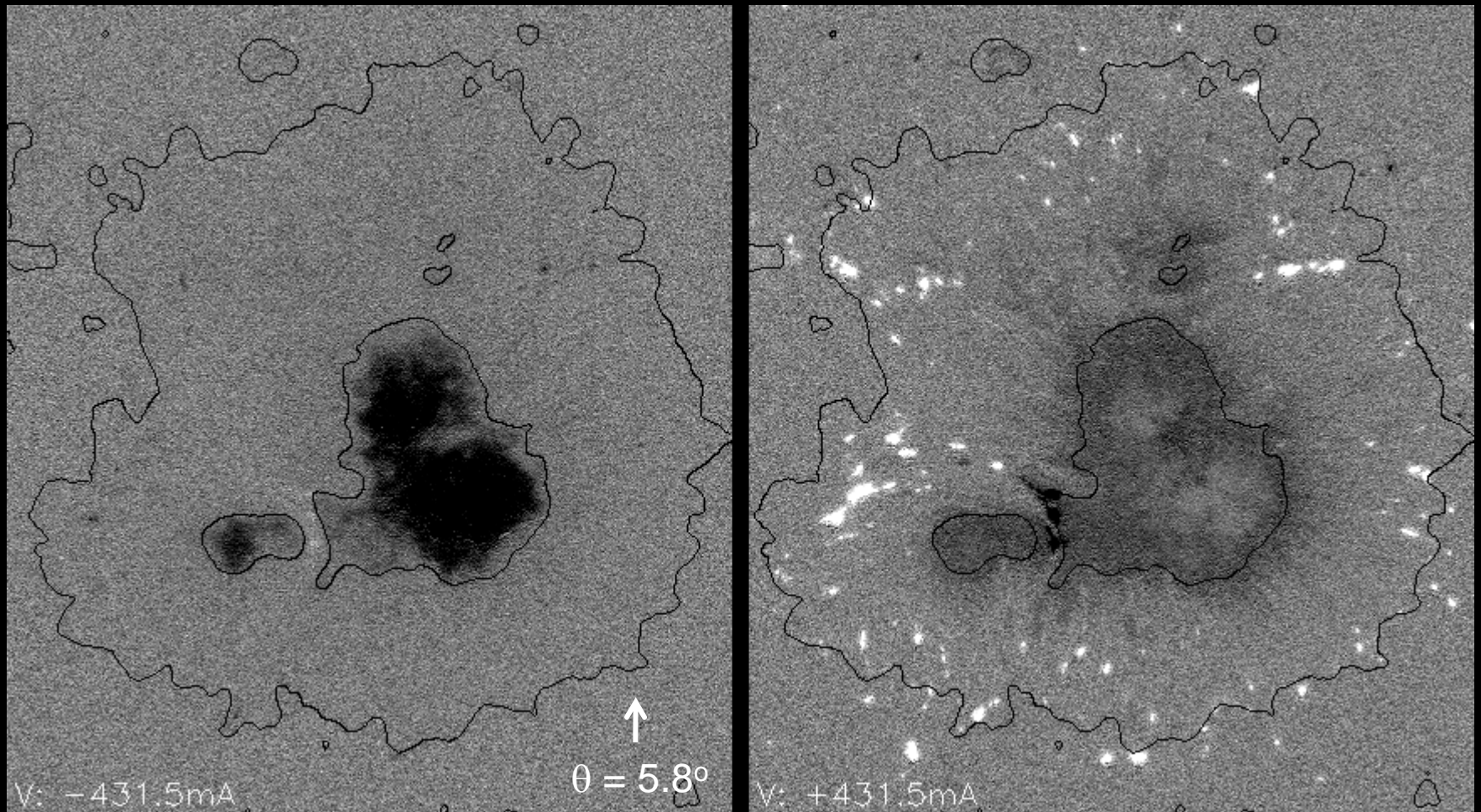
-387mÅ



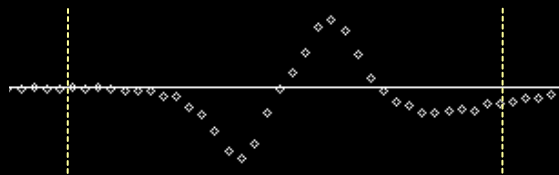
+387mÅ



# Stokes-V at 6302.5A $\pm$ 343mA



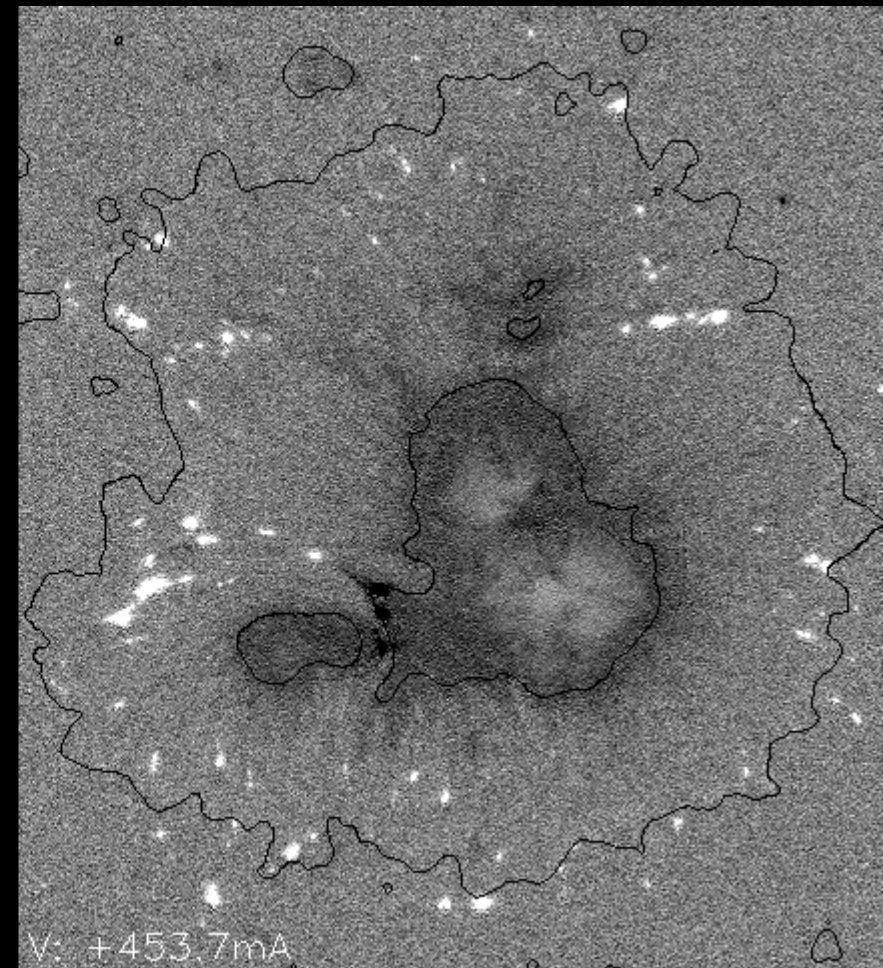
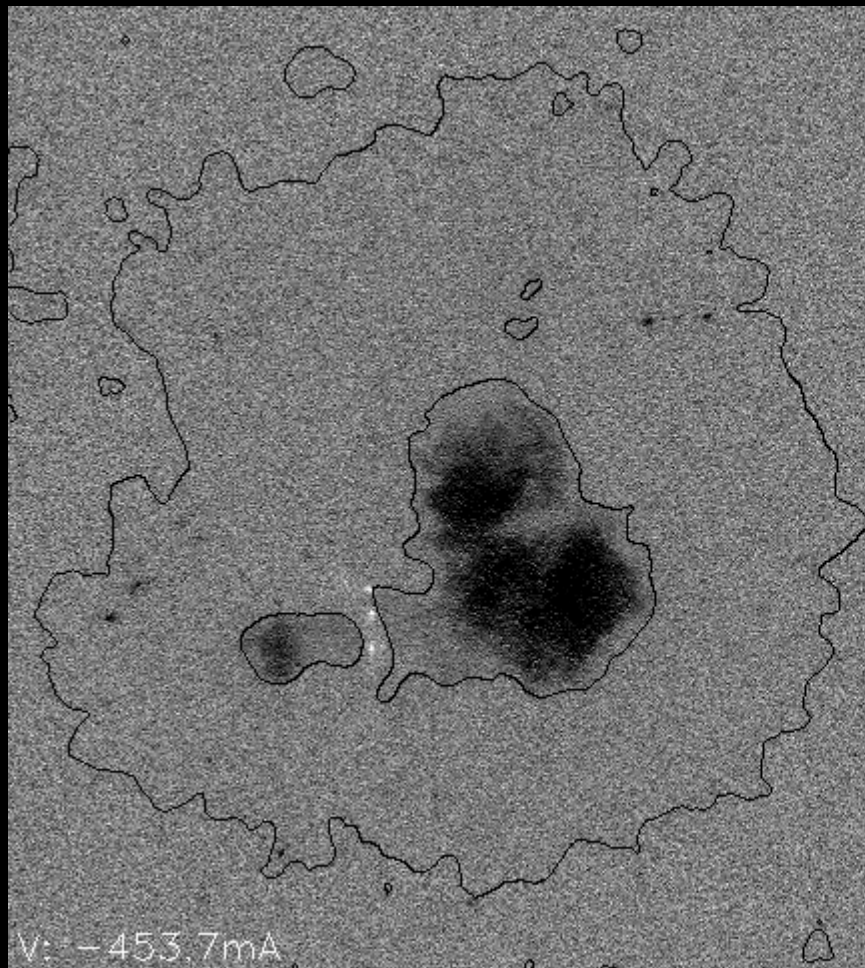
-431mA



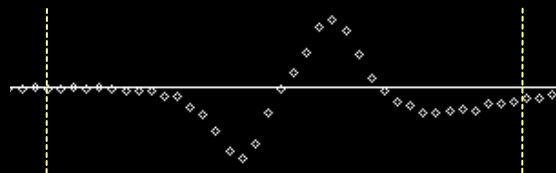
+431mA



# Stokes-V at 6302.5A $\pm$ 454mA



-454mA

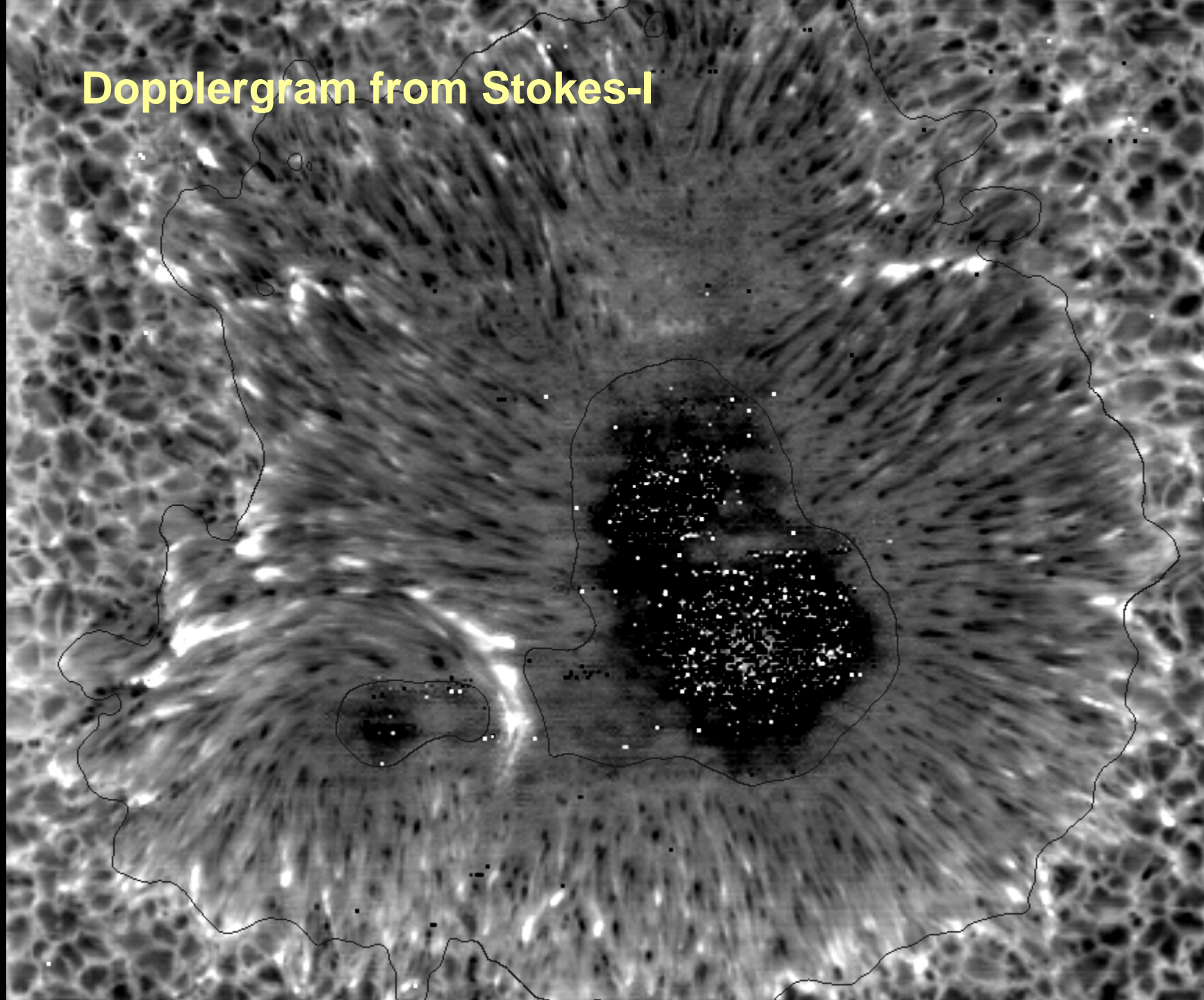


+454mA

→ 22km/s (supersonic?)

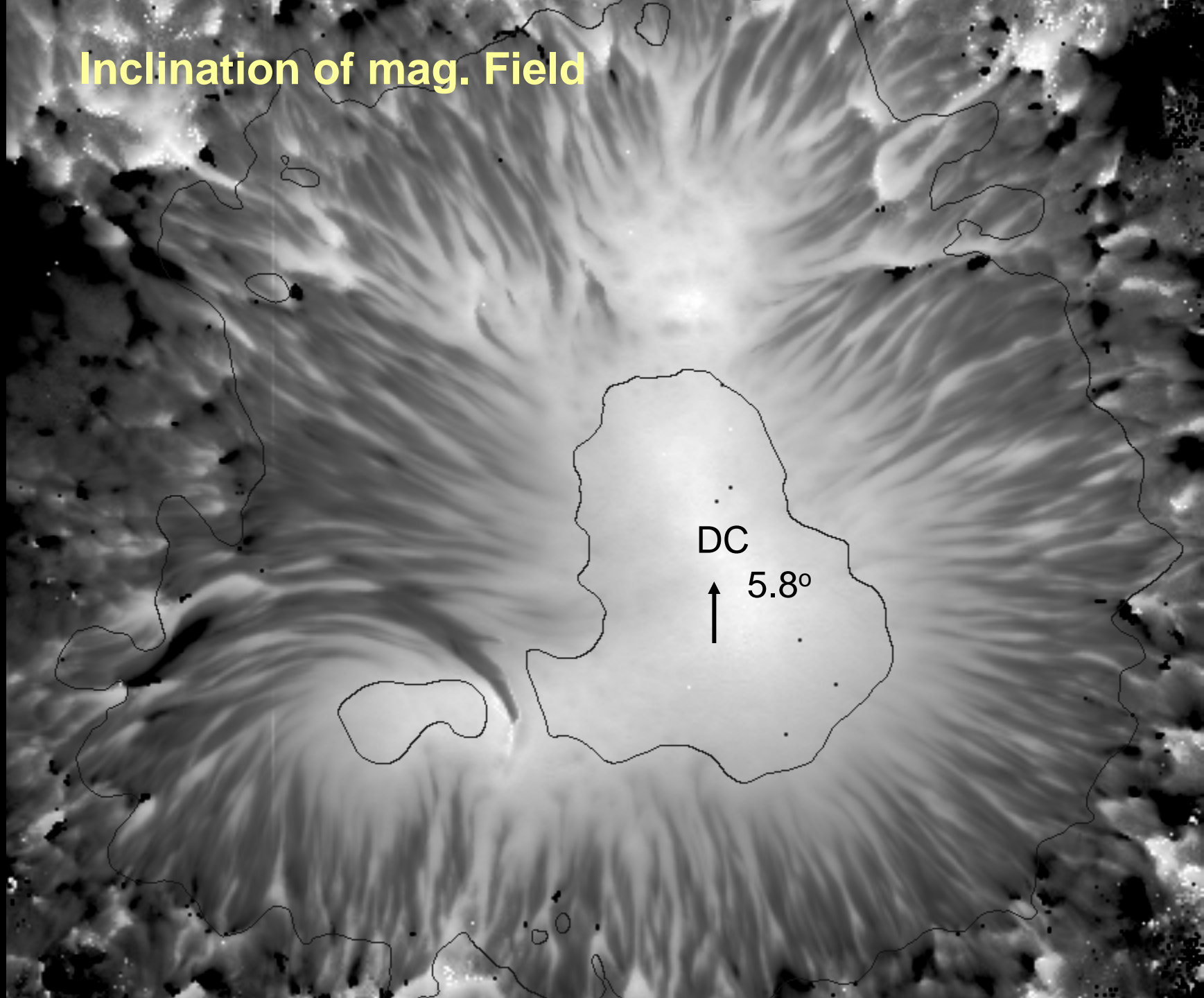


# Dopplergram from Stokes-I

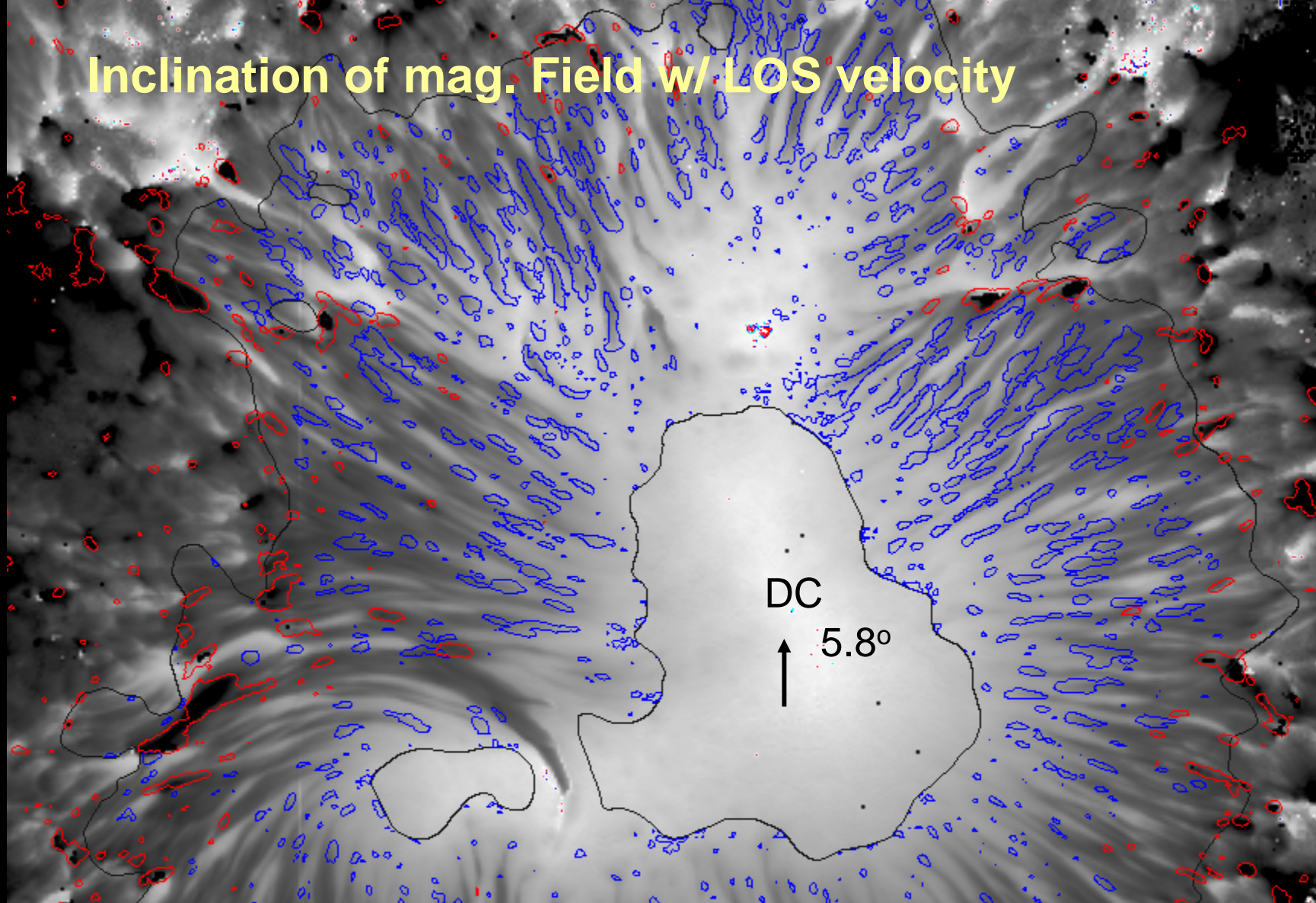




# Inclination of mag. Field



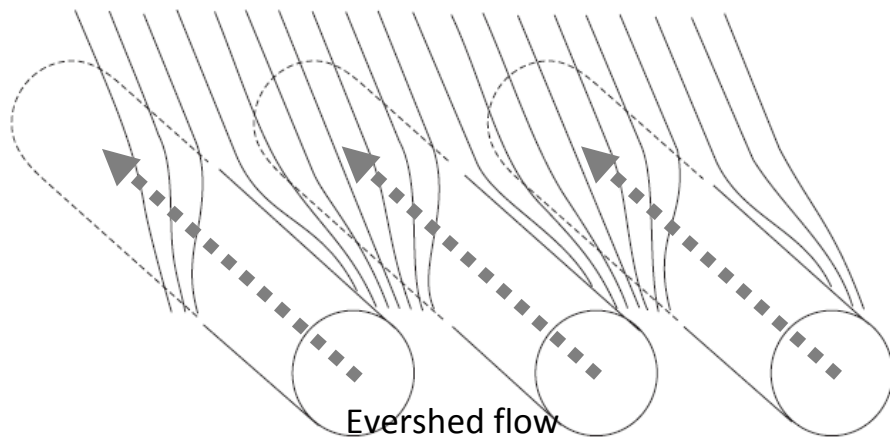
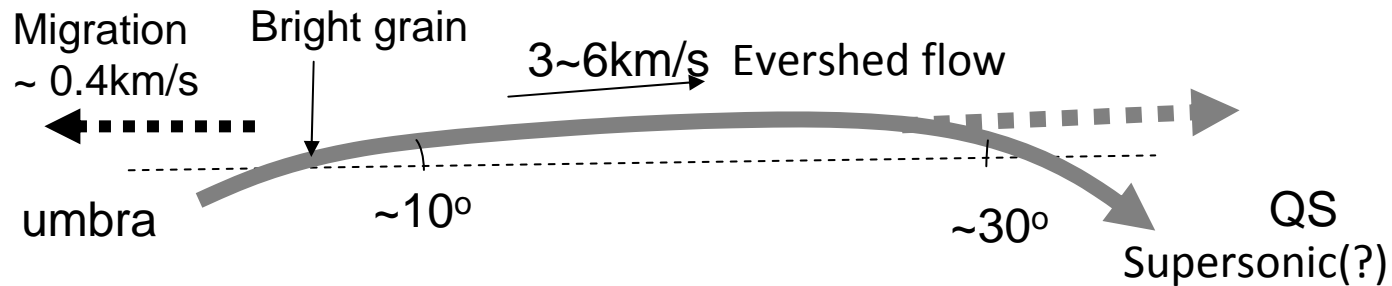
# Inclination of mag. Field w/ LOS velocity



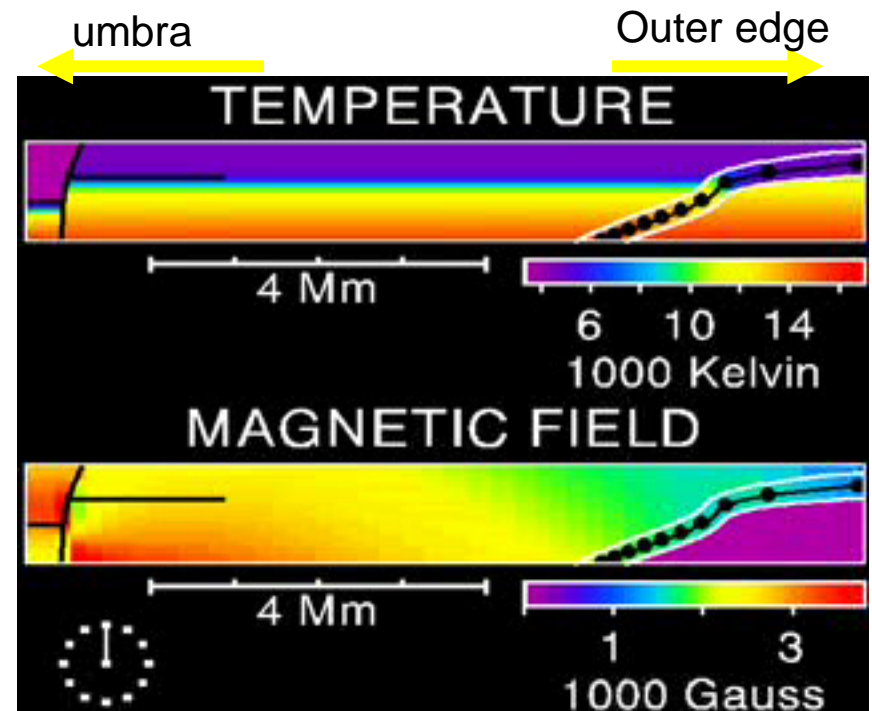
**Upflow and downflow patches are aligned on horizontal field filaments that carries the Evershed flow.**

**→ Source and sink of individual Evershed flow channel!**

Individual Evershed flow channels consistent with the rising flux tube model w/ uncombed structure.

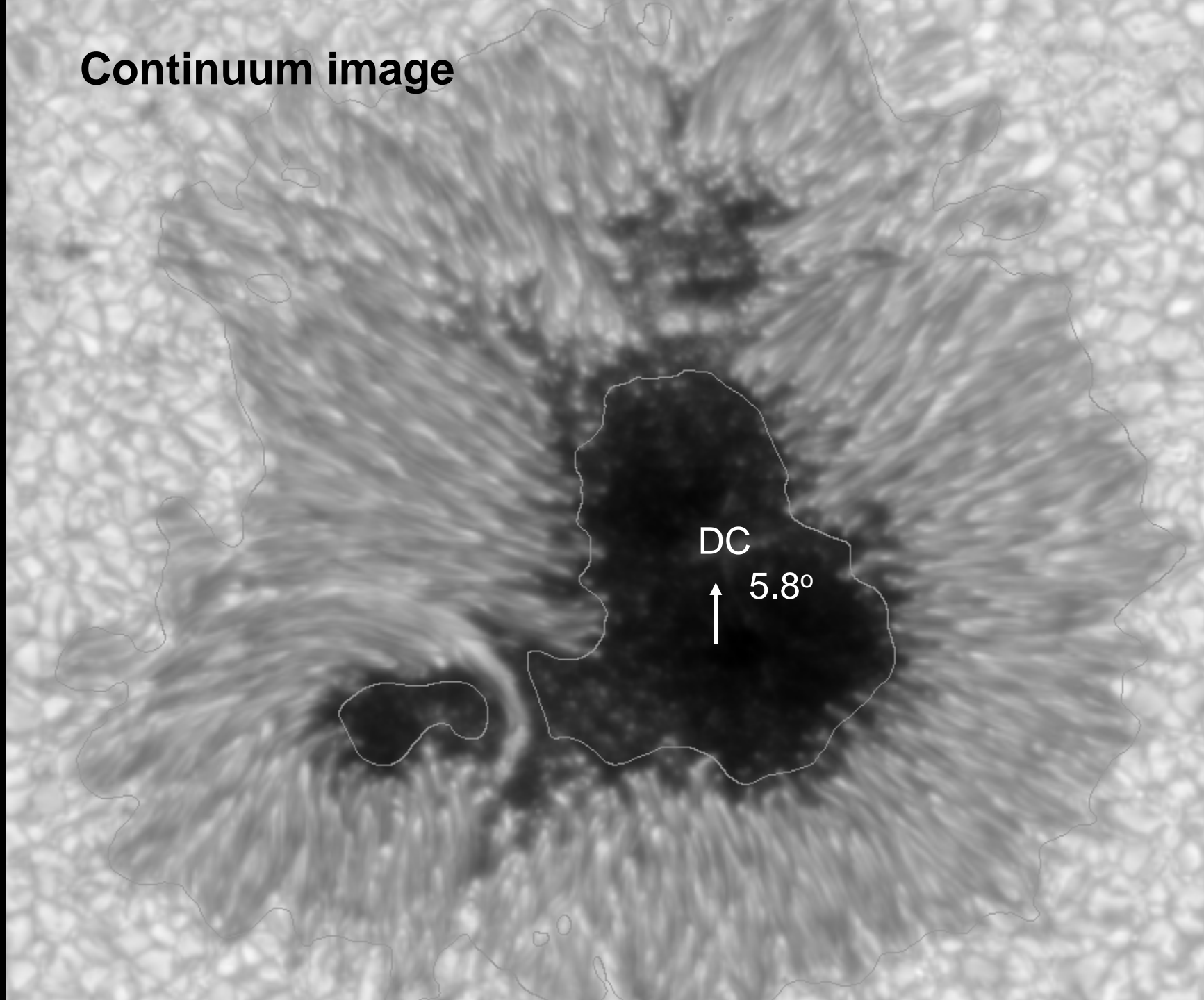


'uncombed' penumbral model  
Solanki & Motavon 1993



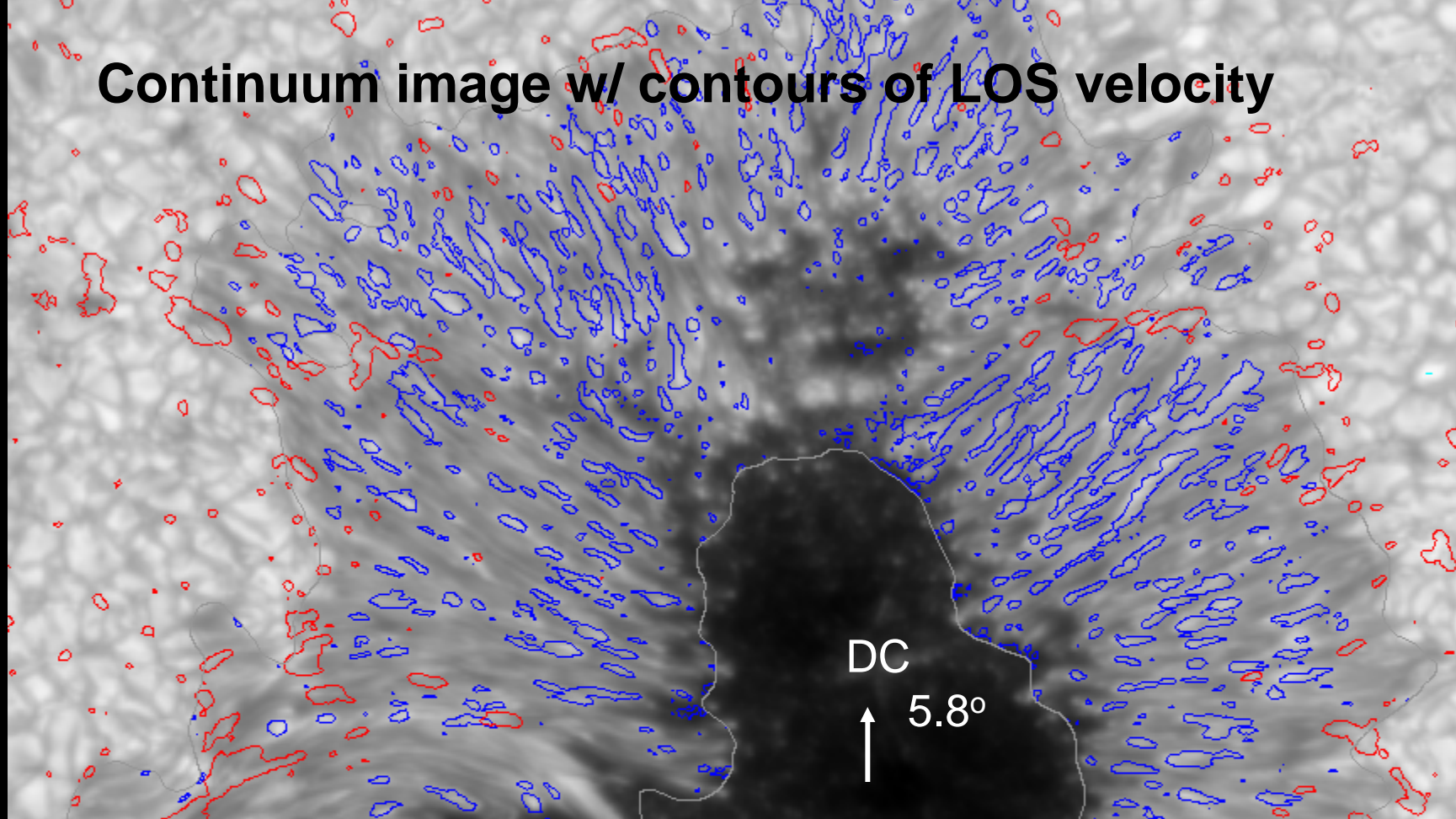
flux tube model (Schlichenmeier et al 1998)

# Continuum image





# Continuum image w/ contours of LOS velocity

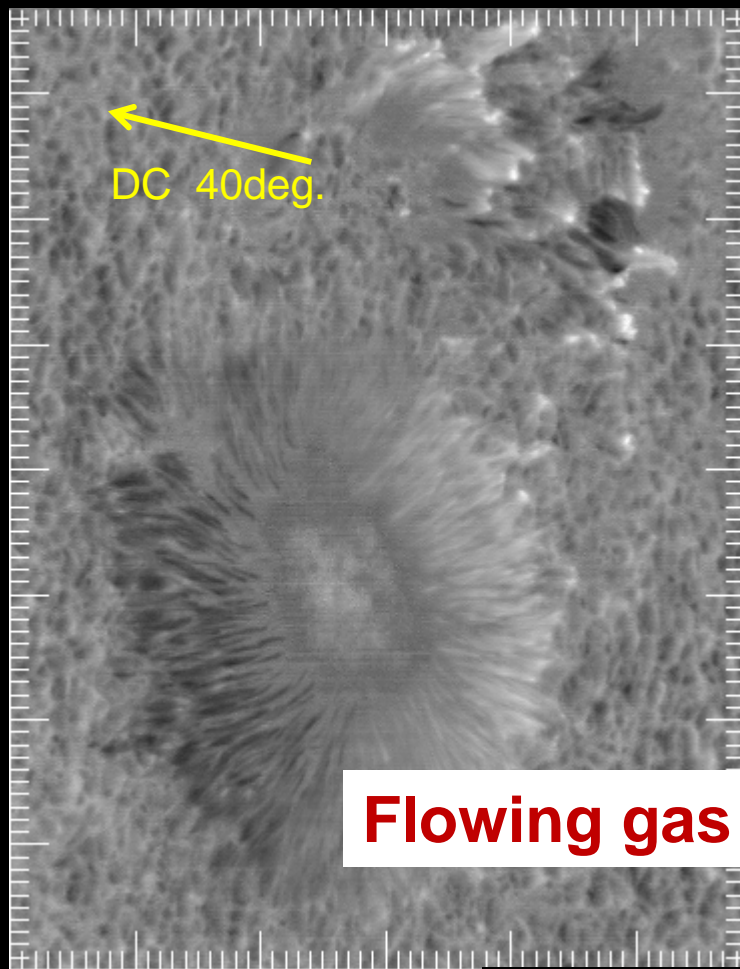


**Very good correlation between bright grains and upflows.  
→ Evershed flow carries the energy to maintain the penumbral brightness!**

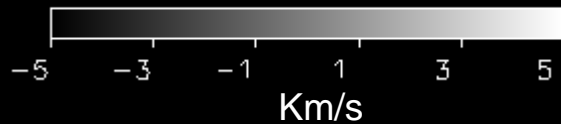
6302.5A Doppler shift, 2007.1.8

CG of Stokes-I

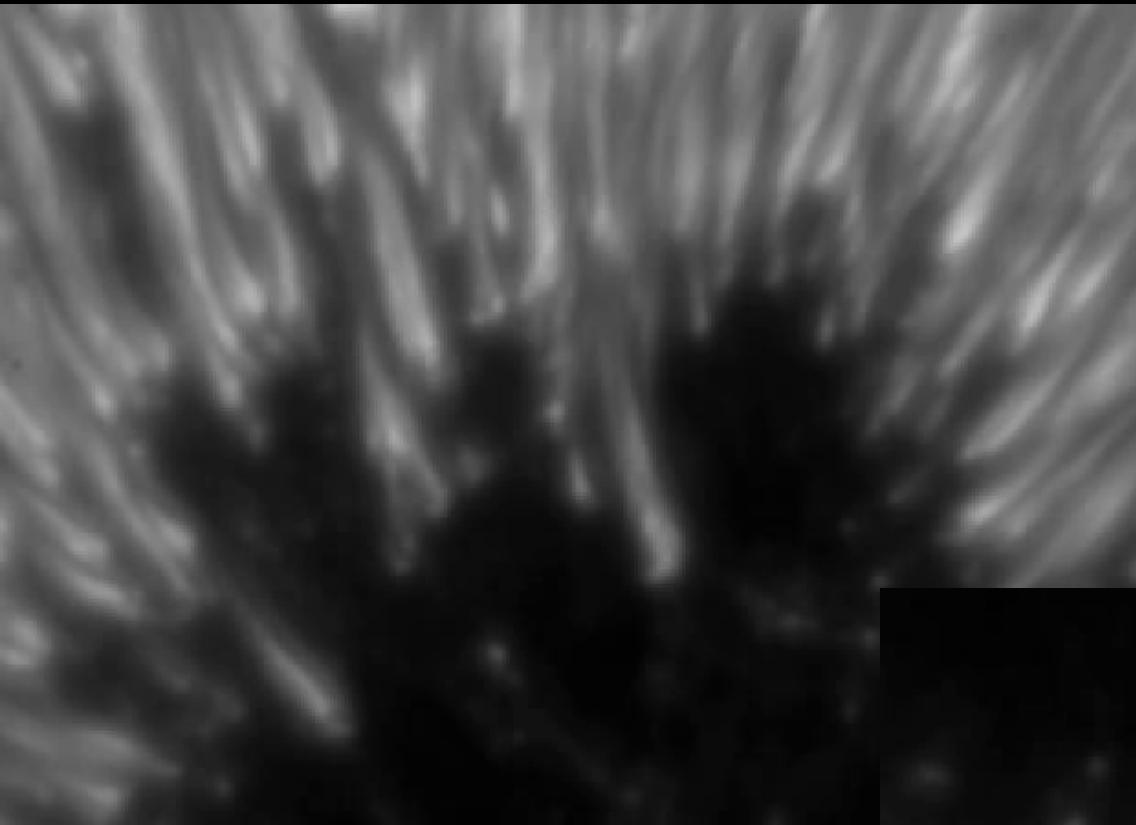
CG of  $\sqrt{V^2+Q^2+U^2}$



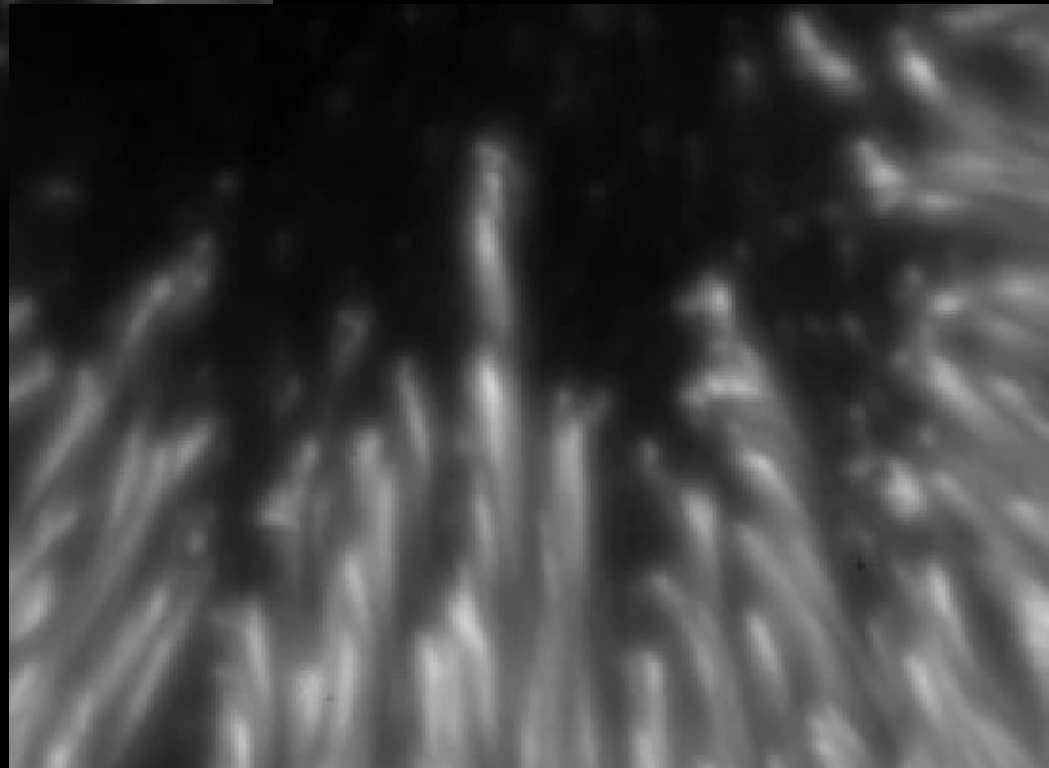
**Flowing gas is magnetized!**



2007.1.7

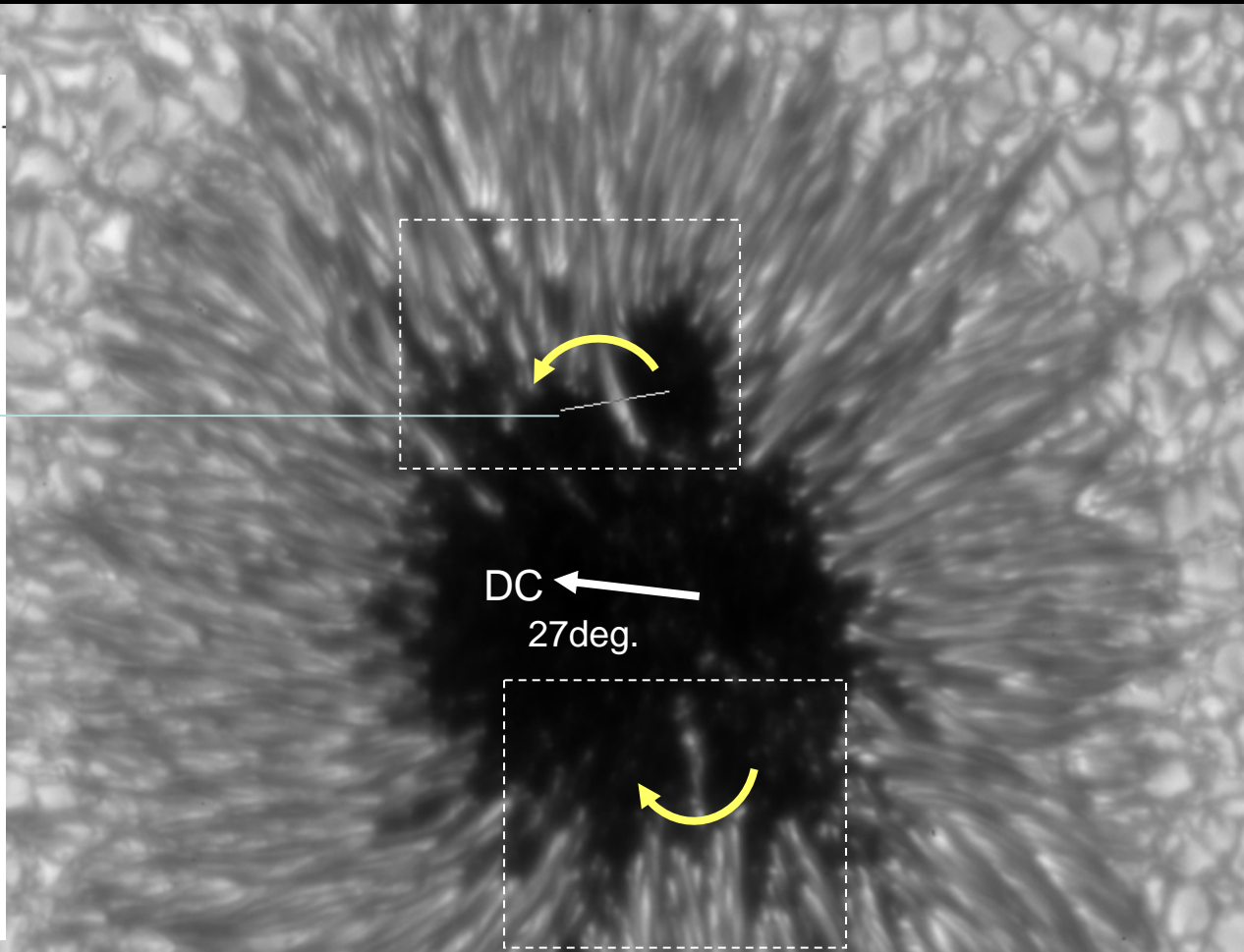
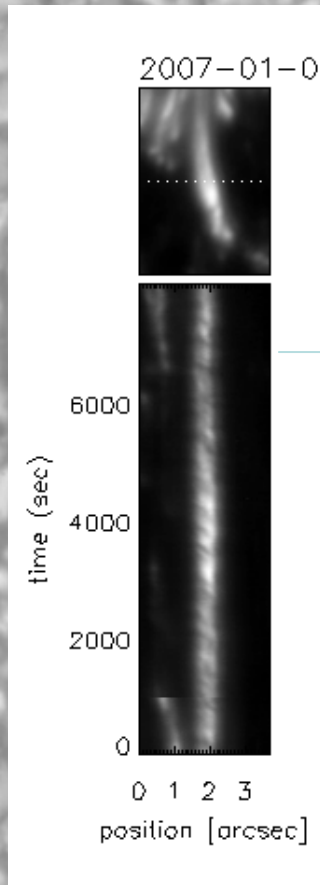


*Twisting filaments...*





2007.1.7

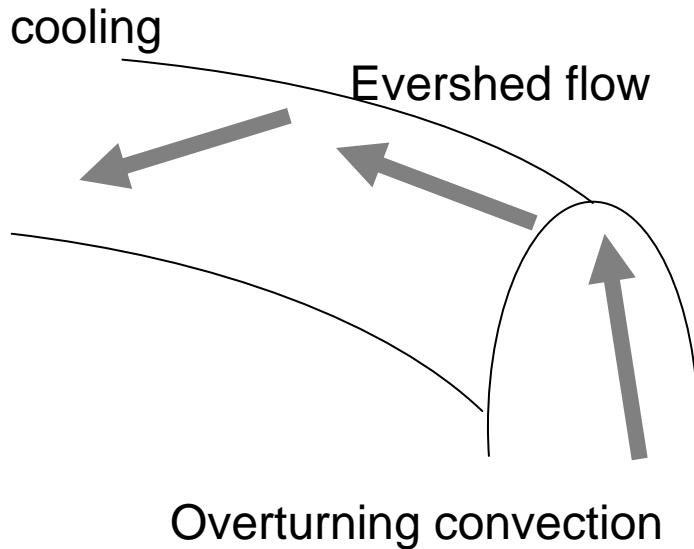


The 'twisting motion' of penumbral filaments is not a real turn of individual filaments, but is a manifestation of their dynamical nature such that the appearance depends on the viewing angle.

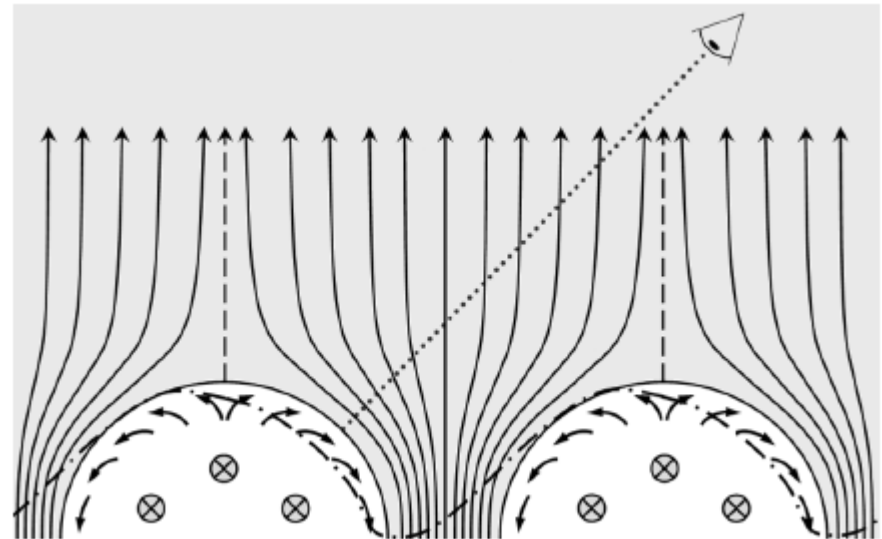


# What is the origin of the twisting appearance?

→ **Overturning-convection seen from a side(!?)**

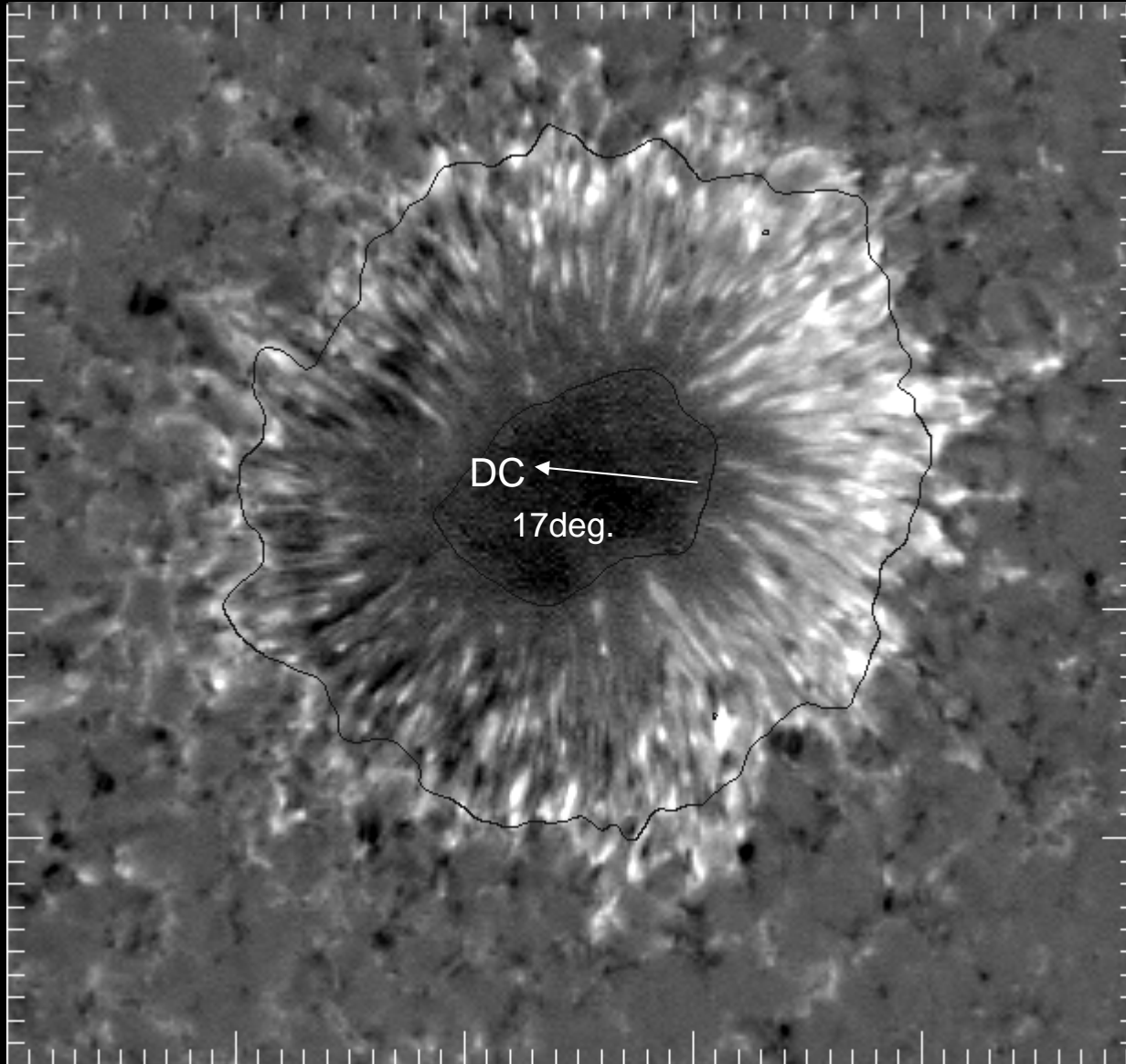


Ichimoto, et al., 2007, Science, 318, 1597

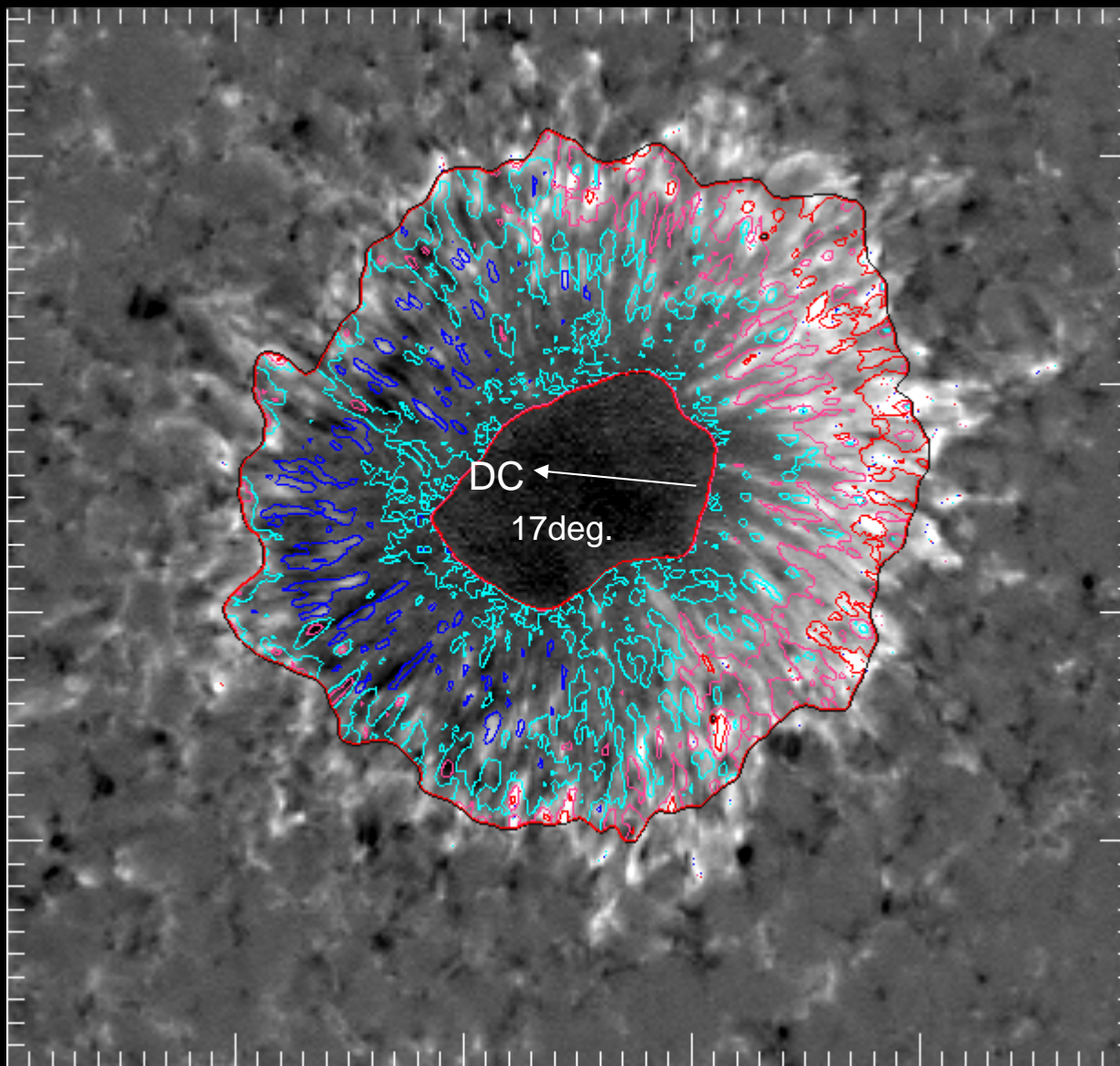


V. Zakharov, et al., 2008,  
A & A manuscript no. 0266 c ESO

# Net circular polarization in SOT resolution



# Net circular polarization in SOT resolution



**Evershed flow channels in both limb-side and DC-side penumbra produce a positive NCP!!**

# Summary (1):

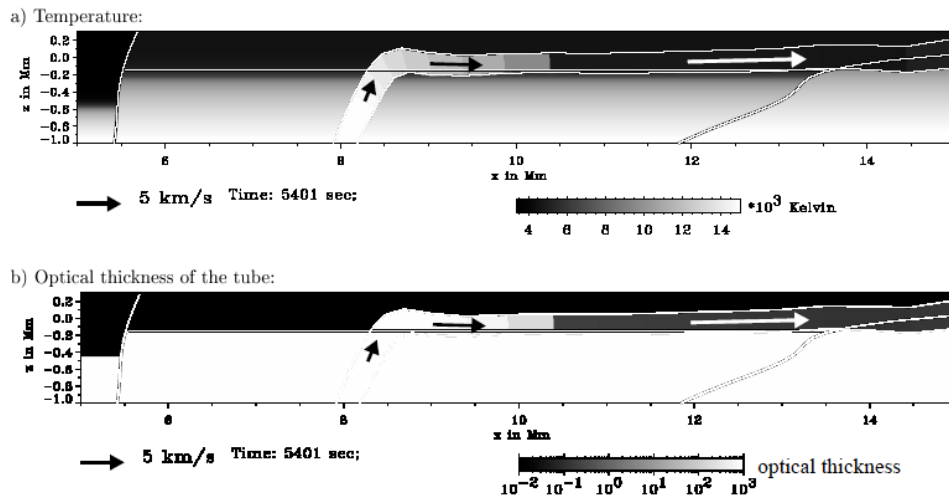
## **‘Convective nature of the Evershed Effect’**

- 1) Source and sink of the Evershed flow are identified; The geometry is consistent with the 3D uncombed penumbral model.
- 2) Evershed flow carries the energy of penumbra.
- 3) Source region of Evershed flow channels shows a hint of overturning convection.
- 4) Flowing plasma is not field free, but magnetized.
- 5) Flow velocity (and magnetic field strength) increase with depth in flowing channel (← NCP).

# Flux tube model vs. gap model

## Embedded flux tube model

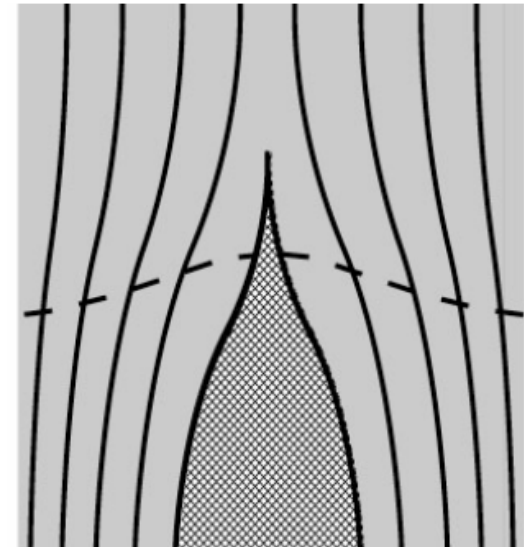
(e.g., Solanki & Motavon 1993  
Schlichenmaier et al 1998)



There is no observational evidence of the lower boundary of flux tubes.

## Gap model

(e.g., Spruit & Schermer 2006)



Field free gap  
Penetrating convection

Flowing gas is not field free.

*In both models, buoyancy drives the rising motion.*

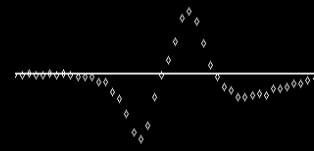
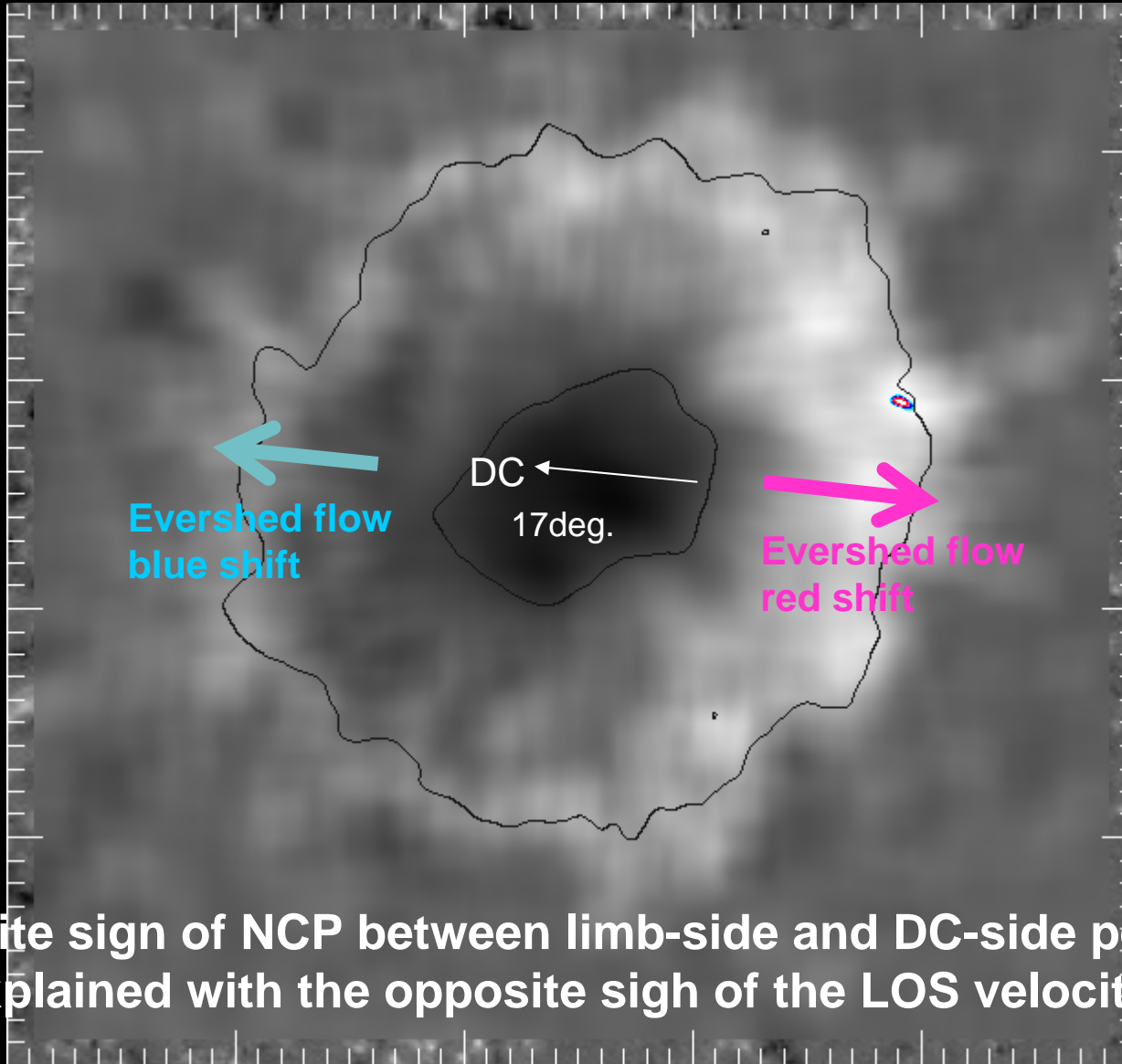
# Summary (2):

- If the flux tube model allows vertically elongated ‘flux tubes’, and if the gap model discard the word “field free”, then *there is no fundamental difference between the two models*. And SOT observations suggest this direction.
- Evershed effect may be interpreted as a natural consequence of ‘thermal convection’ under a strong, inclined magnetic fields.

Thank you!



# Net circular polarization in low resolution



$$\int v d\lambda$$

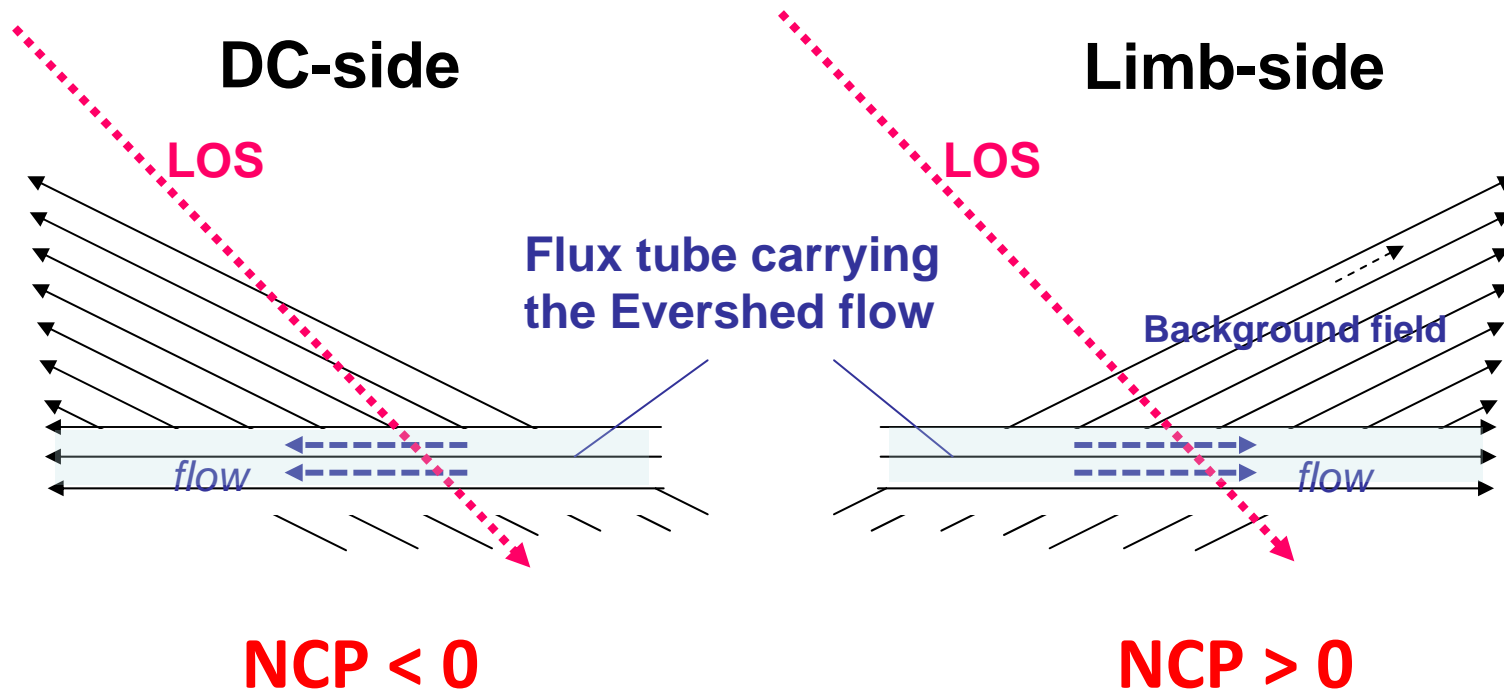
Depth veri.  
 $dv/d\tau$ ,  $dB/d\tau$

Opposite sign of NCP between limb-side and DC-side penumbra was explained with the opposite sign of the LOS velocity in deep layer.

Consistent with  
Horizontal Evershed flow in deep layer,

$$dy/dz$$

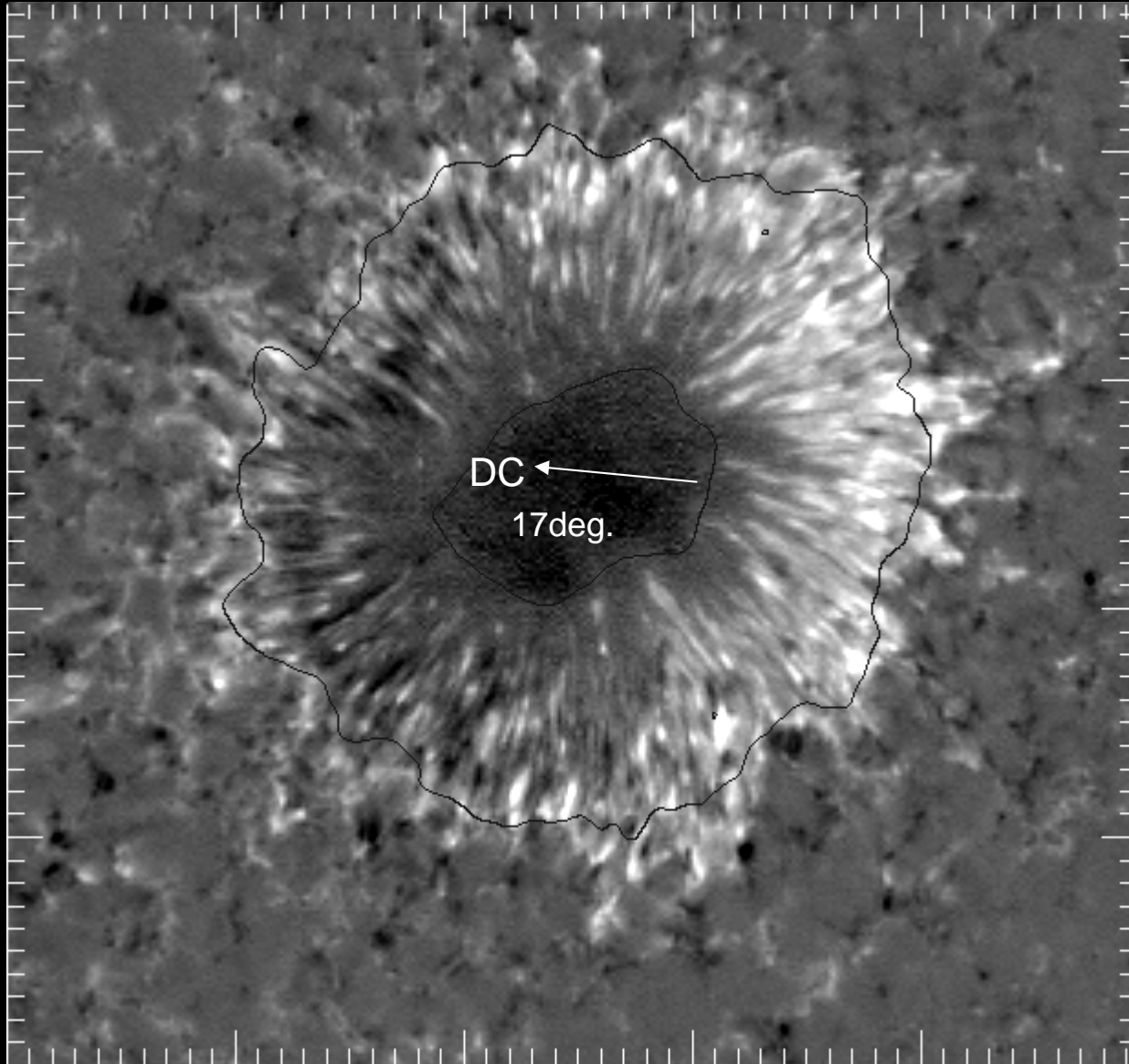
Solanki and Montavon (1992), Martinez Pillet (2000), Muller et al (2002,2006)



$$B_{\parallel} > 0, \quad \frac{d|\cos \gamma|}{d\tau} < 0, \quad \frac{dv_{LOS}}{d\tau} < 0$$

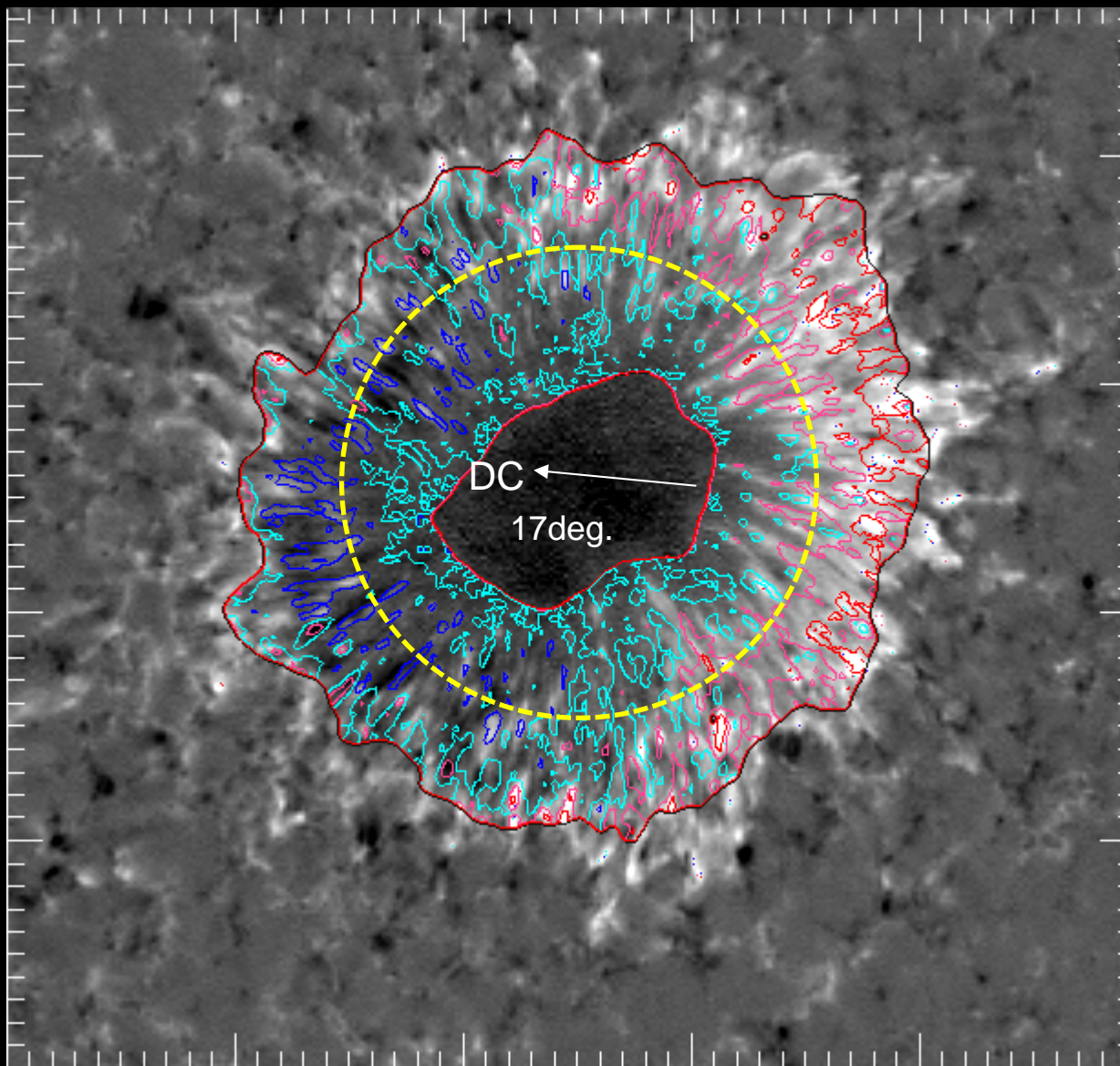
$$B_{\parallel} < 0, \quad \frac{d|\cos \gamma|}{d\tau} > 0, \quad \frac{dv_{LOS}}{d\tau} > 0$$

# Net circular polarization in SOT resolution



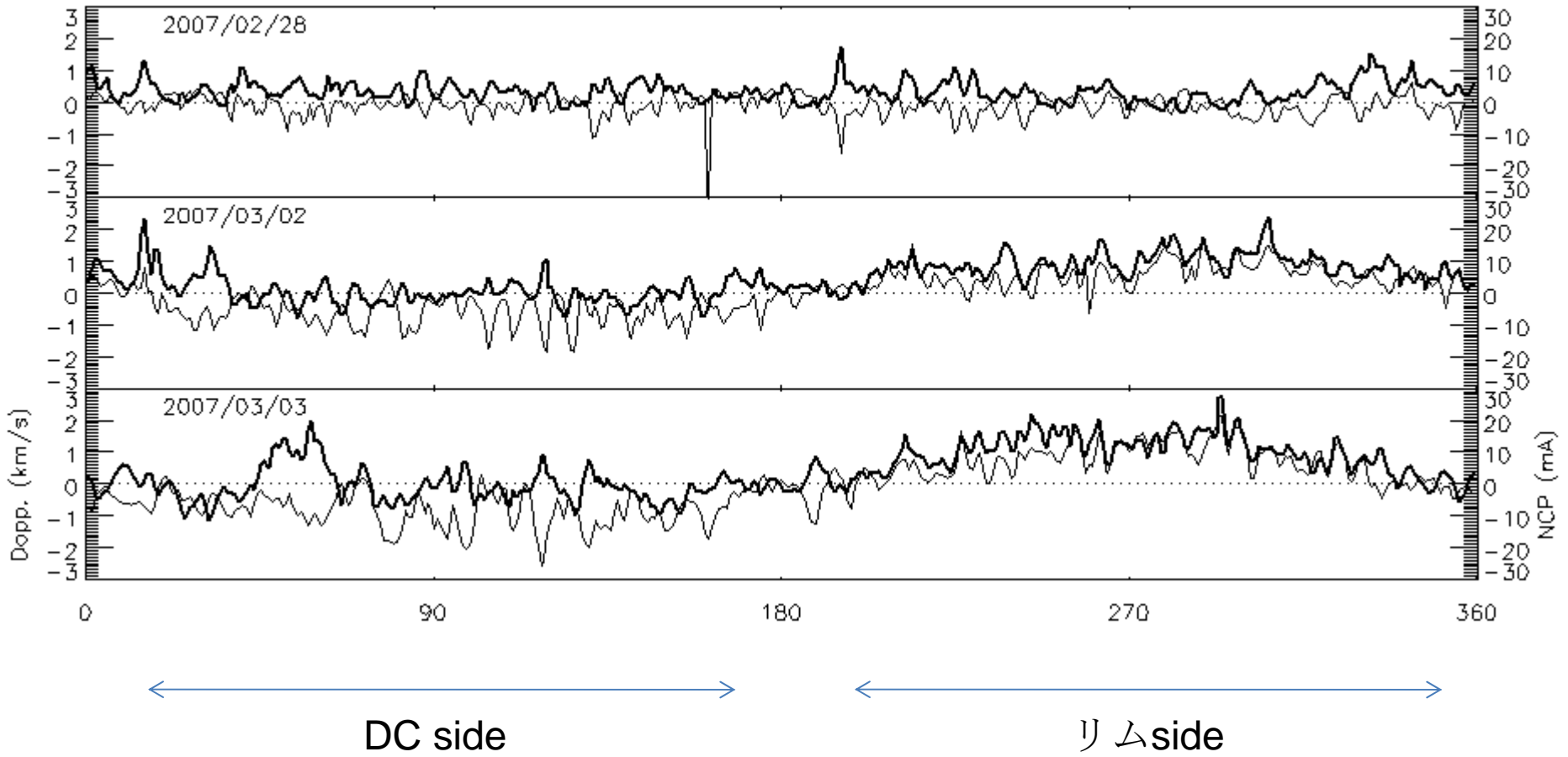


# Net circular polarization in SOT resolution



**Evershed flow channels in both limb-side and DC-side penumbra produce a positive NCP!!**

Thick: NCP  
Thin: Doppler



Thus penumbral NCP cannot be explained by  $d\gamma/d\tau$  effect, but require a positive correlation between  $|v_{LOS}|$  and  $|B|$  along the LOS.

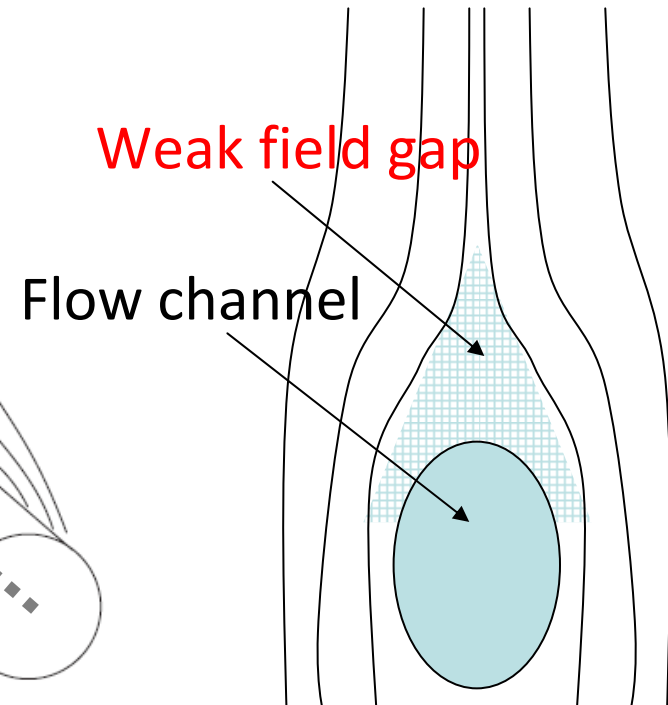
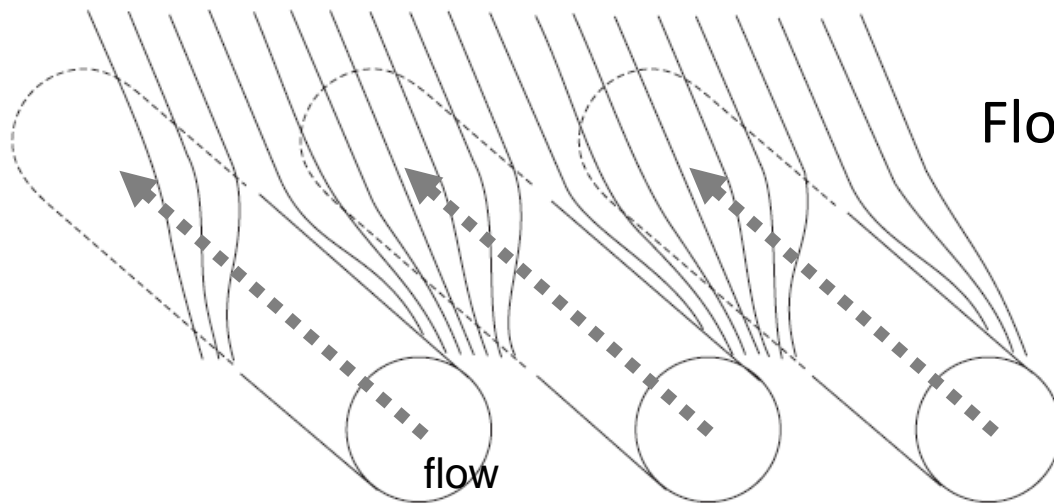
# Suggesting the presence of

## **weak field gap above the flow channel**

Ichimoto et al. 2008, A&A, 481, L9

Uncombed structure

Solanki & Motavon 1993



**Flowing gas is not field free, but magnetized!**

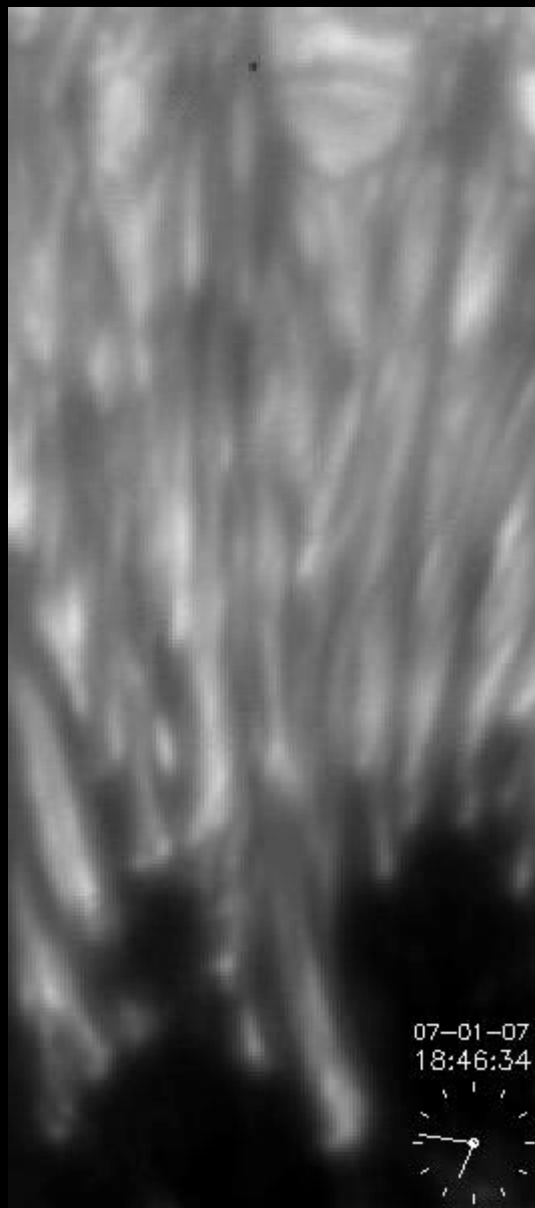
Positive correlation between  $v_{\text{LOS}}$  and  $B$  at

**Negative NCP in DC-side penumbra is produced in inter-Evershed flow channels! → presence of flows there too!**





07-01-07  
18:46:34

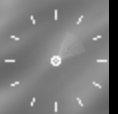


07-01-07  
18:46:34



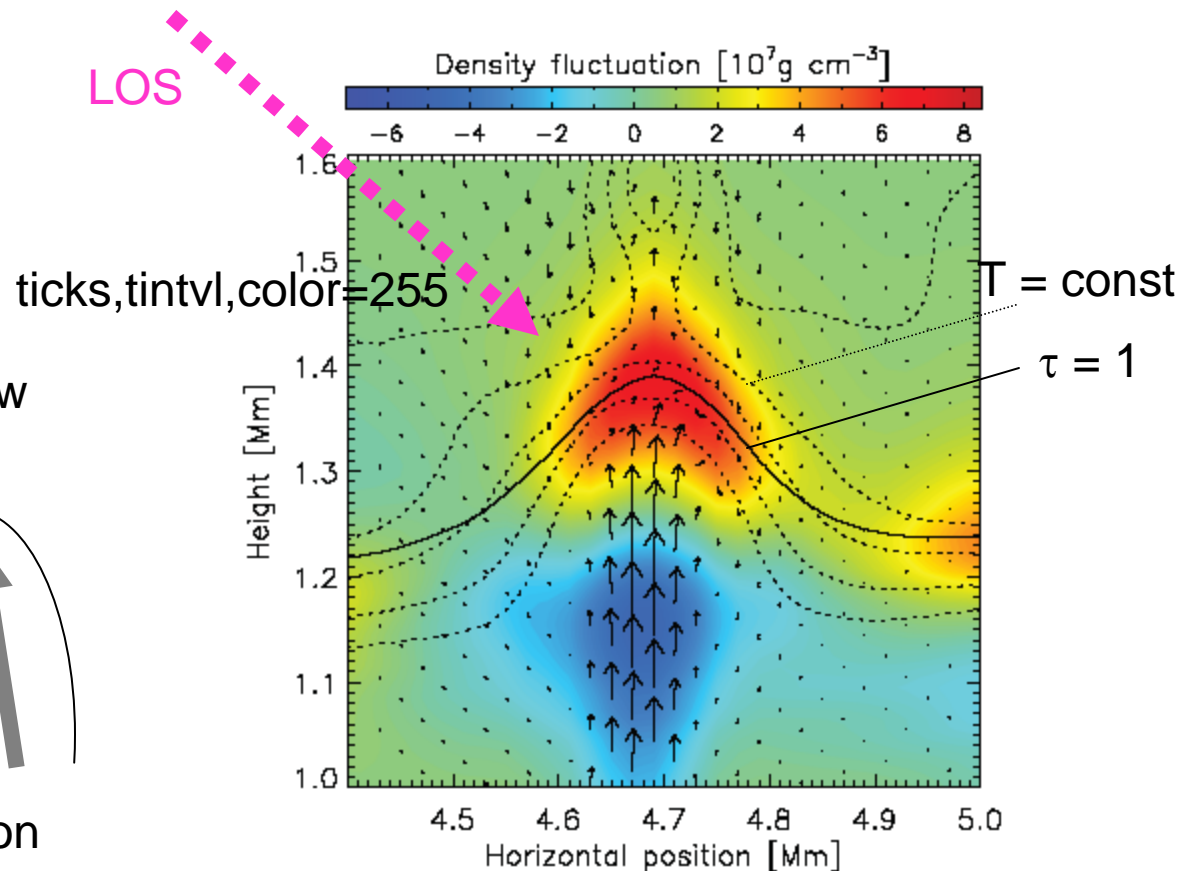
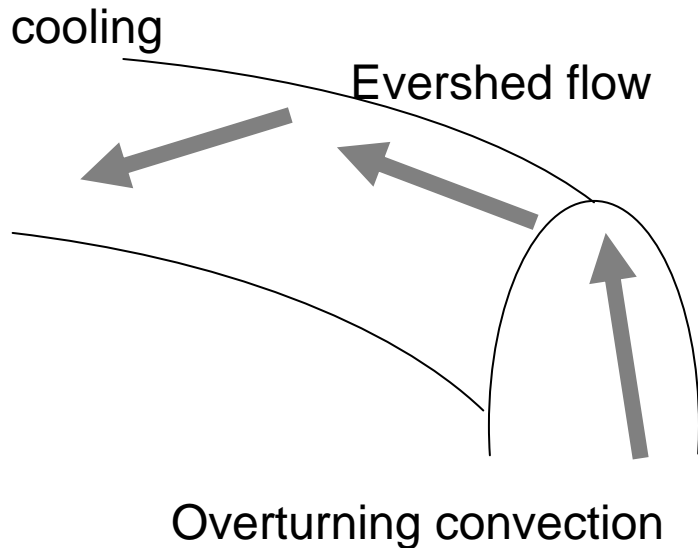
## 110min average

Local intensity fluctuations move in radial direction.  
Flow channels have life time longer than 1 hour.  
Evershed flow is not a stationary (or uniform) flow  
(eg. Shine et al. 1994, Solana et al. 2007, 2008)



# What is the origin of the twisting appearance?

→ **Overturning-convection seen from a side(!?)**



Umbral dot simulation

Schussler and Vogler, 2006, ApJ, 641, L73