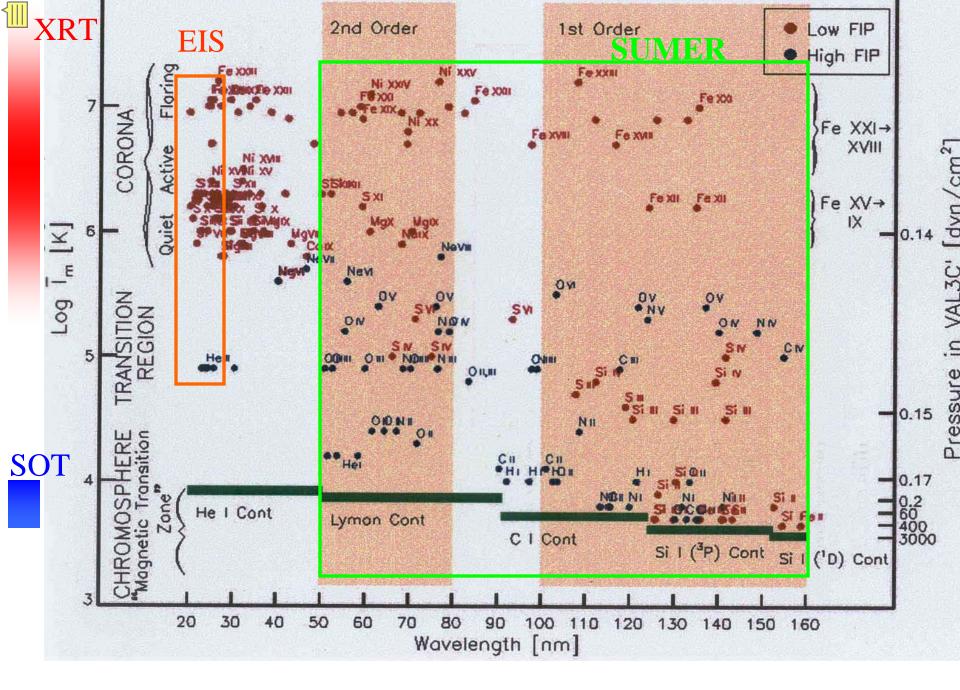
Results from the Hinode/SUMER campaigns

L. Teriaca, Max-Planck-Institut für Sonnensystemforschung



SUMER/Hinode past joint observations and campaigns



SUMER following Hinode targets during commissioning

• October and November 2006

Spring 2007 campaign

- Campaign from April 2 to 29
- Run 22 observing proposals + TOO observations

Fall 2007 campaign

- Campaign from November 2 to 16
- Run 14 observing proposals

<u>April 2008</u>

• Observations in the WHI program from 10 to 16 April

<u>June 2008</u>

Joint observations





Teriaca:	Spatial and temporal evolution of the temperature response during VUV explosive events				
Gomory:	Energy transport and dynamics in/above the network and coronal heating mechanisms				
Landi:	The thermal structure of off-disk quiet Sun and active region plasmas				
Teriaca: The average Doppler shift of coronal lines on quiet and active regions					
Madjarska: Small-scale transient flows in the quiet Sun and active regions					
Madjarska: Coronal Bright Point plasma characteristics and evolution					
Solanki:	Centre-to-limb variation of active region and quiet Sun brightness				
Doyle:	Further insight into the spicules/blinker connection: a search for blinkers using EIS				
Doyle:	Oscillations in chromospheric and coronal bright points				
Doyle:	Armagh explosive event study				
Wilhelm:	Im: N _c and T _c diagnostic in polar plumes (W. Curdt et al., 2008, A&A 481, L61;				
	L. Feng et al. 2008, in preparation)				
Fontenla:	Quiet-Sun radiance distribution and UV variability	black	Nothing done/No infos		
Marsh:	MHD wave propagation		6		
Teriaca:	Detection of waves in the solar atmosphere	Blue: Red:	Work in progress paper in preparation		
Bewsher:	Dynamic events in the network				
Doschek:	Transition region T_{e} diagnostics	Red:	Paper accepted/published		
Curdt:	Super disk atlas: (Tian et al. 2008, ApJ 681, L121) - Poster 1-11				
Doschek:					
Innes:	Chromospheric heating in quiet Sun - Talk 1-9				
Landi:	Diagnostics of quiescent active region loops				
Schmieder: Prominence / JOP 178 (Heinzel et al., 2008, ApJ, In Press)					
Kamio: Velocity field in a coronal hole (paper in preparation) – Talk 7-6					
тоо	Observations of AD 10052 turnends the end of the company	m (DE L	nnog 2008 A 8-A 481 T 41)		

TOO Observations of AR 10953 tuwards the end of the campaign (D.E. Innes, 2008, A&A 481, L41)

Fall 2007 campaign



G. Del Zanna, K. Reardon,	Multi-wavelength observations of coronal hole plumes at solar minimum Talk 5-2 Spectral Observations of Spicule Dynamics			
D. Innes,	Temperature, density and 3-D structure of active region loops			
D. Innes,	Doppler shifts in X-ray jets - Talk 1-9			
D. Innes,	3-D structure and evolution of filaments/prominences			
M. Madjarska,	Coronal holes boundaries evolution			
S. Imada,	Waves in front of/back side of (north/south) polar jets			
M.P. Miralles,	Characterization of Fast and Slow Solar Wind Source Regions			
S. Kamio,	SUMER campaign - coronal hole			
J.G. Doyle,	Magnetic structure of macrospicules	black	Nothing done/No infos	
K. Matsuzaki,	DEM analysis in lower corona	Blue:	Work in progress	
L. Teriaca,	Detection of waves in the solar atmosphere	Red:	paper in preparation	
D. Rabin,	Coordinated observations with EUNIS	Red:	Paper accepted/published	
S. Patsourakos,	Moss Observations			

April 2008 observations during the WHI

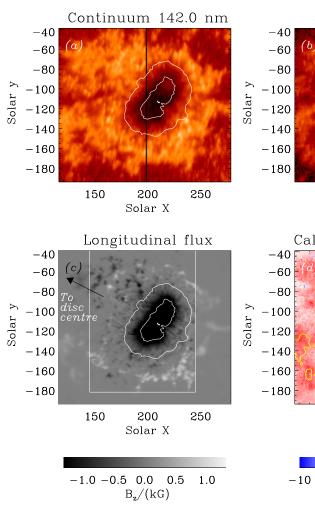
S. McIntosh, Characterizing the energetics and dynamics of the quit Sun -Talk 1.2, Poster 1.5

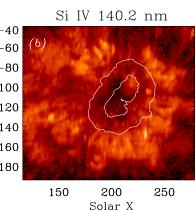
June 2008 observations

A Pietarila, G. Poletto Chromospheric network structure and dynamics Plume study

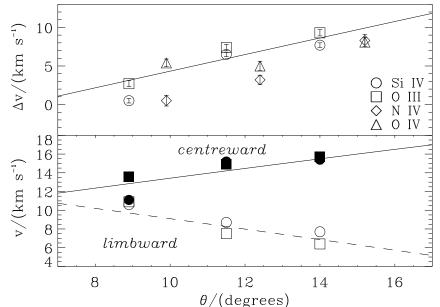
SUMER observations of a sunspot in November 2006

Teriaca et al. 2008, A&A Letter, accepted





Cal. v - Si IV 140.2 nm -40 -60 -80 -100 -120 -140 -160 -180 -150 200 250 Solar X -10 0 10 20 30 Velocity/(km s⁻¹)

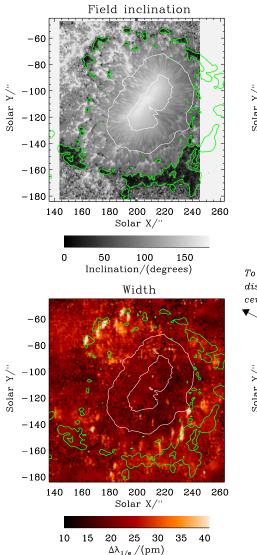


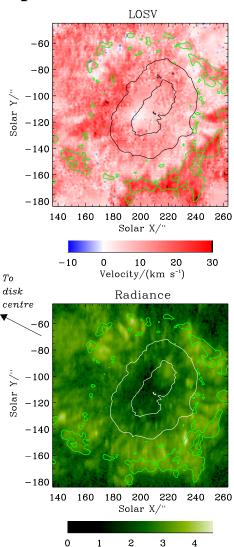
- •We observe a downflow pattern compatible with the presence of an Inverse Eveshed flow.
- •The flow is visible in lines from 0.08 to 0.18 MK.

•It occours in a collar of radially directed filamentary structures with widths < 1 Mm and inclined between 10° and 25° relative to the solar surface.

SUMER observations of a sunspot in November 2006

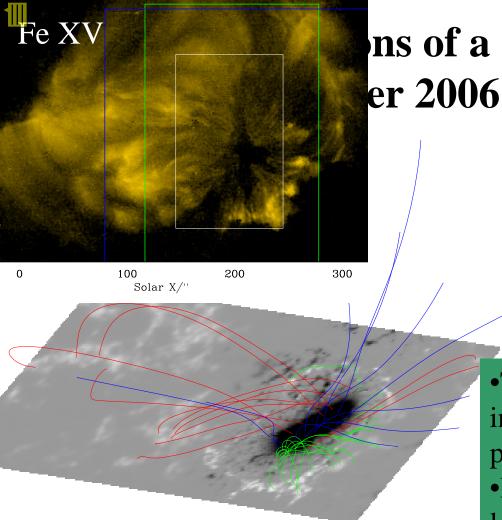


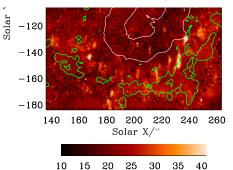




 $Log L/[mW m^{-2} sr^{-1}]$







 $\Delta 15 20 25 30 35 4$ $\Delta \lambda_{1/e} / (pm)$

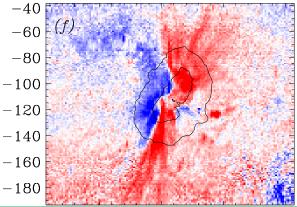


140 160 180 200 220 240 260 Solar X/"

> 1 2 3 4 Log L/[mW m⁻² sr⁻¹]

20 –10 0 10 20 Velocity/(km s⁻¹)

v – Ne VIII 77.0 nm



nua

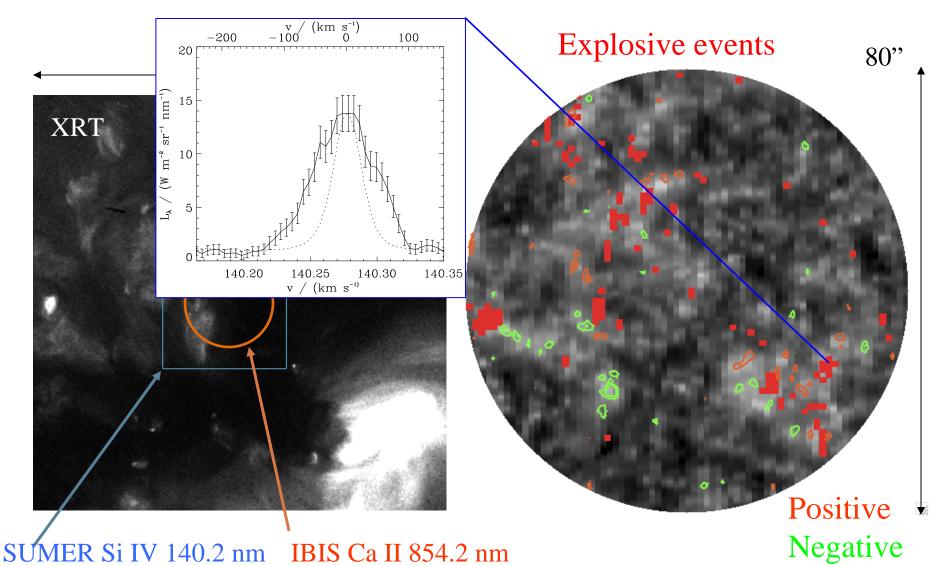
•The inverse Evershed flow is visible in a region roughly twice the penumbra.

•It seems occurring in far-reaching loops that only above the umbra and (extended) penumbra are cold enough to be seen in Si IV.

•Further out, emission is dominated by strong downflows at the footpoints of hot (> 2 MK) loops.



Spatial and temporal response of the solar atmosphere during VUV explosive events Cauzzi, Reardon, Teriaca, Pitterle, Curdt



Chromospheric velocity: 3 min power map



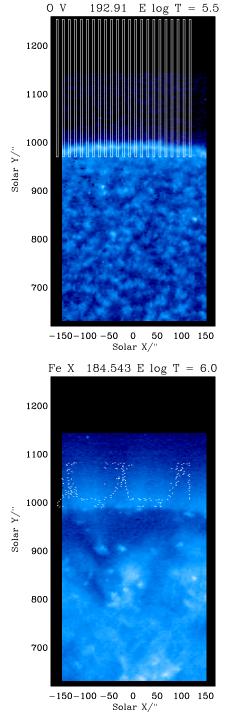
•Explosive events are located at the edges of the magnetic regions, at locations where the chromospheric fibrils stop being visible. •We are searching for chromospheric signatures either "vertically" either following the magnetic connectivity as shown by the fibrils. •XRT and EIS simultaneous observations are being investigated to study these events in the context of the whole solar atmosphere.

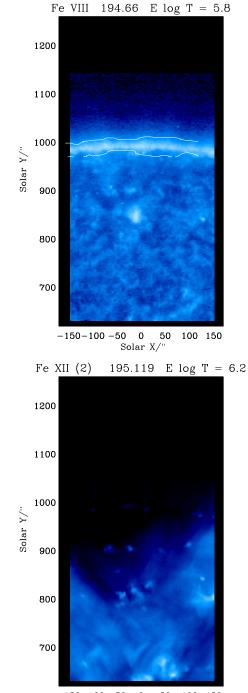
80"

Plume campaign Nov 2007

Del Zanna, Teriaca, Wilhelm, Andretta

34 h SUMER raster scan acquiring at each slit position several 4 nm wide spectra providing diagnostics of density, temperature and abundance to be combined with / compared to EIS data.





-150-100 -50 0 50 100 150 Solar X/"



The present

A SUMER/Hinode campaign devoted to the study of the polar coronal holes is currently under way with about 8 observing programs being run

The future

With SUMER still performing very well, we look forward for more opportunities of join forces with Hinode and other space and ground facilities to serve the scientific community.