

# Active region loops: Hinode/EIS observations

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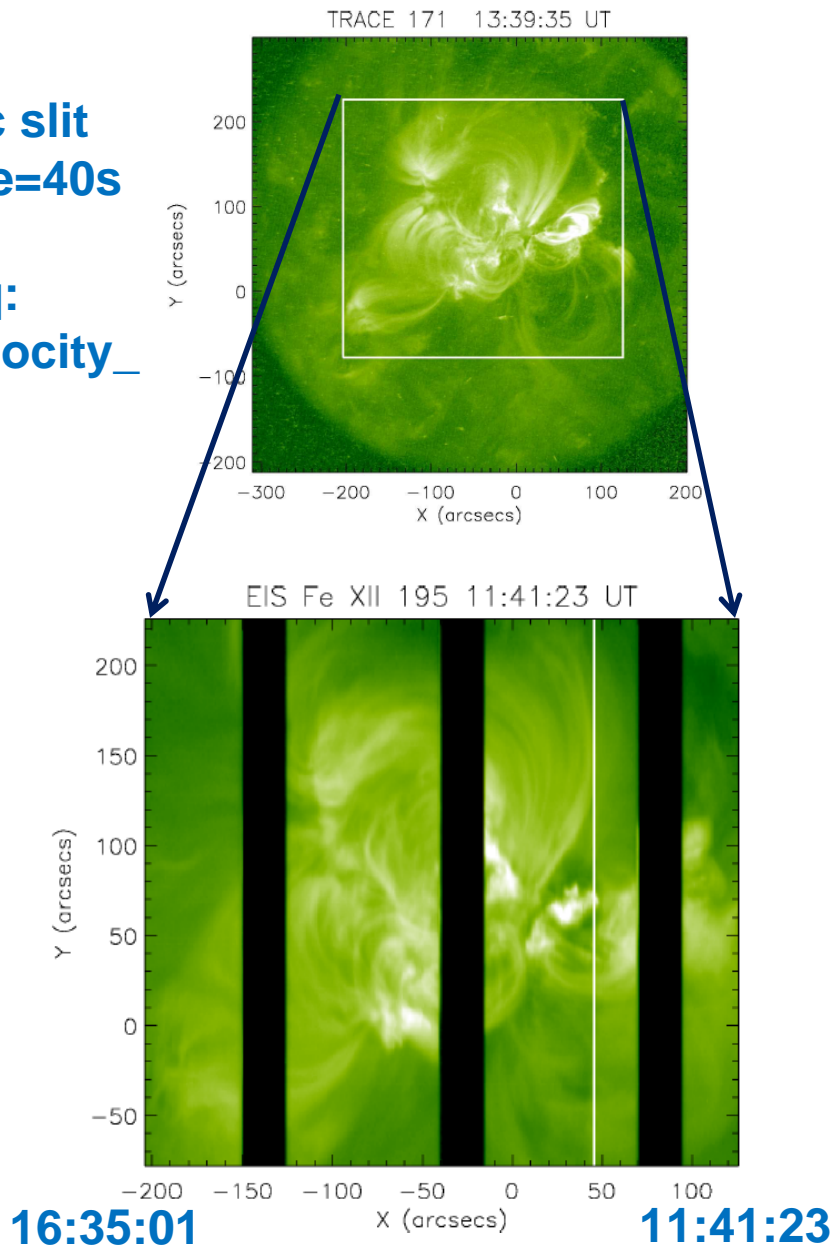
- **Loops are fundamental building blocks of active regions:**
  - Studying coronal loops will help to understand the mechanisms responsible for active region heating and coronal heating in general.
- **Questions of prime importance:**
  - Loop structures appear fuzzier at higher temperatures: is it real or just the matter of spatial resolution?
  - What are the plasma flow structures in the active region loops?
  - What are the values of densities and the filling factors of a coronal loop? How do they vary along the loop length?
  - Are active region loops **iso-thermal or multi-thermal** along the LOS? How does the temperature vary along the loop length?

- We can measure flows with a precision of  $\pm 3$  km/s in spatially resolved coronal structures simultaneously at different temperatures from **log T = 5.6 MK (Fe VIII) to log T = 6.4 (Fe XV)**.
- We can measure electron density more precisely than ever before and therefore, filling factors at different temperatures simultaneously.
- We can study temperatures along the LOS using emission measure analysis.

# EIS observations: May 19, 2007

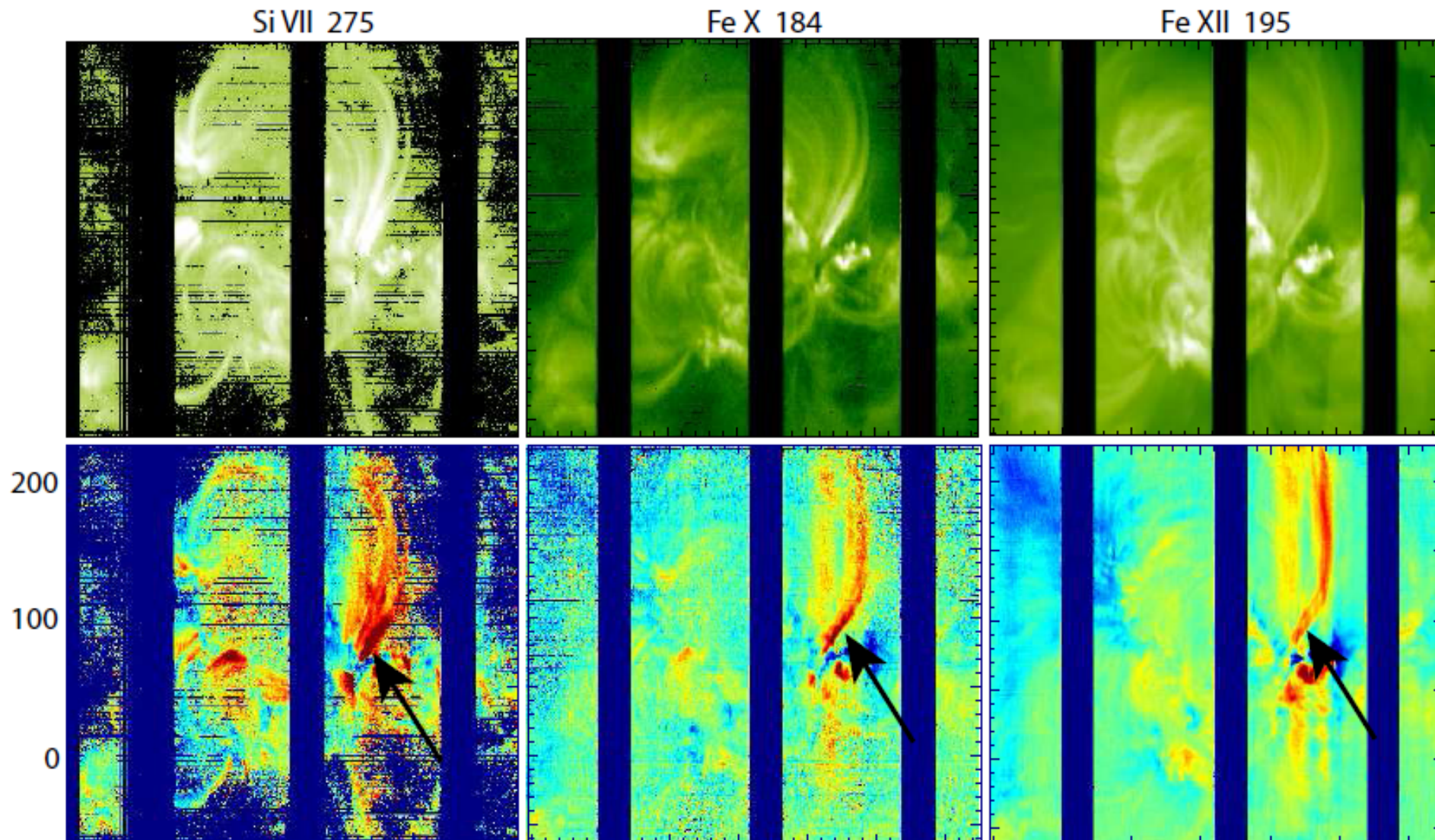
1 arcsec slit  
Exp time=40s

Obs seq:  
"AR\_velocity\_ maps"

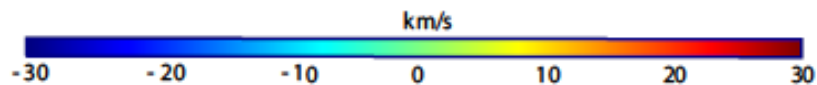


Ion / wavelength	log T [MK]	log Ne (cm <sup>-3</sup> )
Fe VIII (185)	5.6	
Mg VII (278)	5.8	<b>8.0 – 10.5</b>
Mg VIII (280)	5.8	
Si VIII (275)	5.8	
Fe X (184)	6.0	
Fe XII (195)	6.1	<b>7.0 – 12.0</b>
Fe XII (186)	6.1	
Si X (258)	6.1	<b>8.0 – 9.8</b>
Si X (261)	6.1	
Fe XIII (202)	6.2	<b>8.0 – 10.5</b>
Fe XIII (203)	6.2	
Fe XIV (264)	6.3	<b>8.5 – 11.0</b>
Fe XIV (274)	6.3	
Fe XV (284)	6.4	

# Overall intensity and flow structure

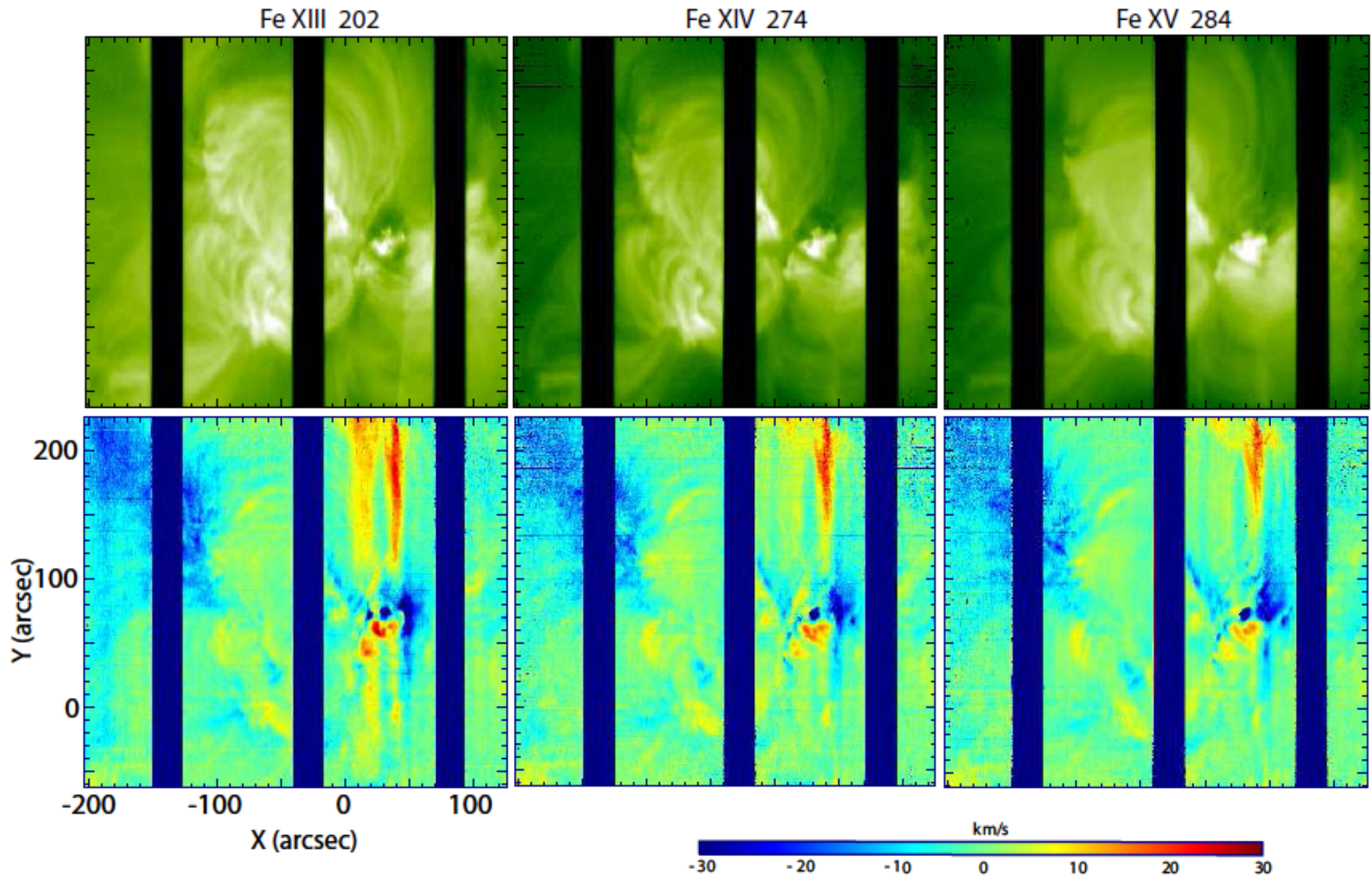


Velocity Range =  $\pm 30$  km/s

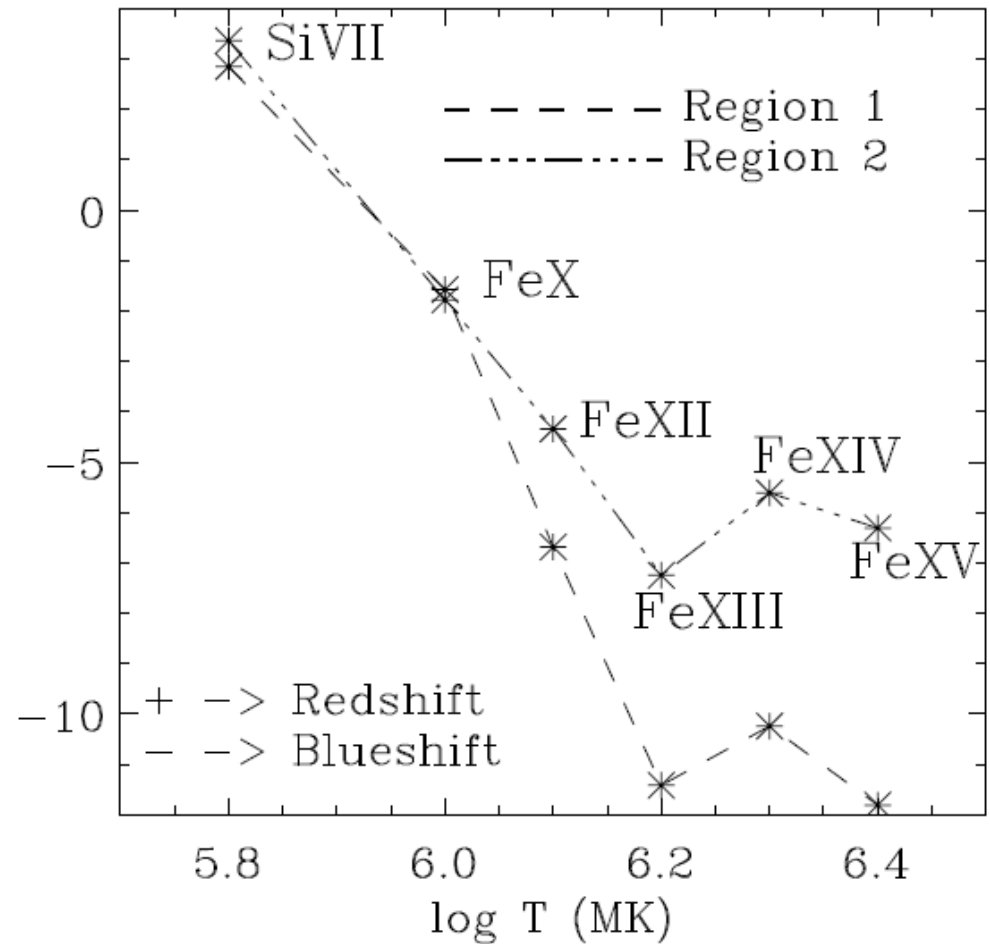
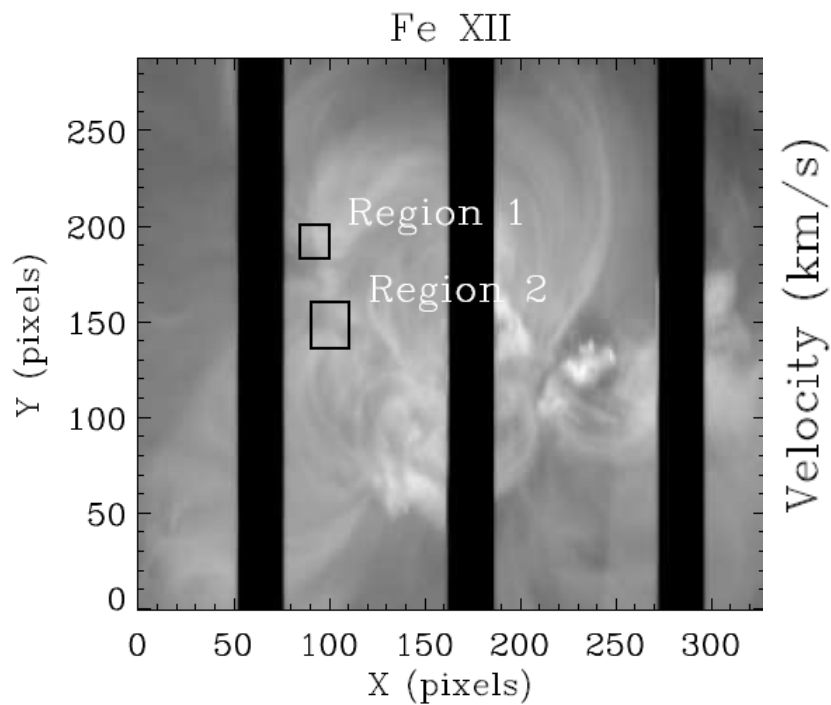




# Overall intensity and flow structure



Velocity Range =  $\pm 30$ km/s



**The foot point regions show red-shifted emission at lower temperatures. However, blue shifts dominate at high temperatures.**

# Intensity variation across loop structures

log T = 5.8

log T = 6.0

log T = 6.1

log T = 6.2

log T = 6.3

log T = 6.4

Si VII

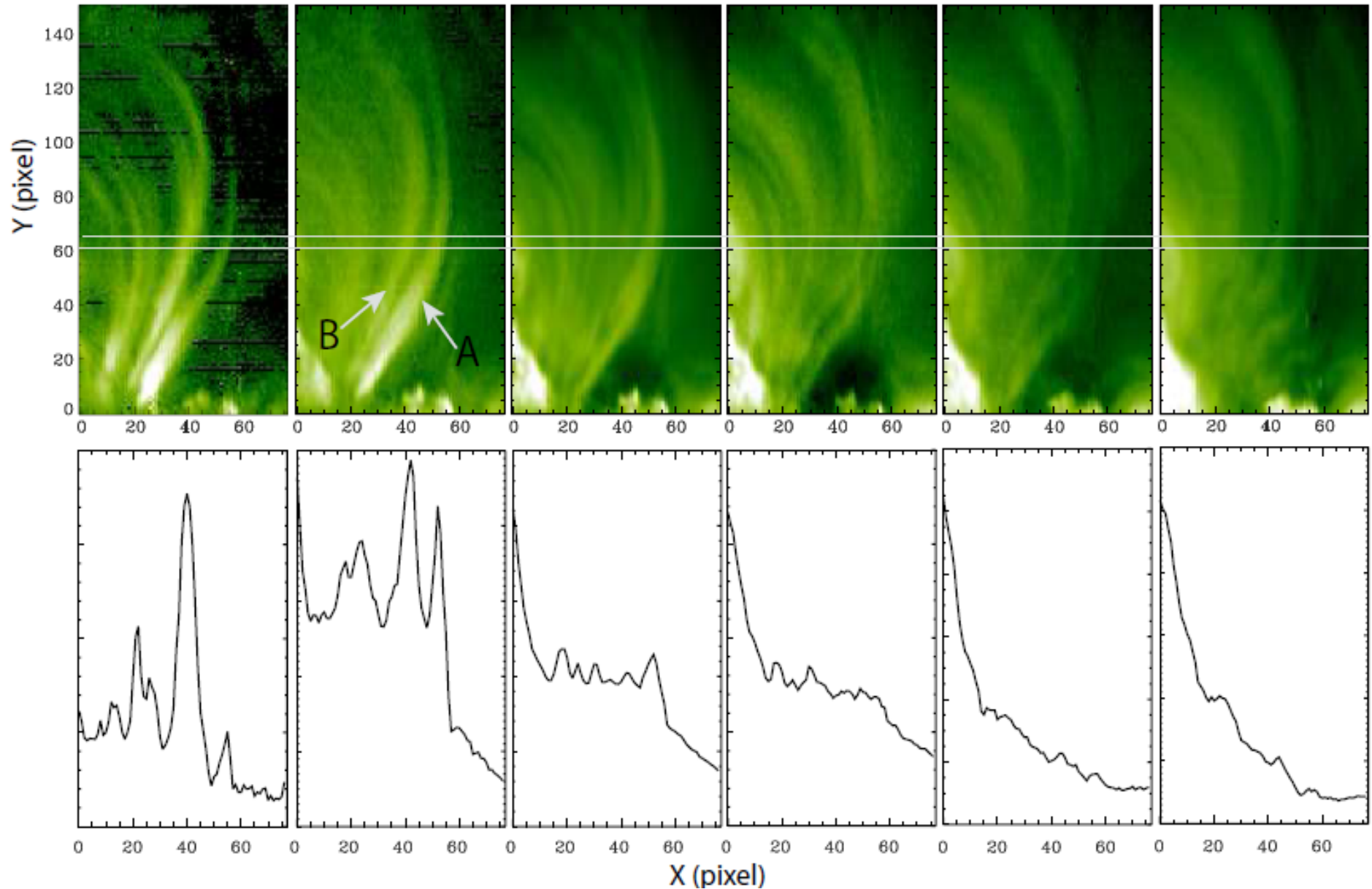
Fe X

Fe XII

Fe XIII

Fe XIV

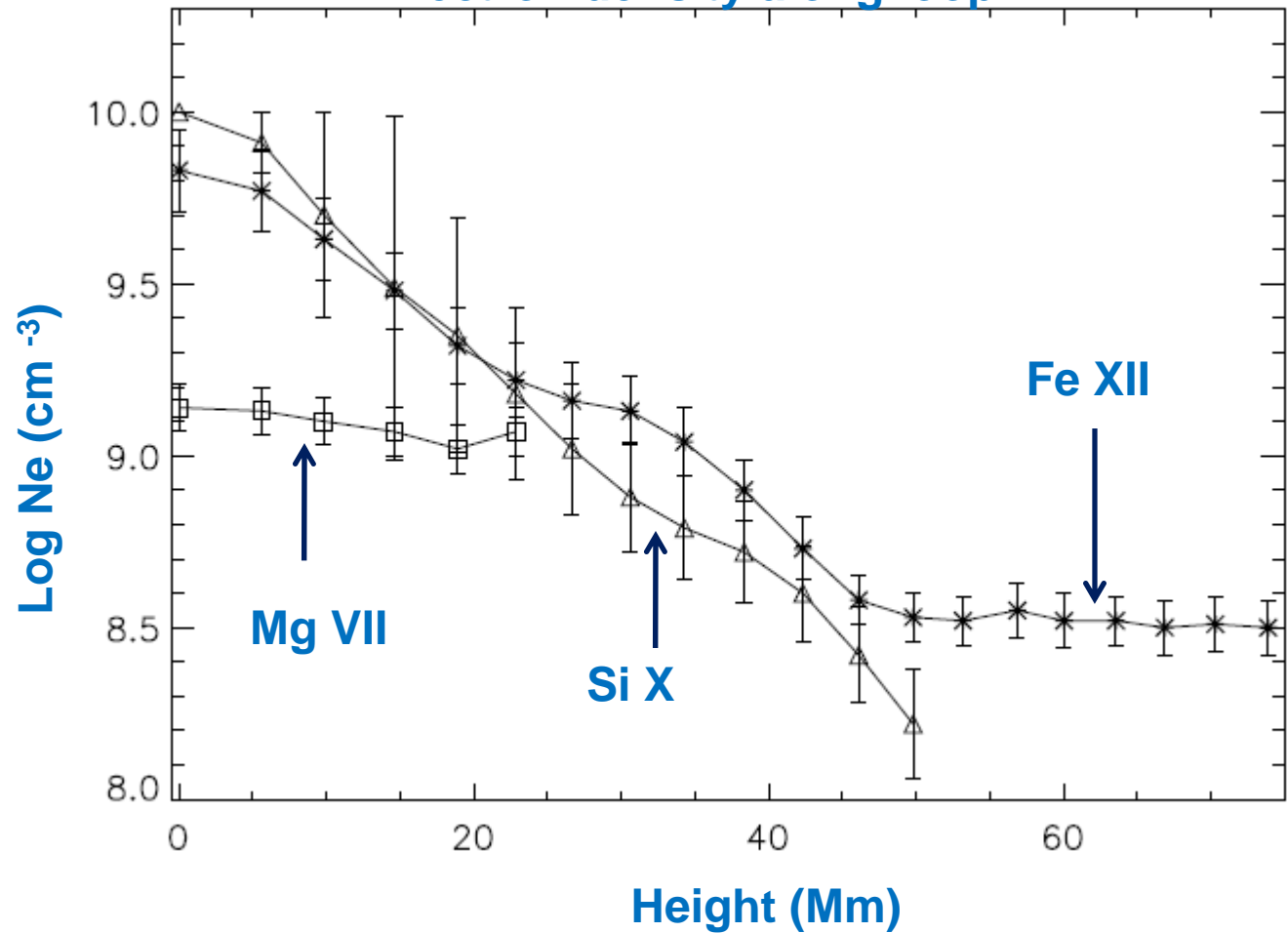
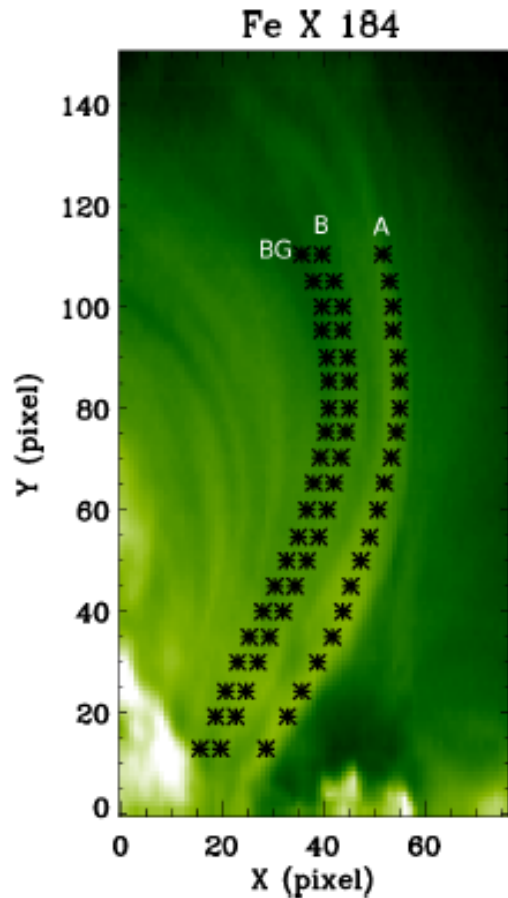
Fe XV





# Densities along the loop

## Electron density along loop A



BG → Background

Similar values for densities were obtained for loop B.

# Filling factor along the loop

Filling factor  $\Phi = EM / (h \cdot N_e^2)$

(Porter & Klimchuk (1995), Cargill & Klimchuk (1997))

Height (Mm)	Fe XII			Si X		Mg VII		
	$h$ (")	$N_e$ $\text{cm}^{-3}$	$\phi$	$N_e$ $\text{cm}^{-3}$	$\phi$	$h$ (")	$N_e$ $\text{cm}^{-3}$	$\phi$
0	7	9.83	0.02	10.00	0.00	20	9.14	1.79
5	8	9.77	0.02	9.91	0.01	17	9.13	1.71
9	6	9.63	0.06	9.70	0.06	15	9.10	1.48
14	10	9.48	0.07	9.49	0.14	15	9.07	0.84
18	10	9.32	0.15	9.35	0.26	15	9.02	0.56
22	13	9.22	0.20	9.18	0.43	10	9.07	0.28
26	12	9.16	0.29	9.02	1.03	...	...	...
30	11	9.13	0.37	8.88	2.11	...	...	...
34	11	9.04	0.47	8.79	2.86	...	...	...
38	10	8.90	0.79	8.72	3.25	...	...	...
42	11	8.73	1.30	8.60	4.14	...	...	...
46	10	8.58	2.63	8.42	7.40	...	...	...
49	7	8.53	4.52	8.22	18.73	...	...	...
53	12	8.52	3.11	...	...	...	...	...
56	10	8.55	3.66	...	...	...	...	...
60	10	8.52	4.21	...	...	...	...	...
63	9	8.52	4.57	...	...	...	...	...
66	11	8.50	4.41	...	...	...	...	...
70	12	8.51	3.80	...	...	...	...	...
73	12	8.50	3.89	...	...	...	...	...

# Temperature: EM-Loci

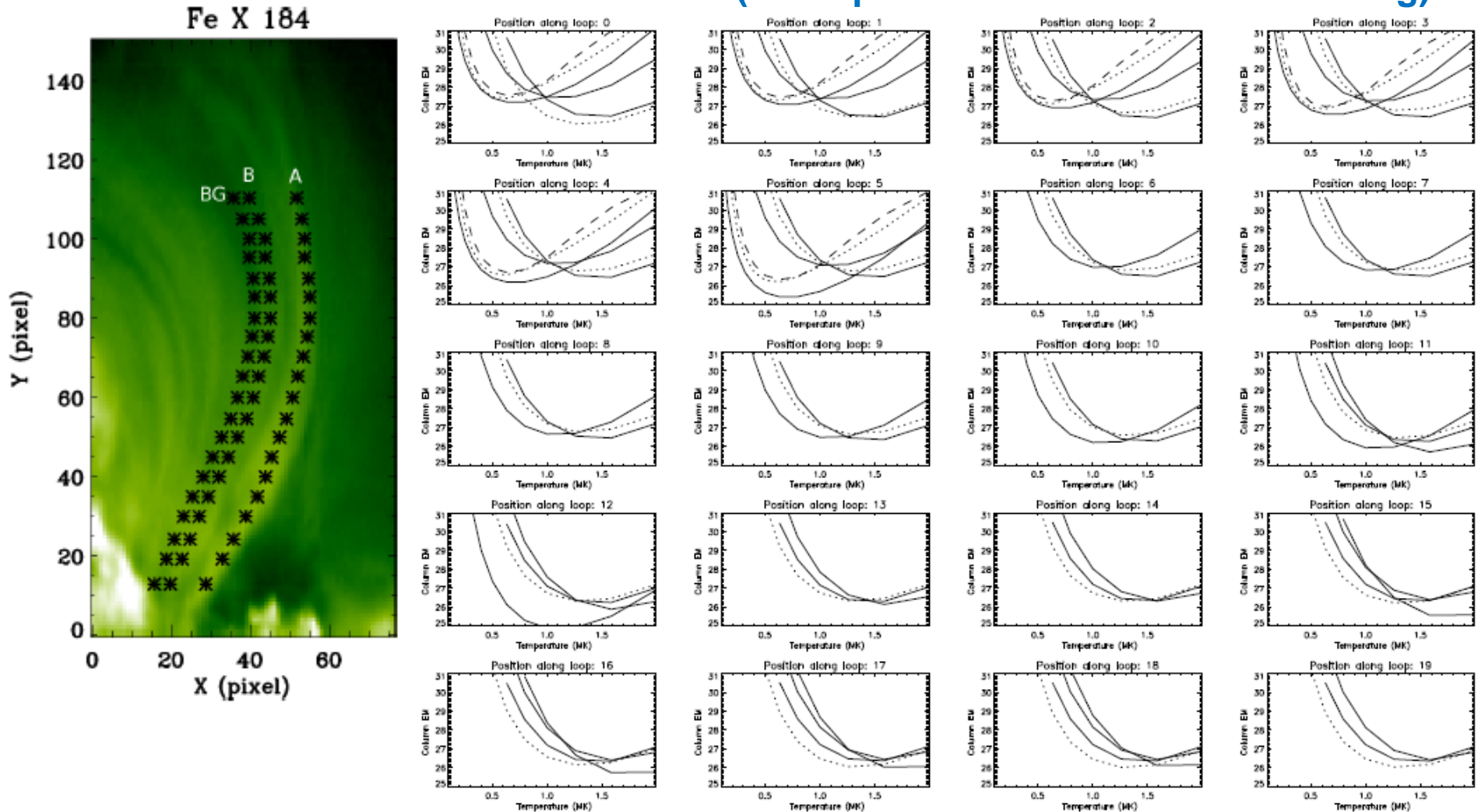
$$EM [T_e] = I_{obs} / [A(b) G(N_e, T_e)]$$

$I_{obs}$  = Background removed observed intensities

$A(b)$  = Abundances (Coronal Abundance)

$G(N_e, T_e)$  = Contribution function

(Ion Equilibrium: Arnaud & Rothenflug)



Warren et al. (2008) showed that the loops are “mildly multi-thermal”.

- In lower temperature lines red shifts (downflows) are seen all along the loop being strongest at foot points.
- For higher temperatures red shifts are seen near the loop top, but turn to blue shifts towards the foot point.
  - **This is indicative of hot material evaporating and cool material condensing.**



- The  $N_e$  along the loop decreased with increasing height. At a given location on the loop we obtained different densities, using spectral lines forming at different temperatures.
- We obtained a very low value of filling factor (0.02 – 0.05) at  $\log T = 6.1$  MK and a filling factor close to 1 at  $\log T = 5.8$  MK towards the foot point of the loop.
- The plasma in the loop along the LOS is “Mildly” MULTI-THERMAL (at least towards the foot point). The temperature increased from 0.8 to 1.5 MK from foot point to loop top.
  - **The loop is comprised of many unresolved strands which are being heated separately (?)**