

# Solar irradiance variability and climate

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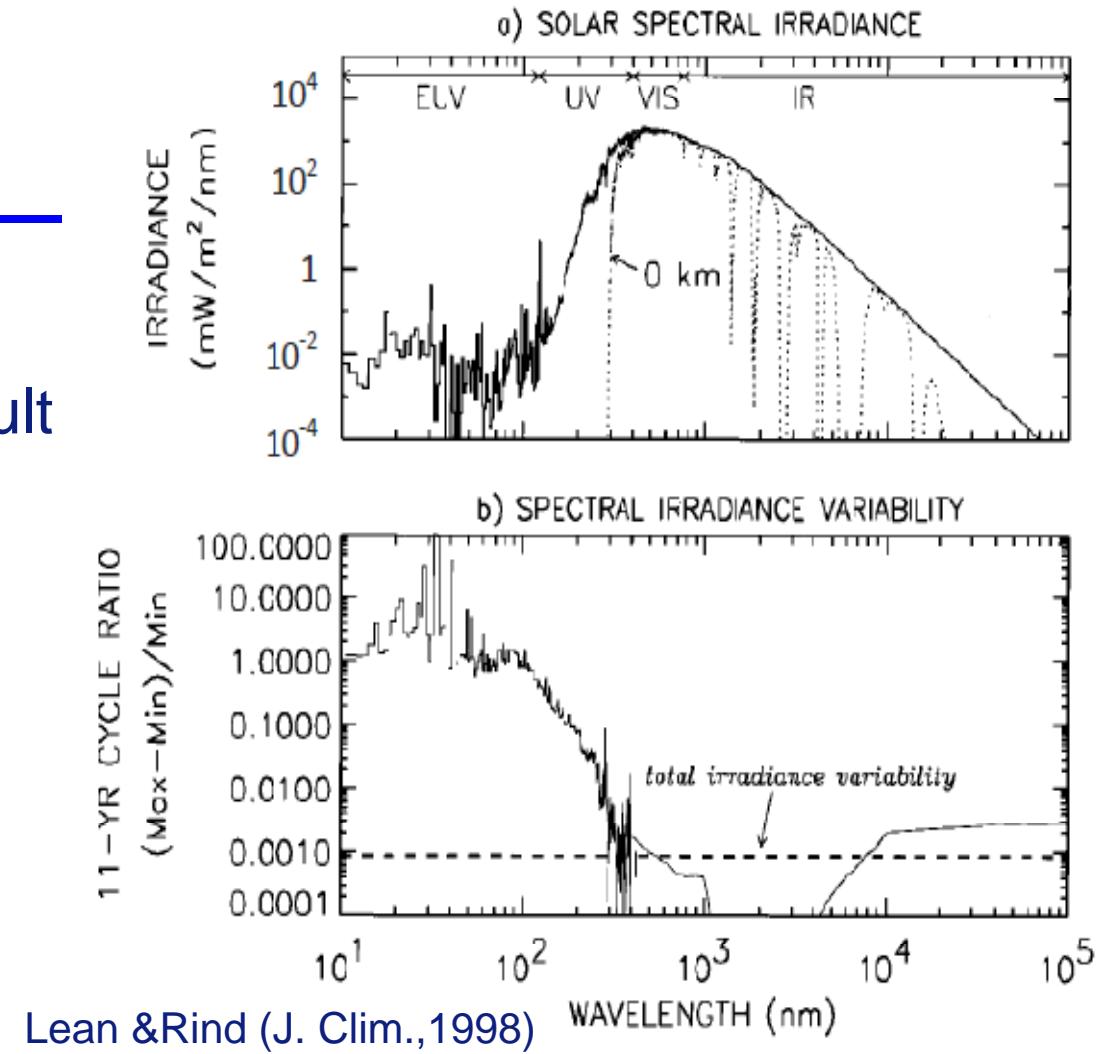
Department of Physics  
Imperial College London, UK

Eddy Symposium Aspen CO 22-24 Oct 2010

# Background

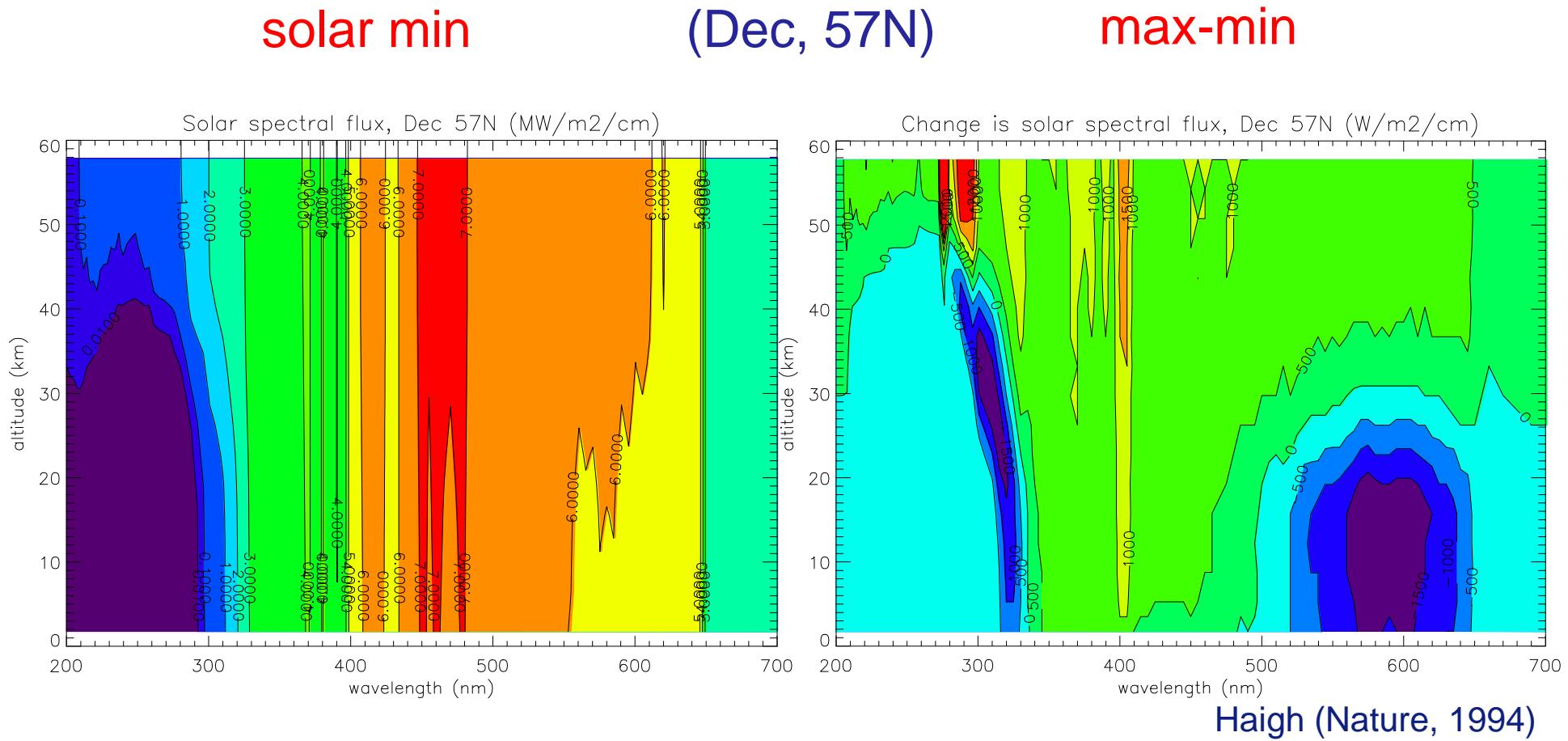
- Evidence for solar influence on climate difficult to explain by small variations in total solar radiative output.

- UV variations (fractionally) much larger.



- Visible wavelength radiation mainly reaches Earth's surface.
- UV radiation mainly absorbed in stratosphere.

# Background: Solar cycle variation in spectral irradiance



Solar irradiance reaching lower atmosphere depends on zenith angle and on response in stratospheric ozone.

# Solar forcing of climate: radiative mechanisms

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“Top-down” via UV heating the stratosphere

and/or

“Bottom-up” via (visible) radiation warming surface ?

# Outline

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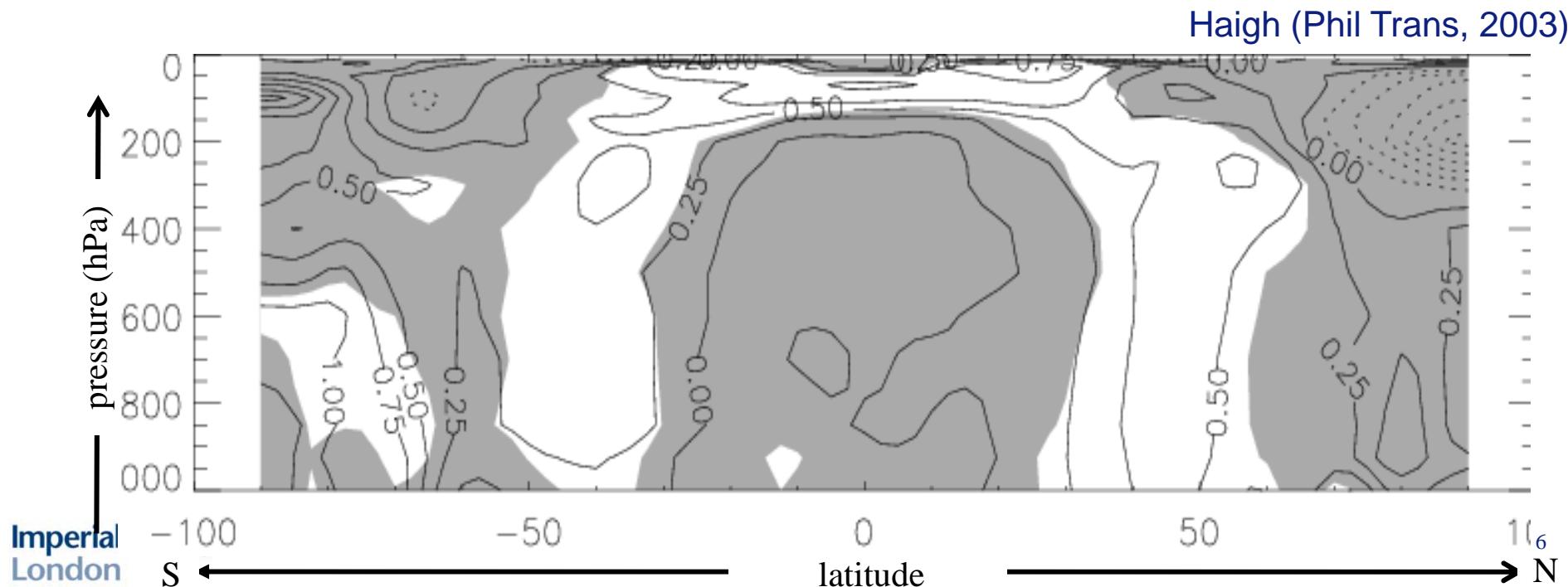
- Impact of solar variability on wind and circulation of troposphere: observations and models.
- Mechanism proposed via UV heating of stratosphere and dynamical coupling.
- Unusual behaviour of solar spectrum over declining phase of solar cycle 23.
- Implications for stratospheric composition and solar forcing of climate.

# Temperature changes over the 11-year solar cycle

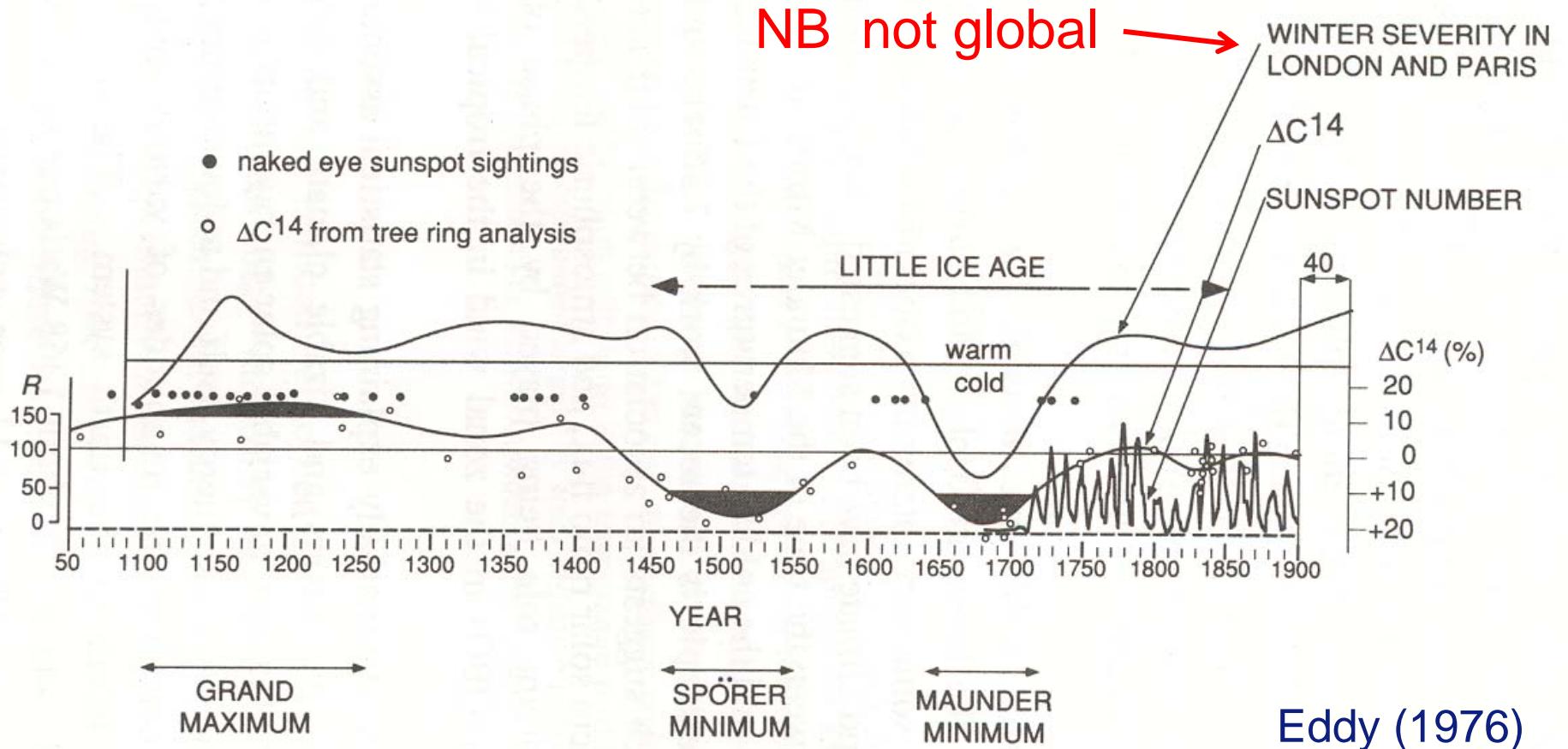
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- Non-uniform.
- Warming in equatorial stratosphere, less towards the poles.
- Bands of warming in mid-latitudes.

Multiple regression analysis of zonal mean temperature 1979-2000

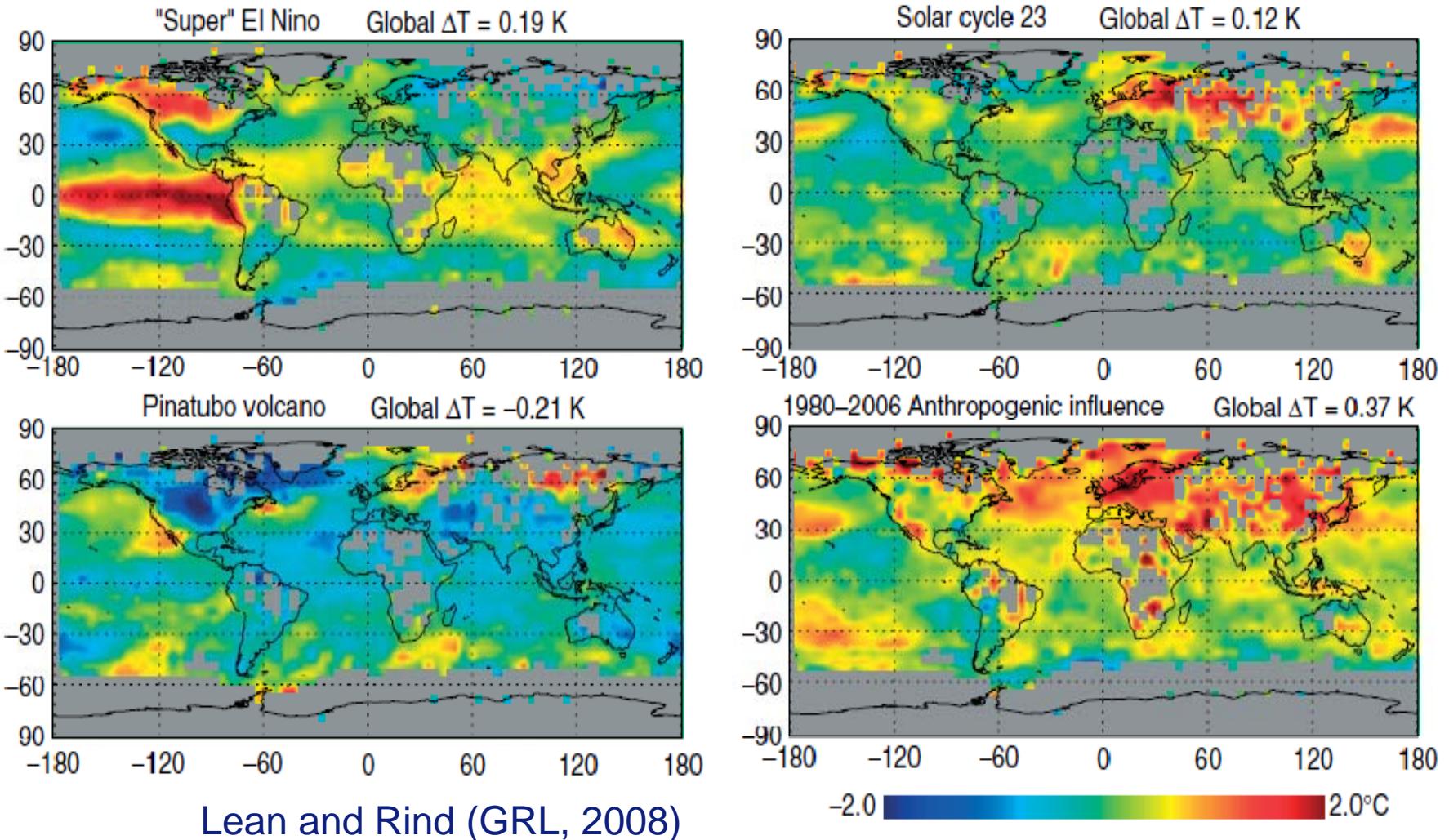


# Temperature in NW Europe



NB  $C^{14}$  produced by cosmic rays: more when Sun less active

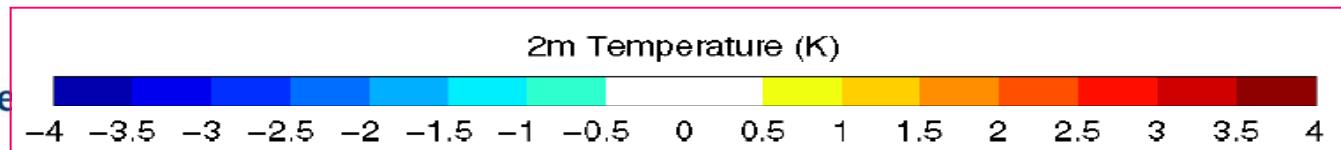
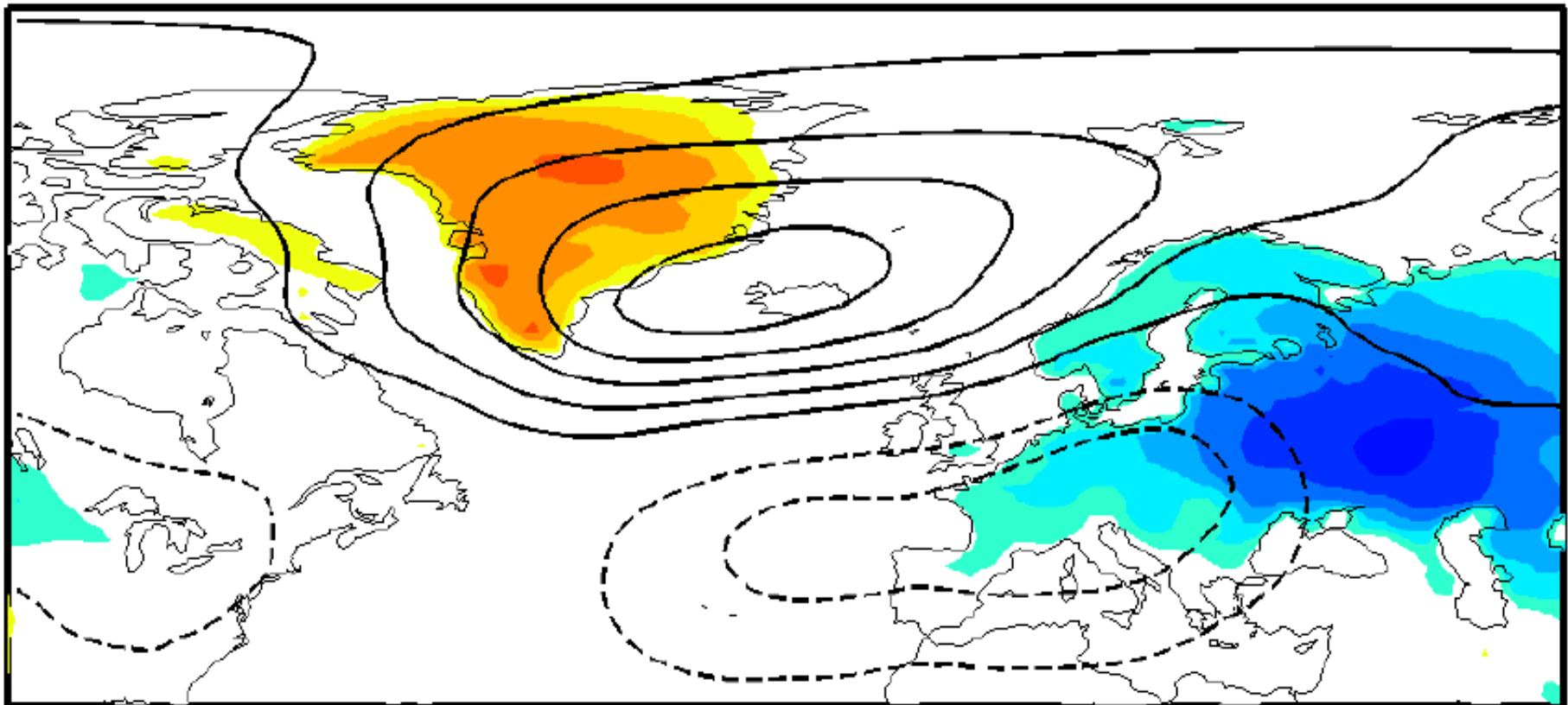
# Signals in surface air temperature



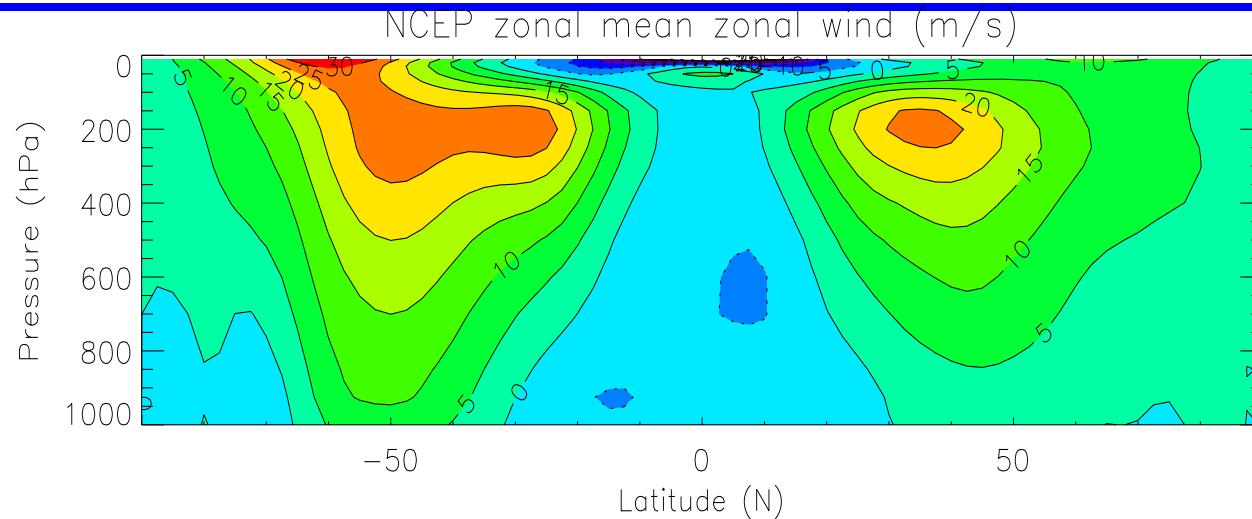
# Winter land surface air temperature and pressure

Lockwood et al. (ERL, 2010)

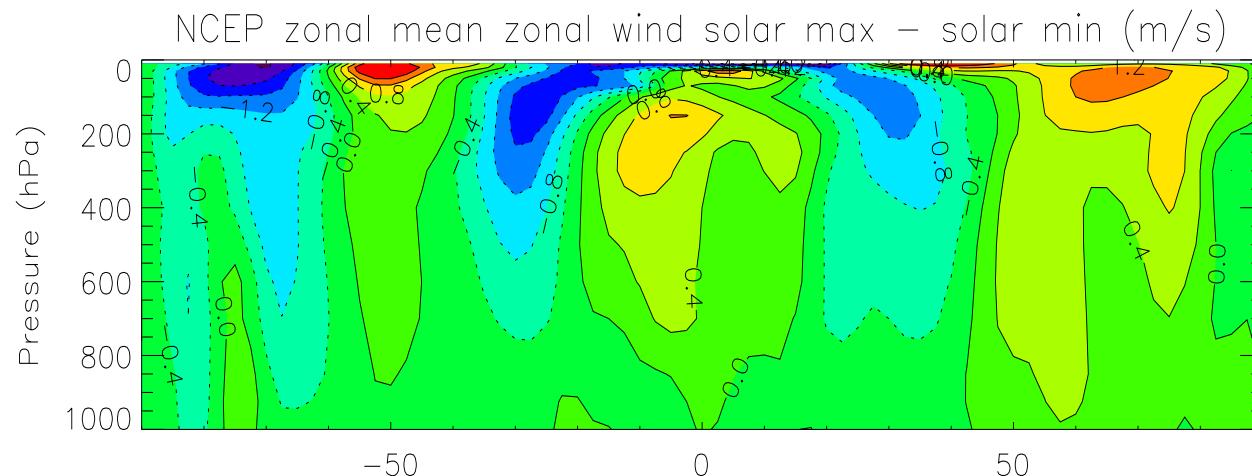
**SOLAR: LOW – HIGH**



# Solar cycle signal in westerly wind (observed)



mean



max-min

Weakening and poleward shift of the mid-latitude jets

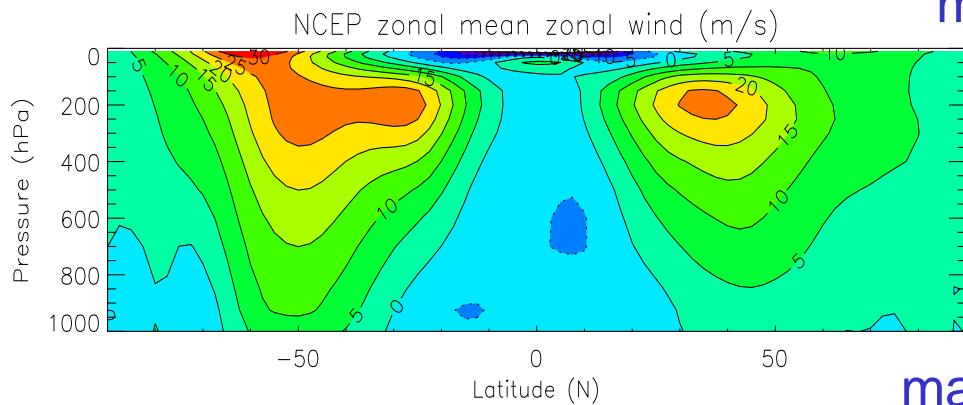
Im  
Lo

Haigh et al (J.Clim., 2005)

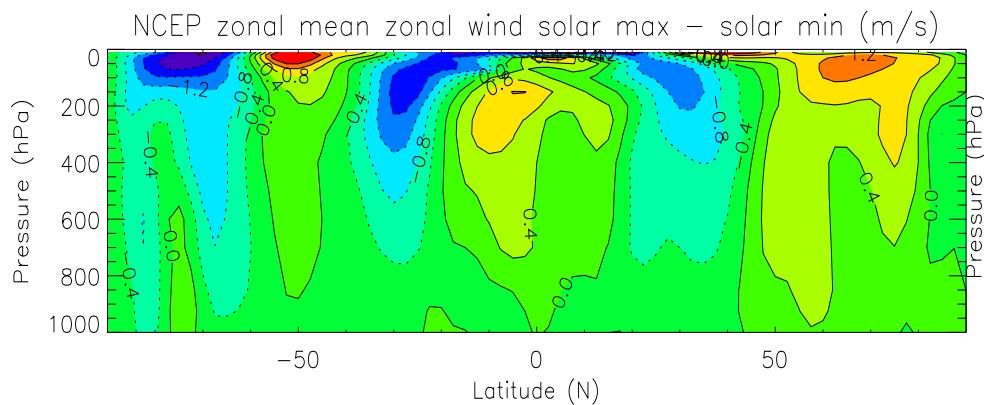
10

# Solar signal in zonal wind: obs cf climate model

Observations

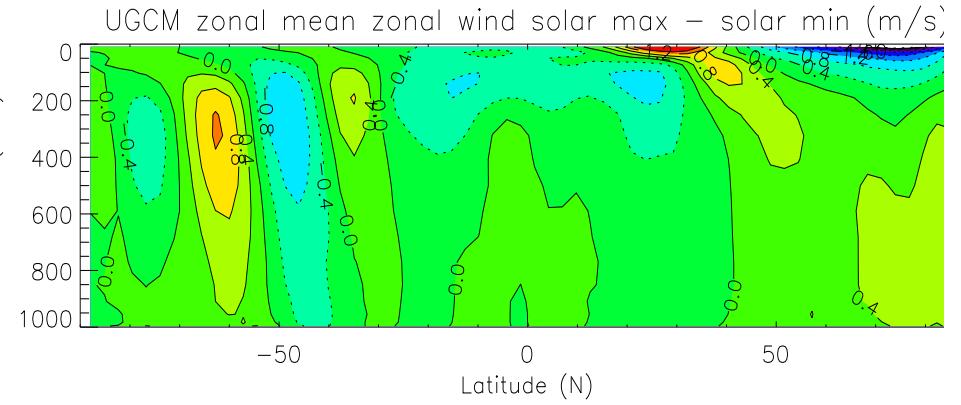
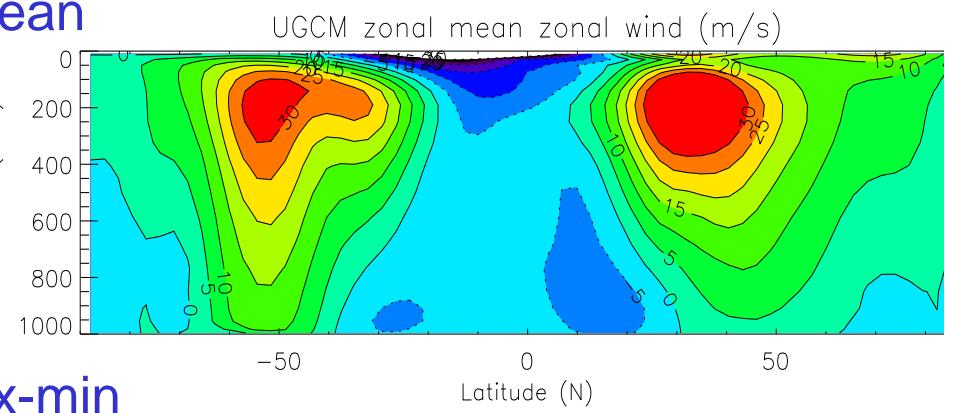


mean



max-min

GCM response to changes in solar radiation & O<sub>3</sub>

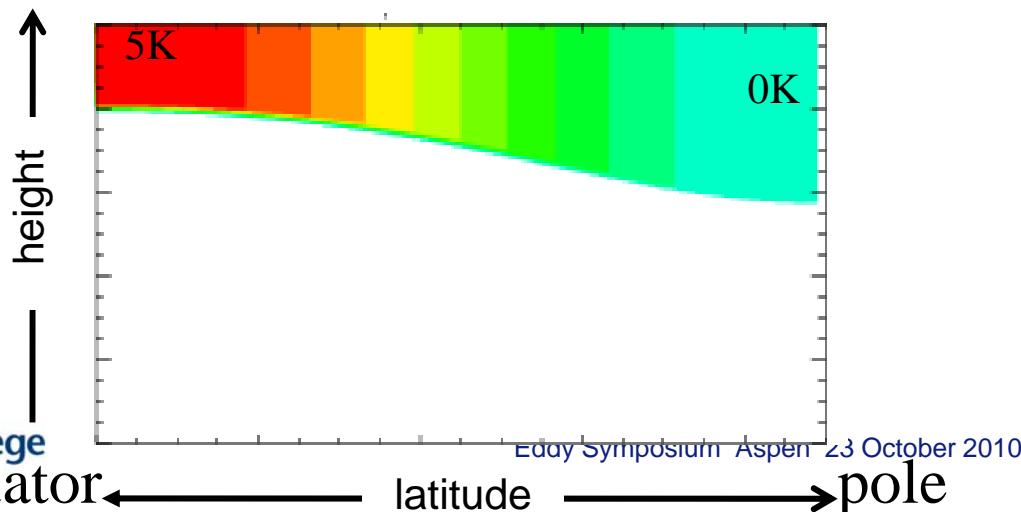


Haigh et al (J. Clim., 2005)

N.B. Similar pattern in response but much smaller amplitude without O<sub>3</sub> changes

# Study using a simplified climate model

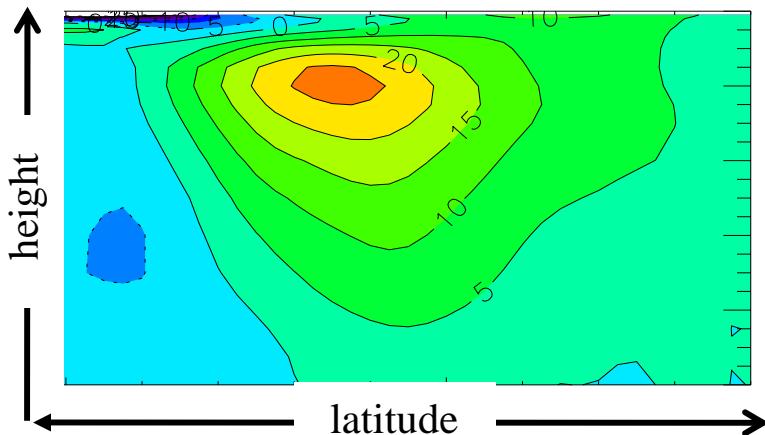
- Uses “Intermediate GCM”
  - full dynamics but highly simplified physics  
(so can try out ideas and do many runs cheaply)
  - no orography  
(so no planetary scale waves but still synoptic scale waves)
- Apply simplistic heating perturbation to the stratosphere:



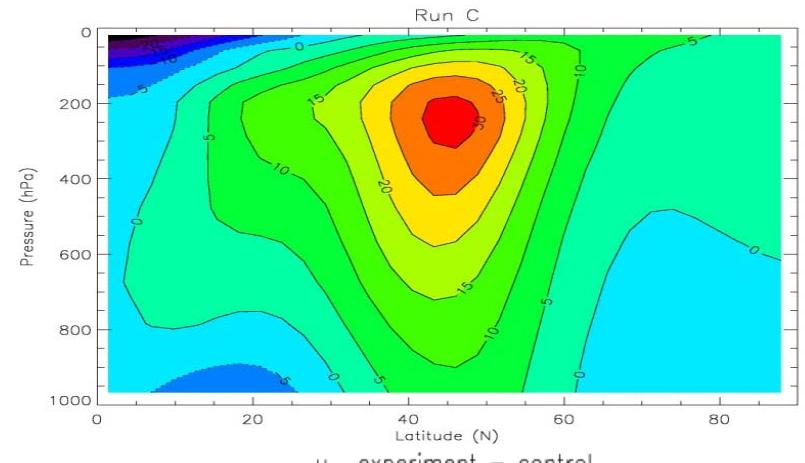
Haigh, Blackburn & Day (2005)

# Heating applied ONLY in the stratosphere:

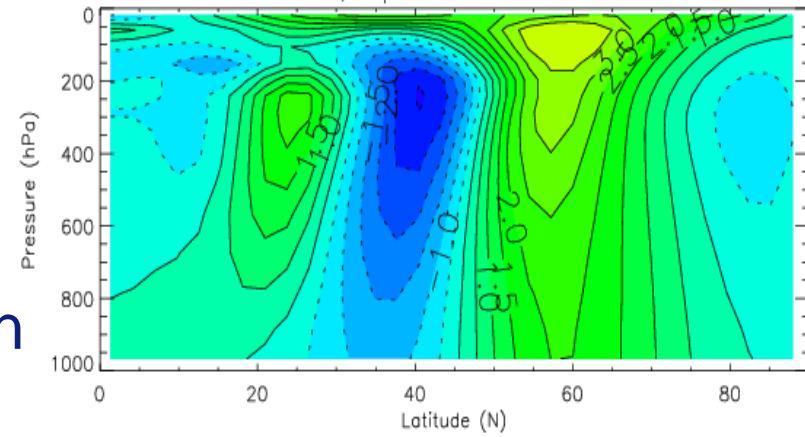
Observations of solar impact zonal wind Simple model



mean

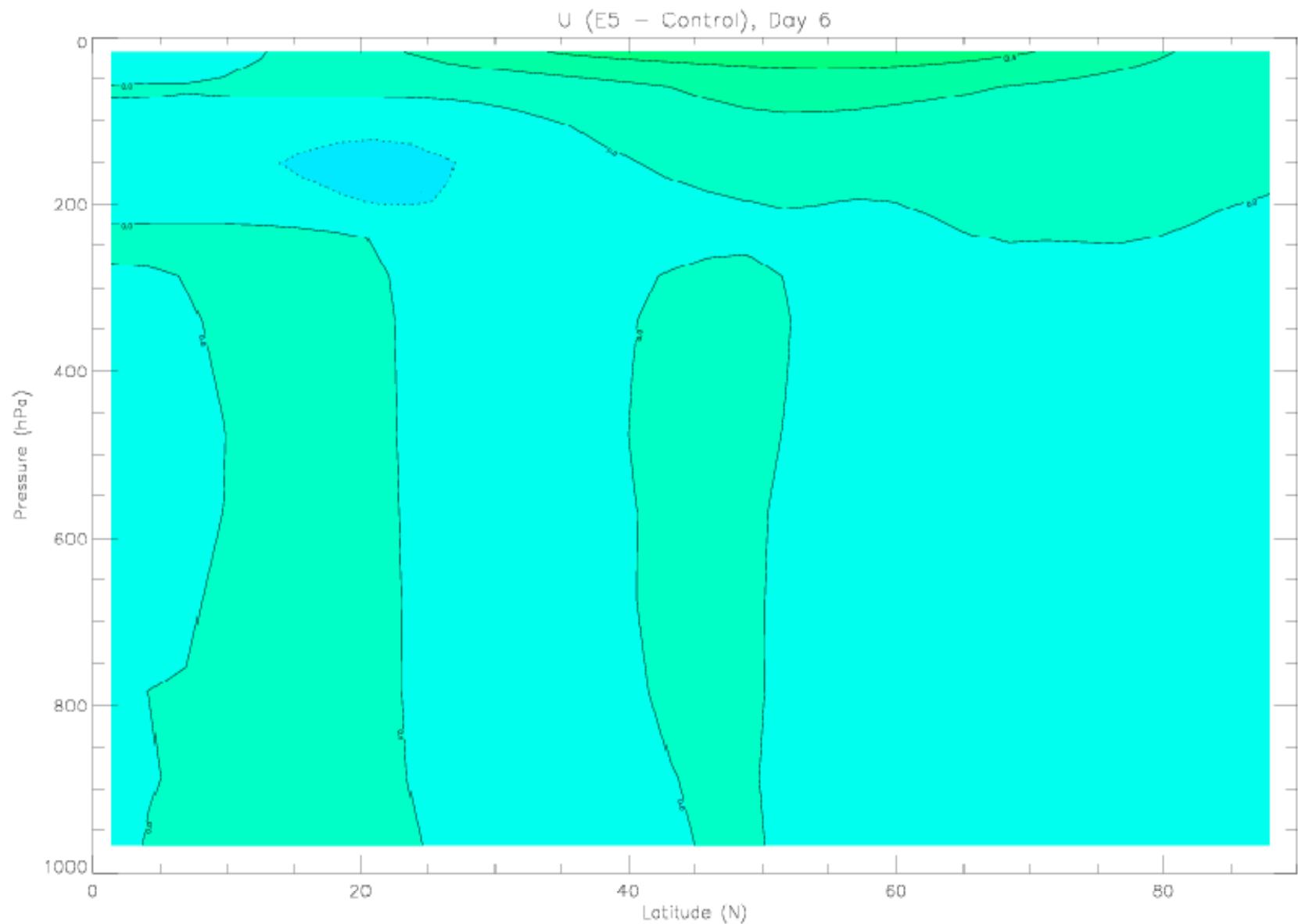


solar  
max - min

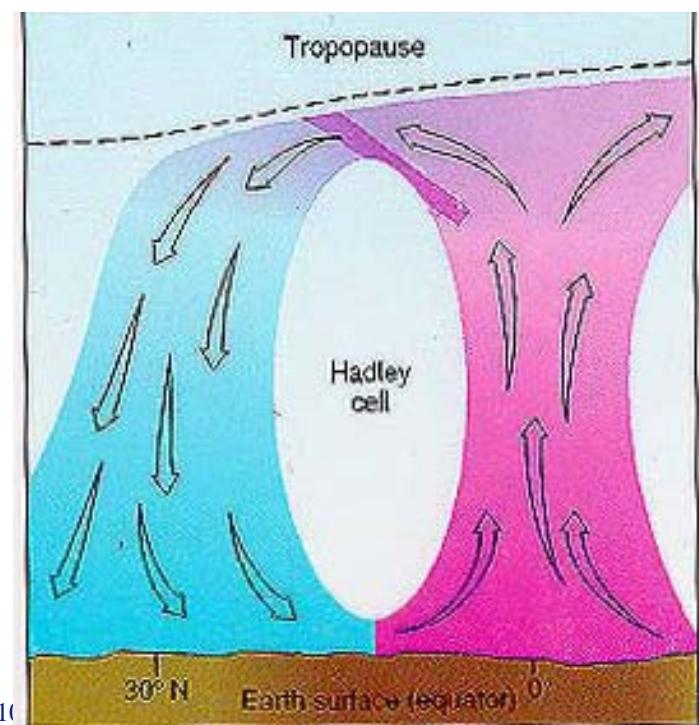
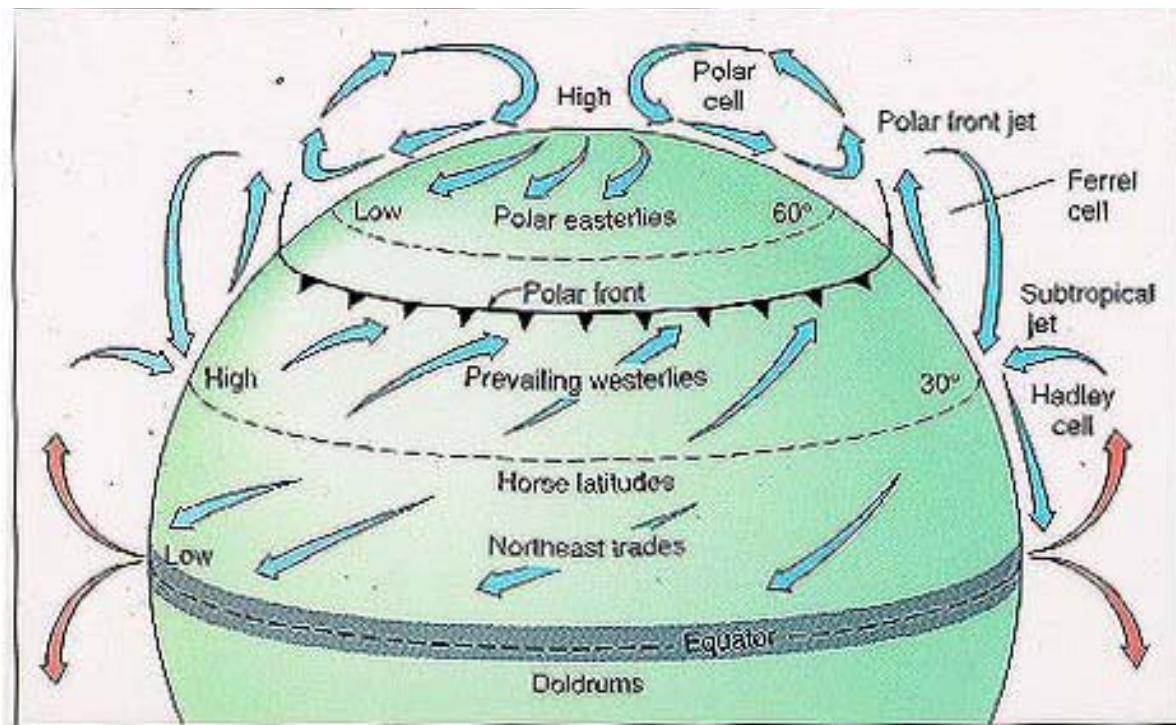


the jet weakens and moves polewards

# Evolution of zonal wind response 11-day running means



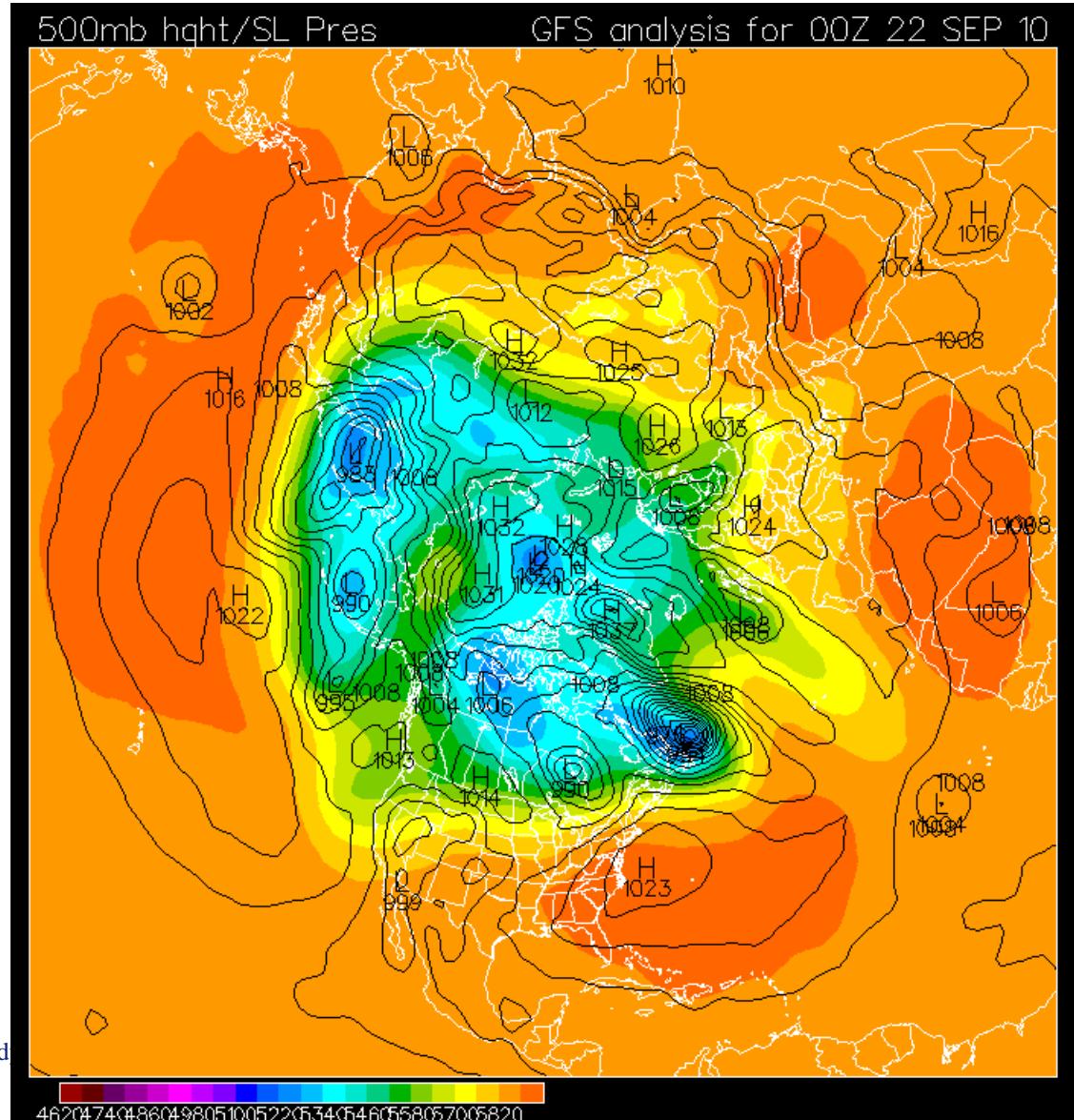
# Mean circulation of the tropical atmosphere: the Hadley cell



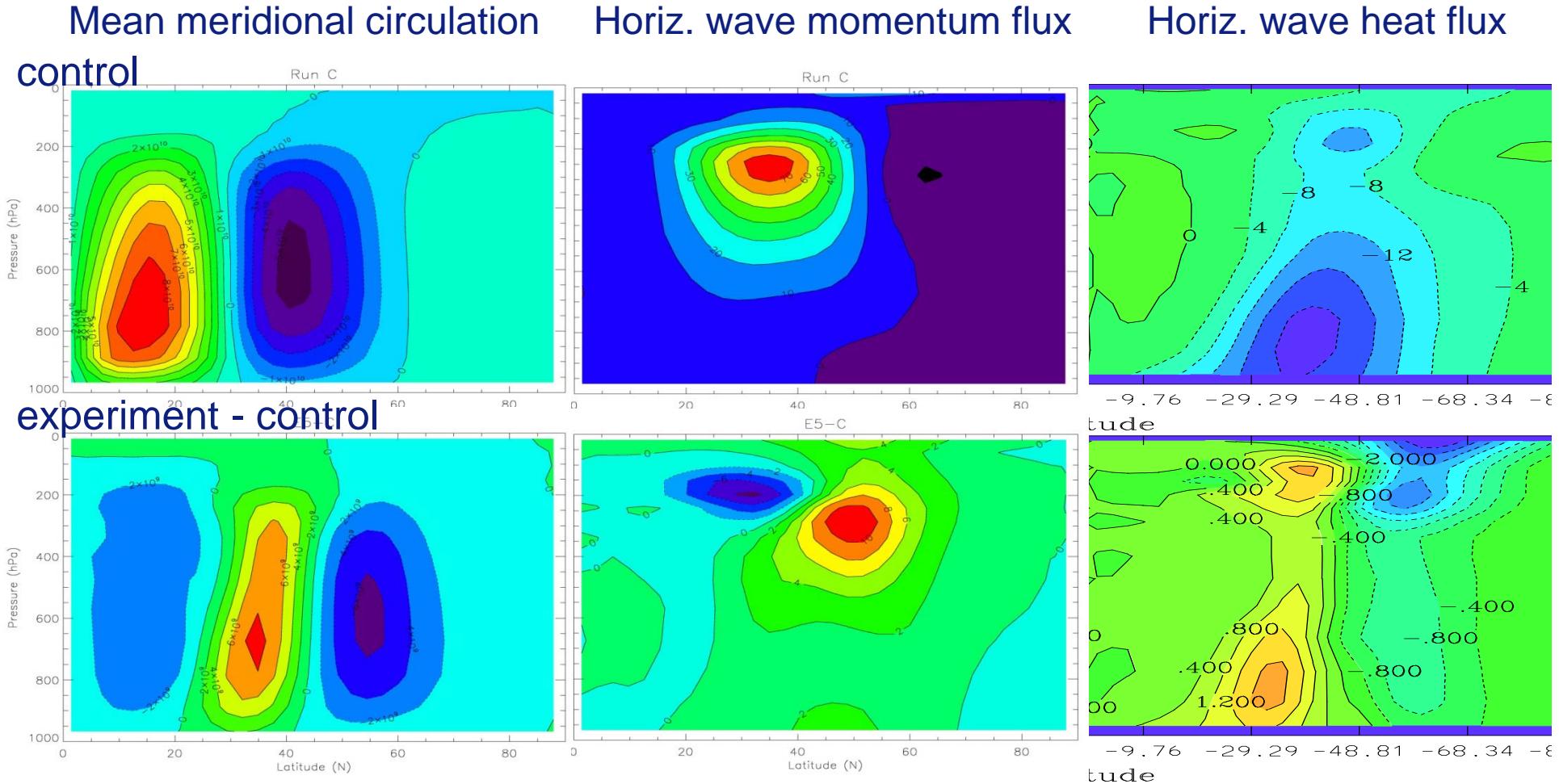
# Waves in the atmosphere

Contours: surface pressure (mb)

Colours: temperature of lowest 5.5km (approx)

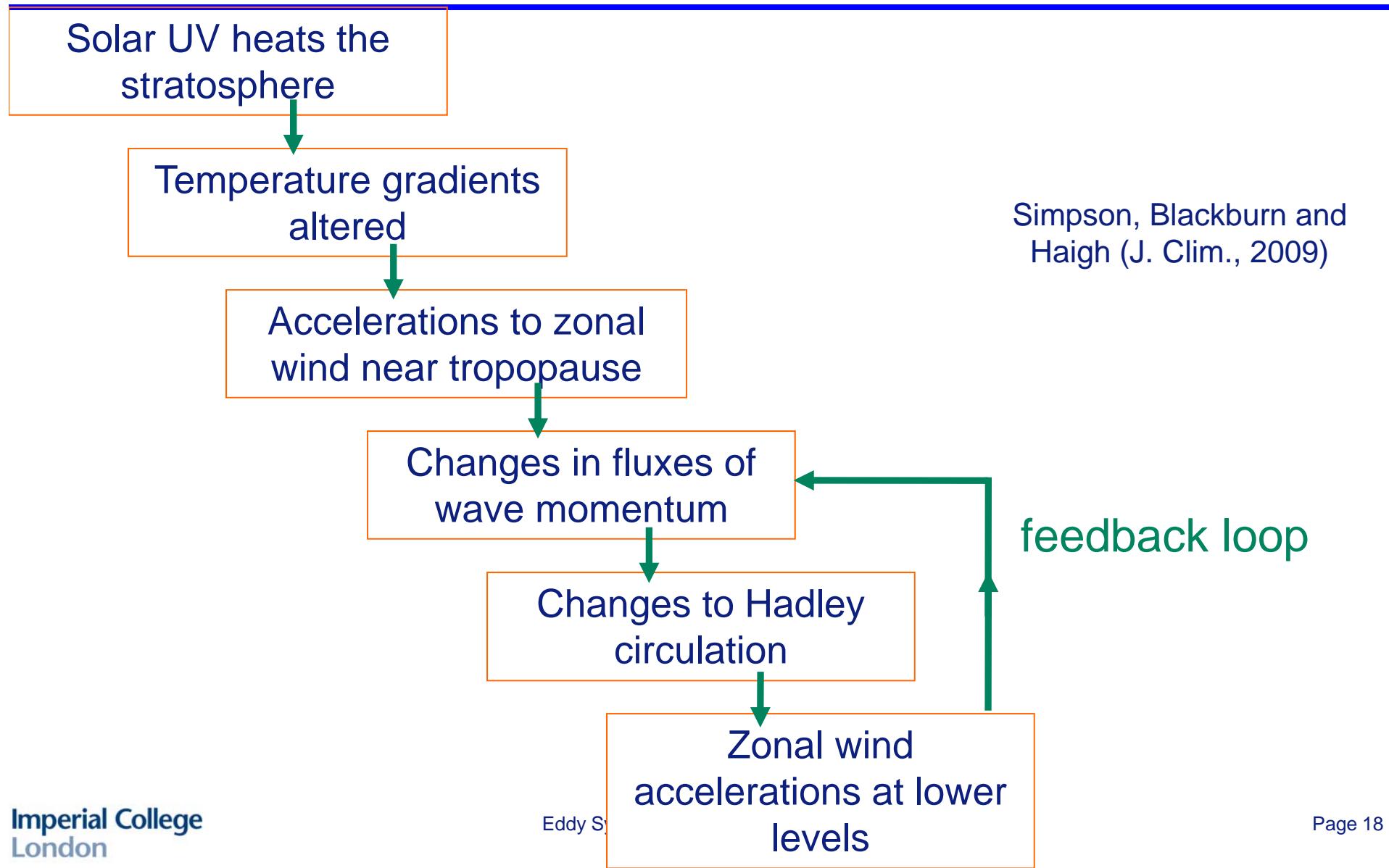


# Simple model equatorial heating (E5) results:



The Hadley cell weakens and expands; wave fluxes moves polewards

# Outline of mechanism:

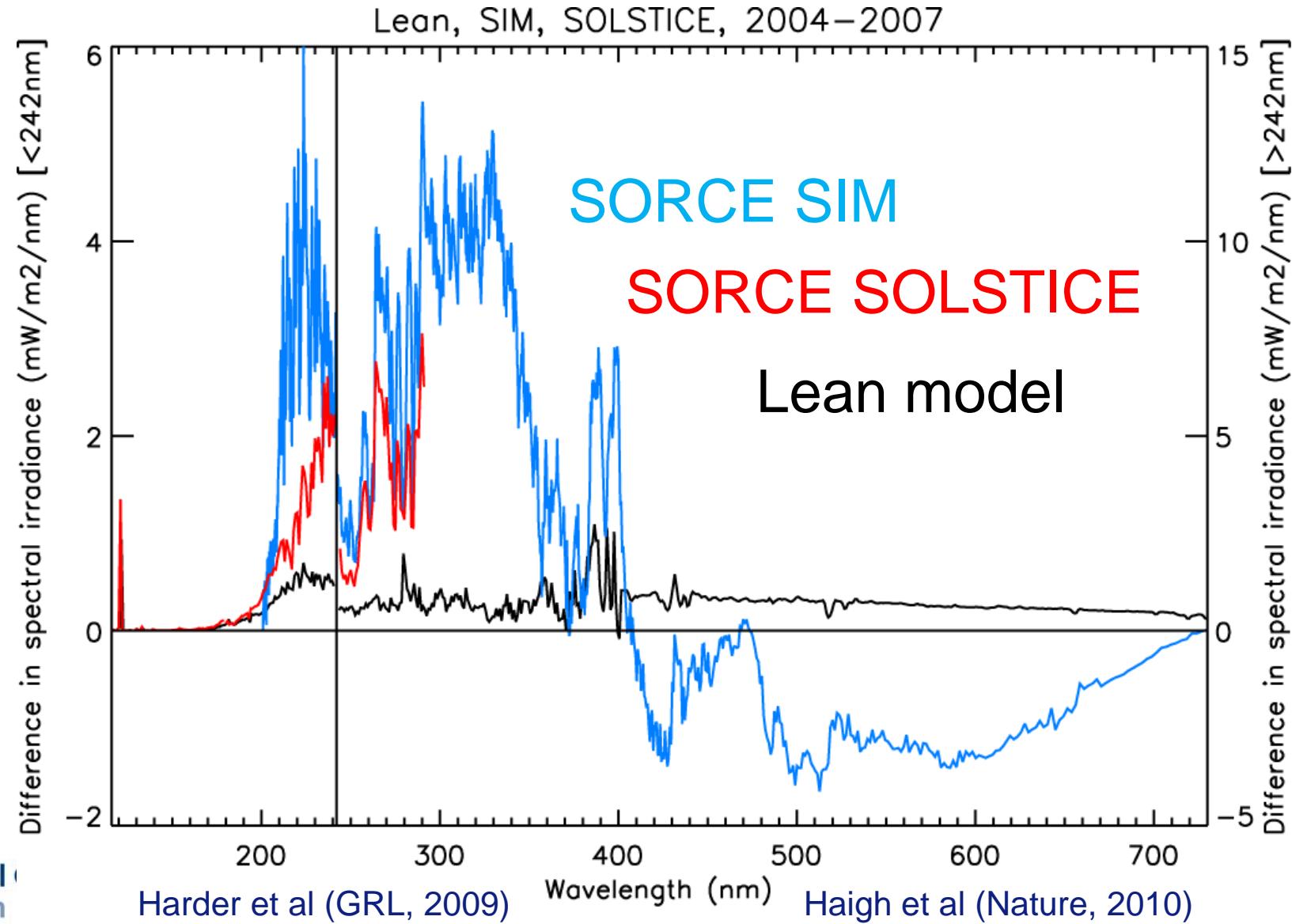


# Summary: Part I

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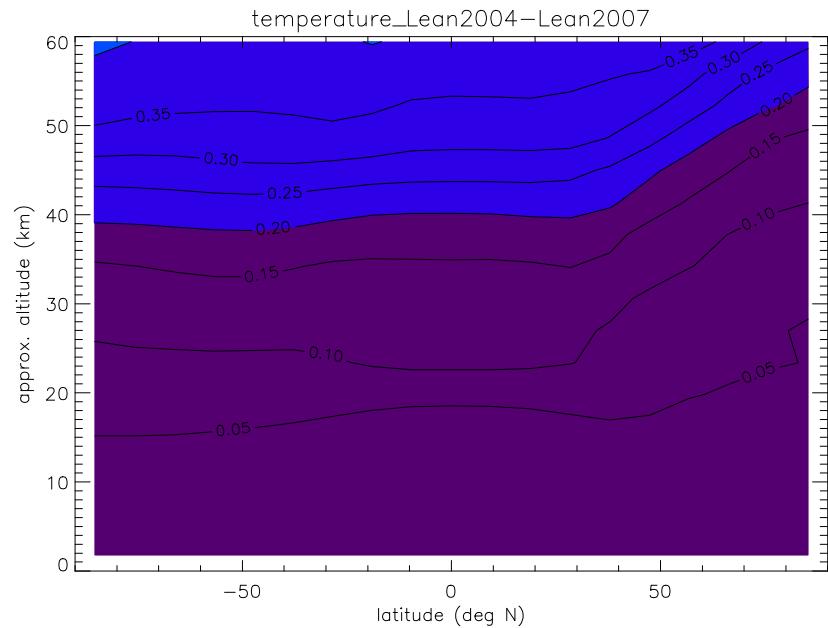
- (Solar) heating of the lower stratosphere produces changes in tropospheric circulation.
- Tropical heating produces a weakening and poleward shifts of the jets and a weakening and expansion of the Hadley cells.
- This results from the impact on the momentum budget of a feedback between vertically propagating, synoptic-scale waves and the mean flow.
- Crucially dependent on magnitude and location of stratospheric heating.

# Solar spectra: differences 2004-2007

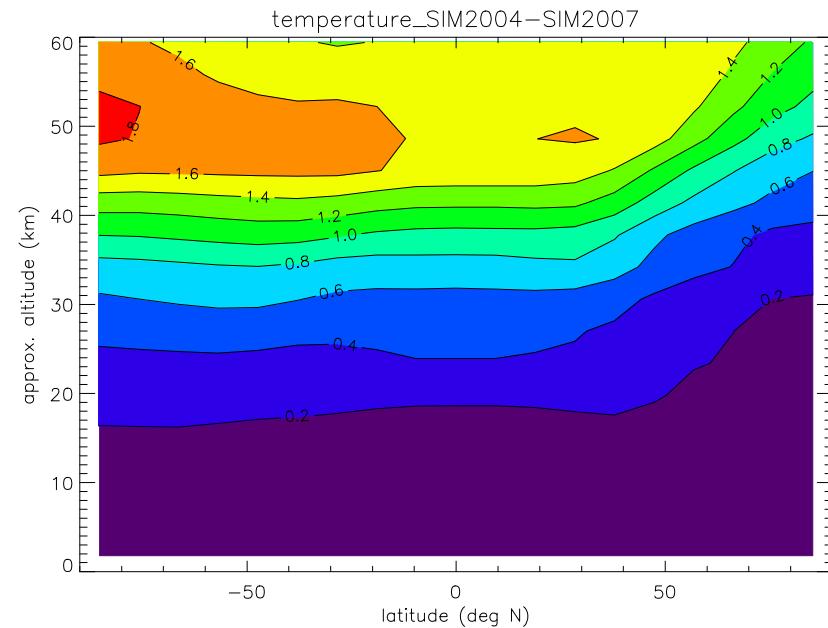


# 2D model temperature differences (K) 2004-2007

Lean



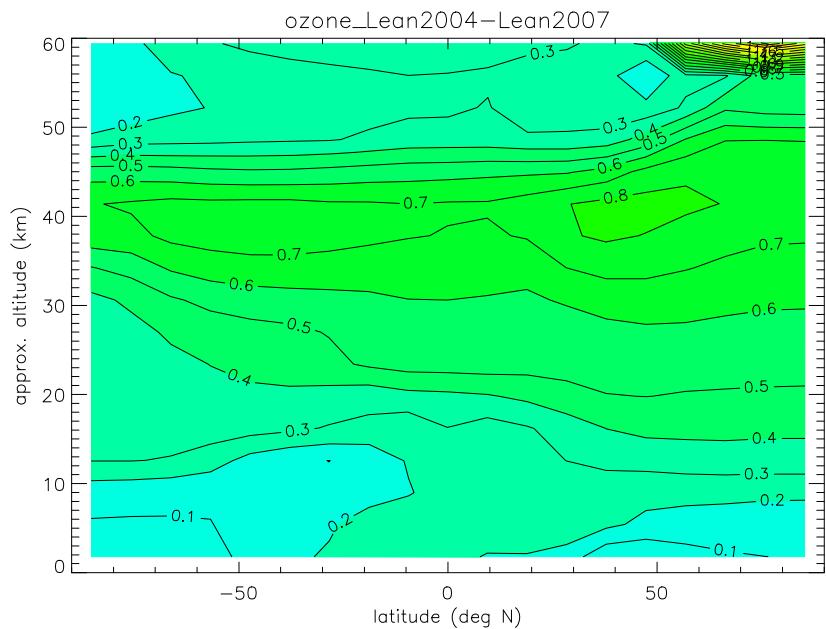
SIM



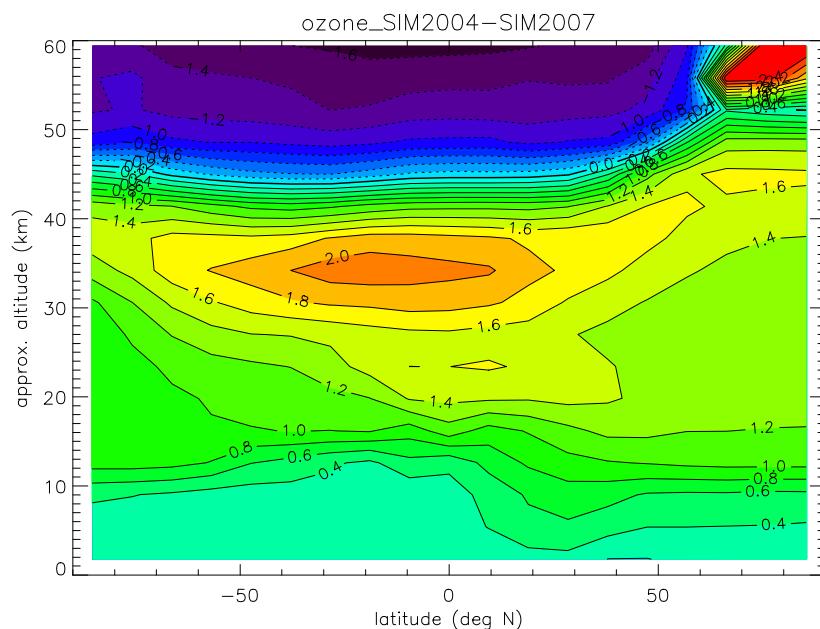
Haigh et al (Nature, 2010)

# 2D model O<sub>3</sub> differences (%) 2004-2007

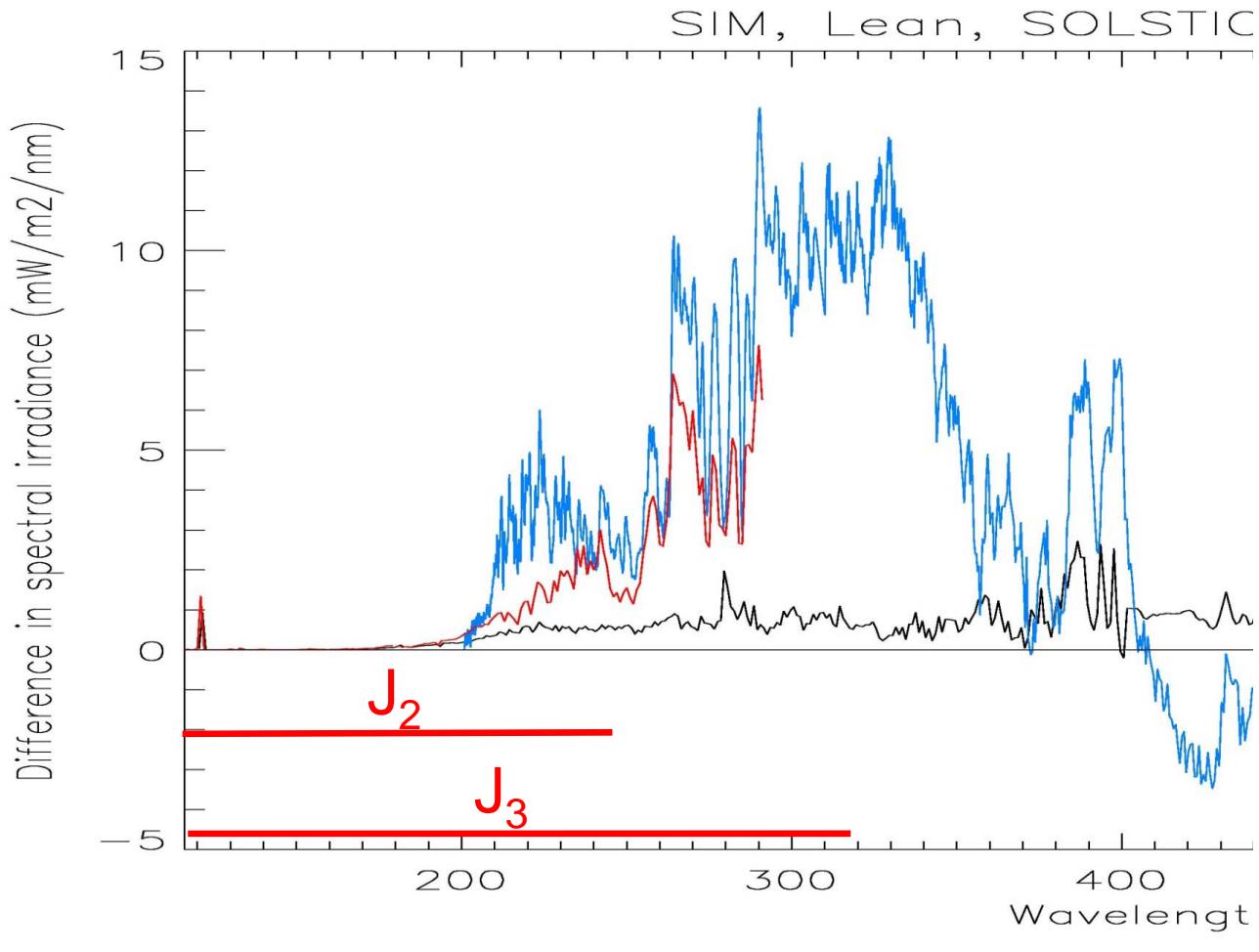
Lean



SIM



# Photodissociation



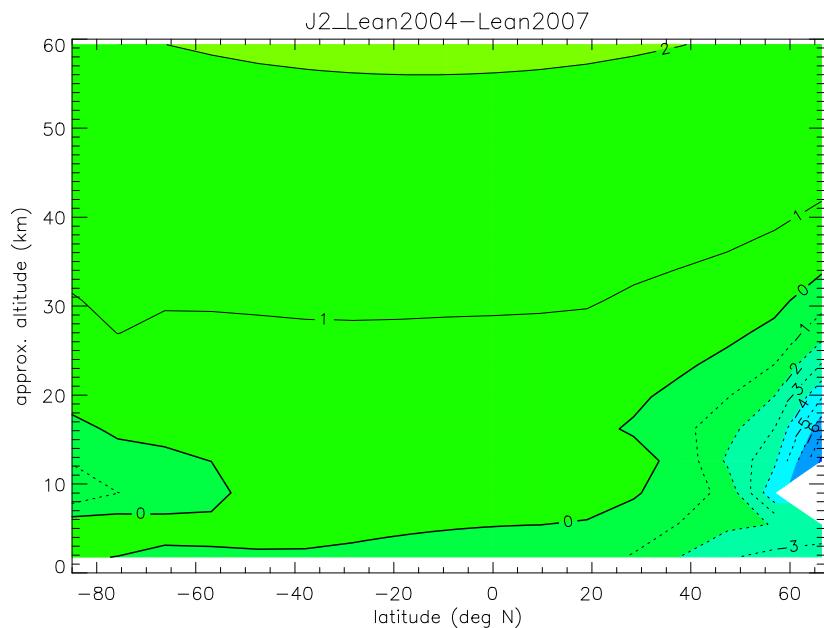
$J_2$  produces  
 $O$  and  $O_3$

$J_3$  produces  
 $O(^1D)$  and  
determines  
 $O/O_3$  ratio

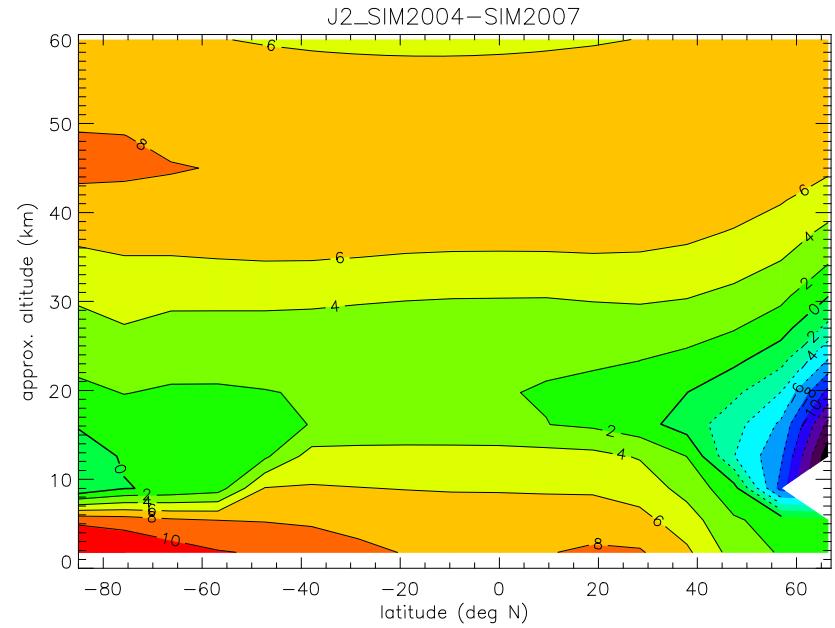
# $J_2$ : $O_2$ Photodissociation rate

2004 – 2007 (%)

Lean



SIM

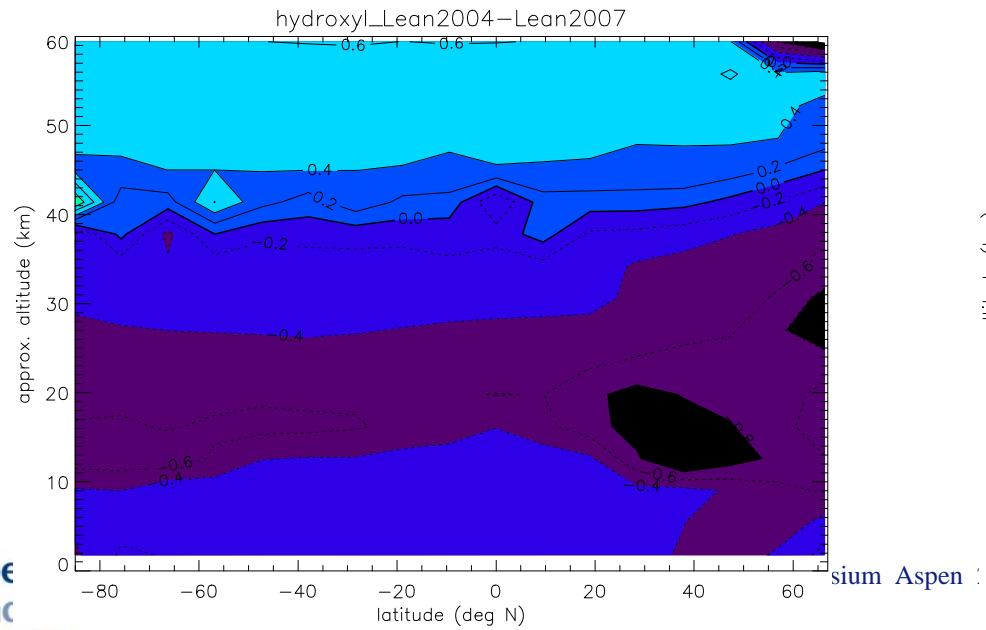


# OH

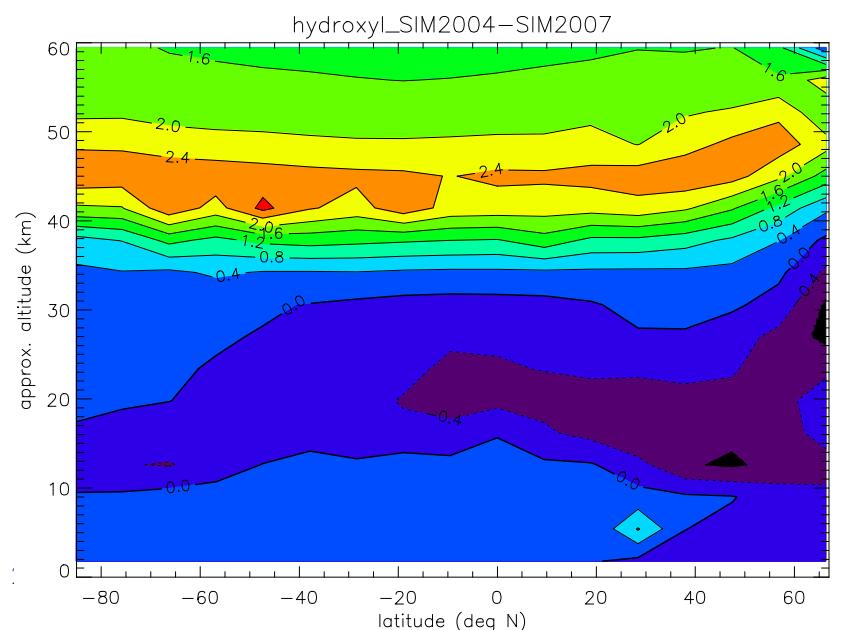
Produced by reaction of  $\text{H}_2\text{O}$  with  $\text{O}(\text{¹D})$

2004 – 2007 (%)

Lean



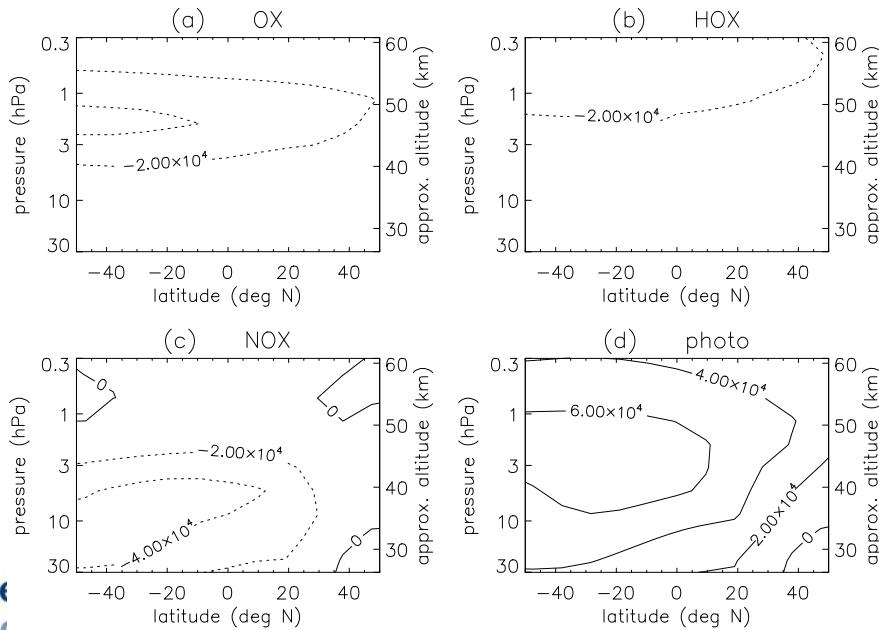
SIM



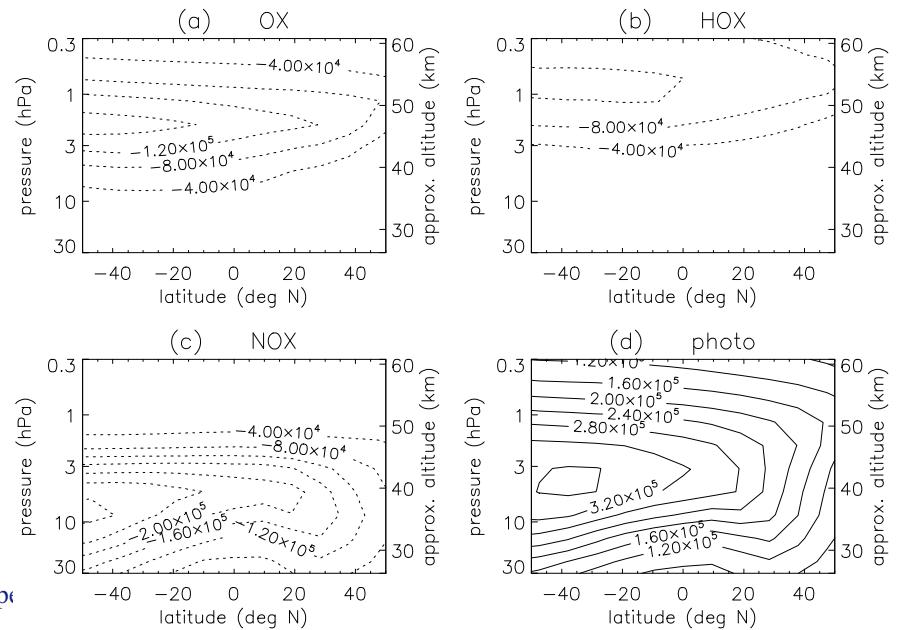
# O<sub>3</sub> production and destruction rates

- a) combination O+O<sub>3</sub>
- b) catalytic destruction by HO<sub>x</sub>
- c) catalytic destruction by NO<sub>x</sub>
- d) production by photodissociation of O<sub>2</sub>

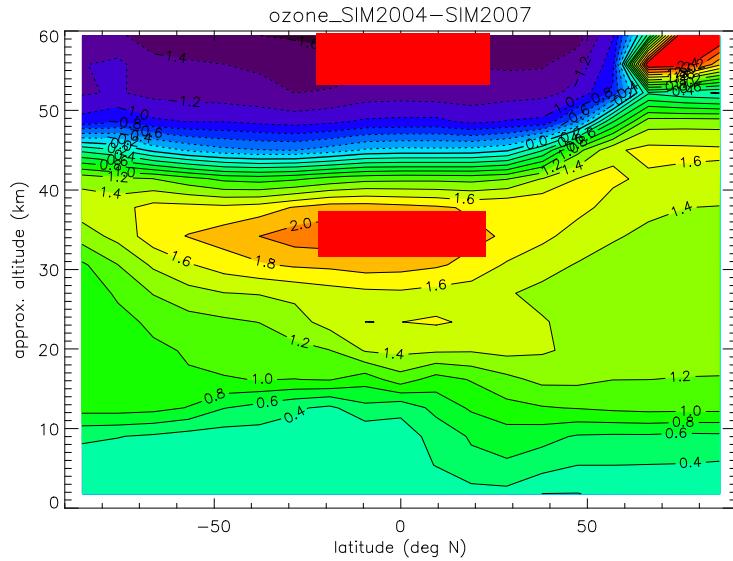
Lean



SIM



# Multiple regression of AURA MLS O<sub>3</sub> data

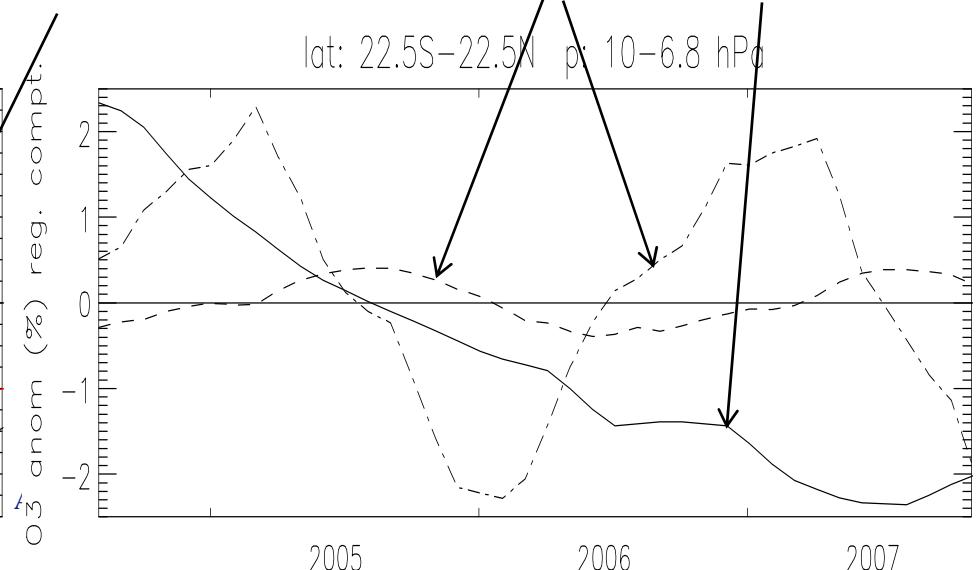
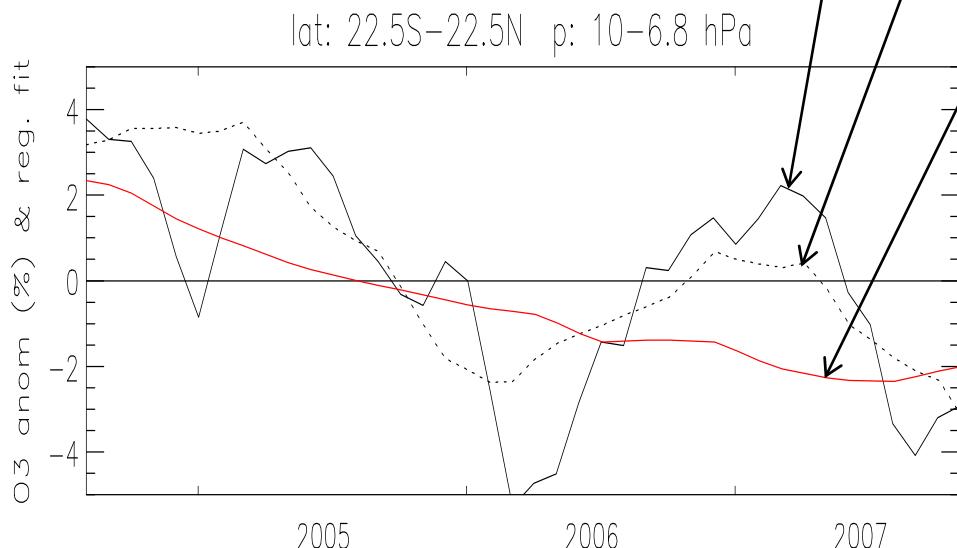


Regions chosen based on model results

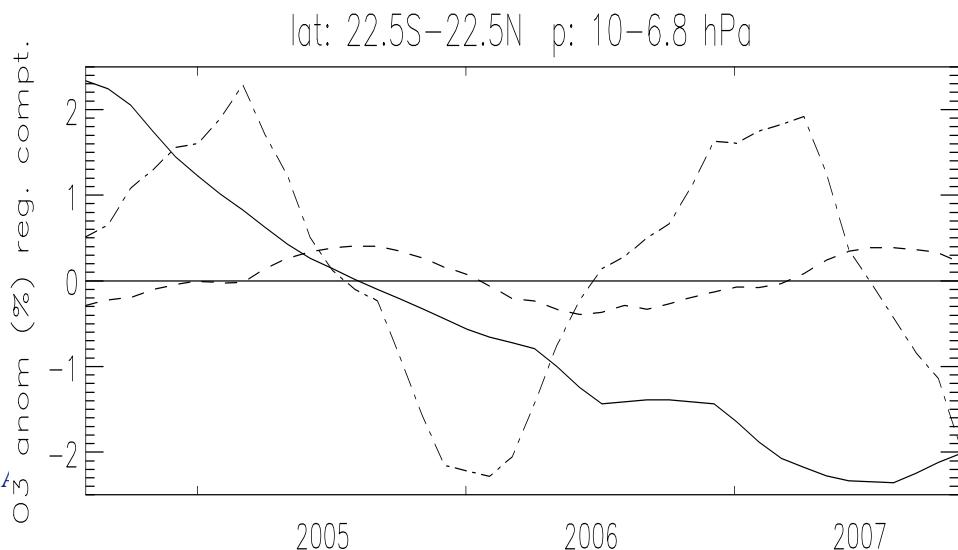
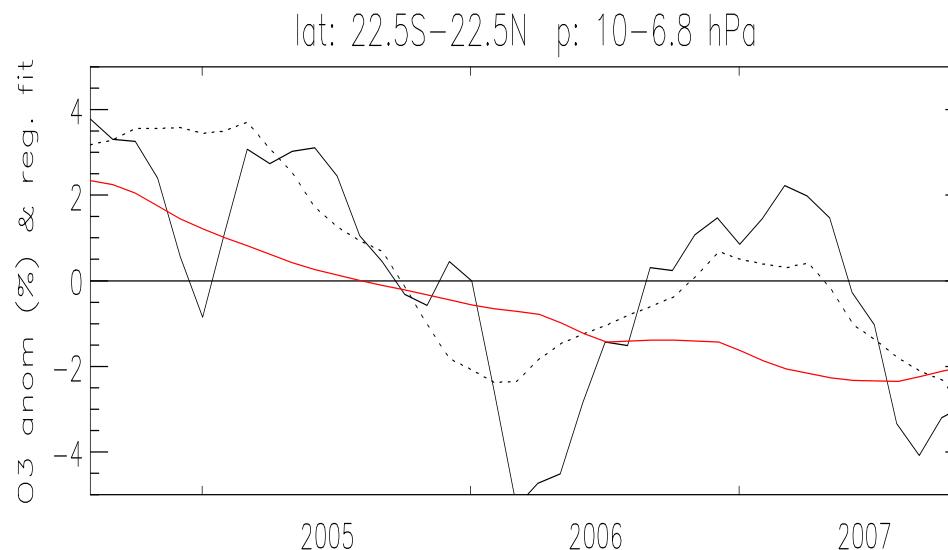
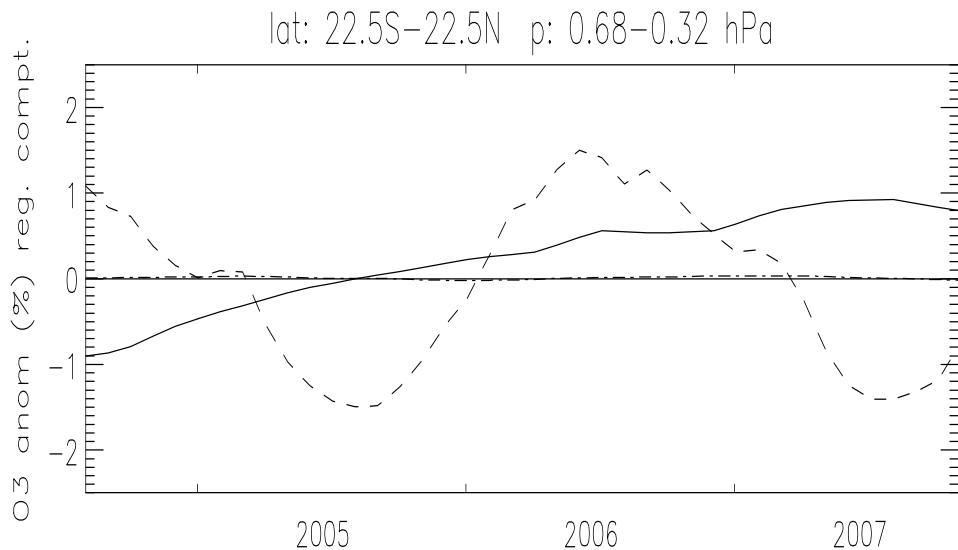
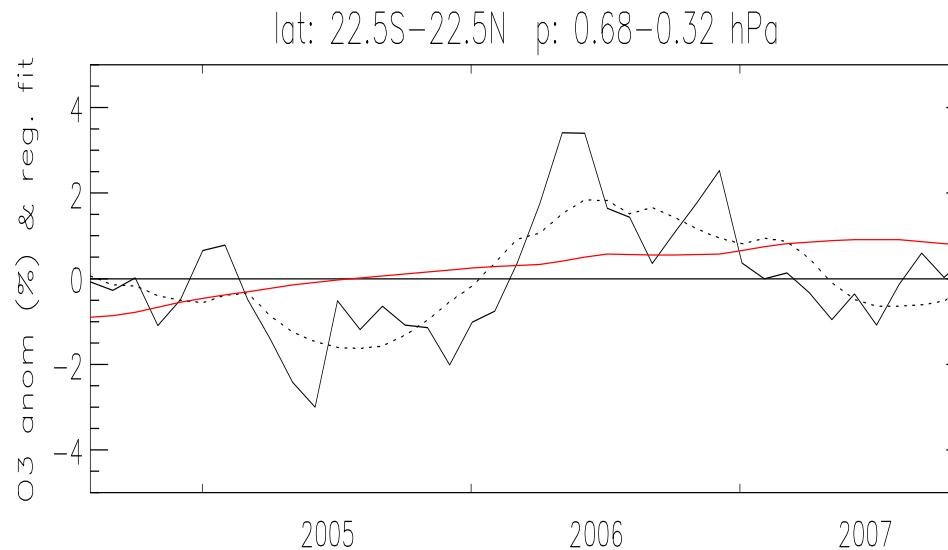
Raw data (deseasonalised monthly means)  
for lower layer

Regression fit

Regression indices:  
Solar? constant, 2 QBO, SIM UV

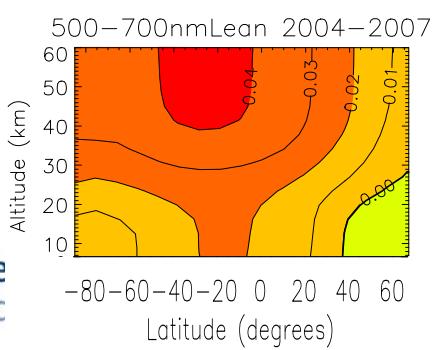
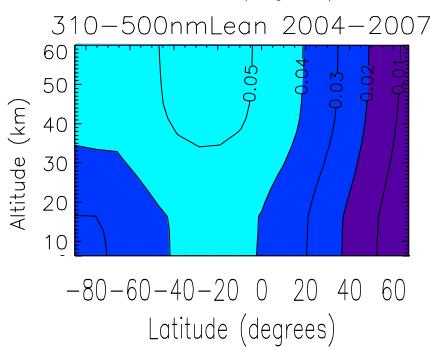
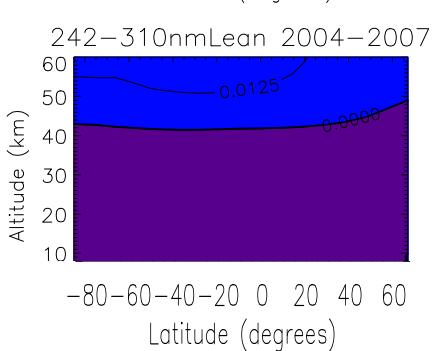
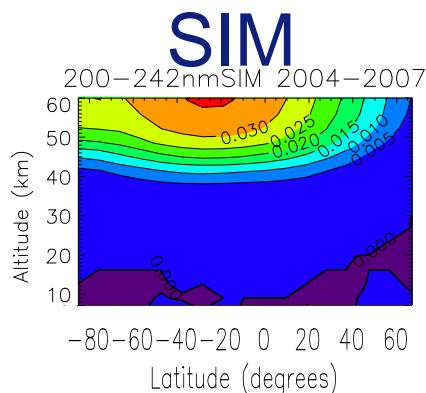


# Multiple regression of AURA MLS O<sub>3</sub> data

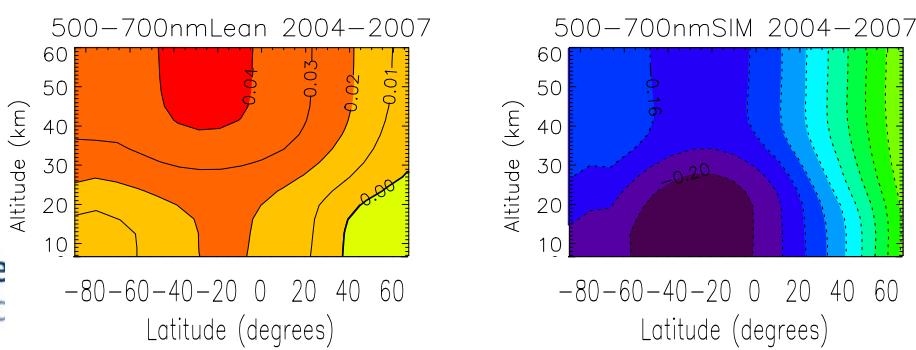
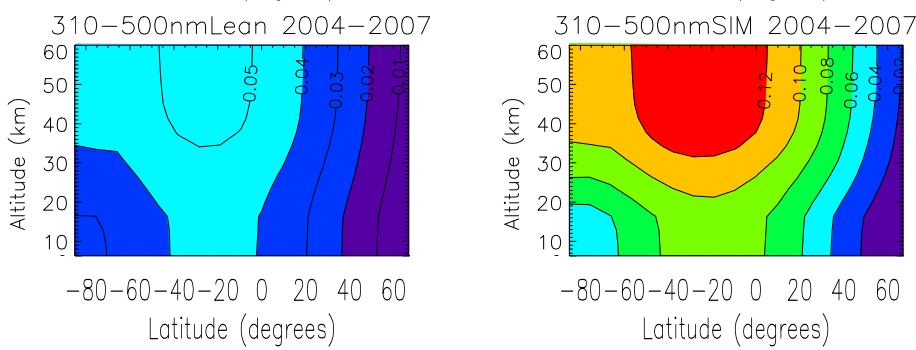
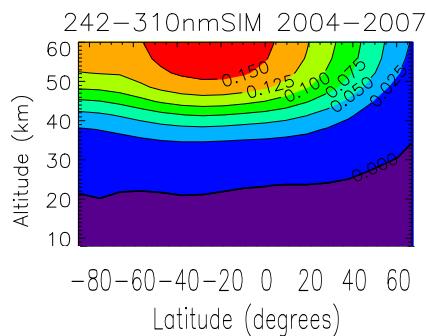


# Integrated solar flux 2004-2007 (Wm<sup>-2</sup>)

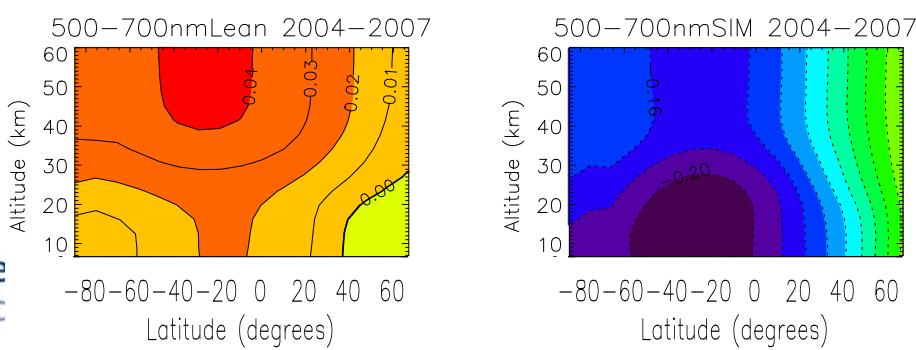
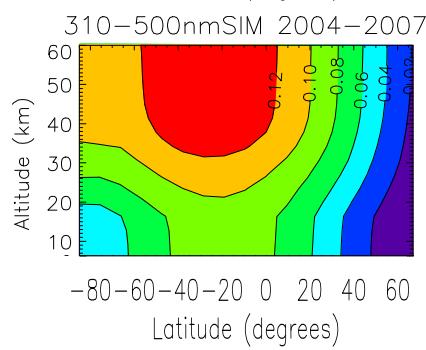
200-242 nm



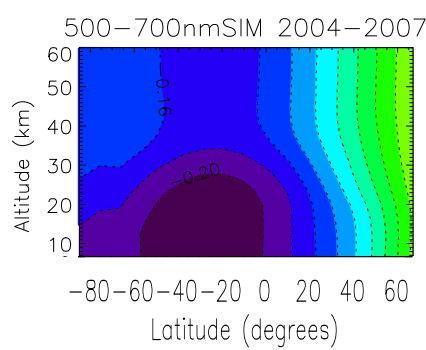
242-310 nm



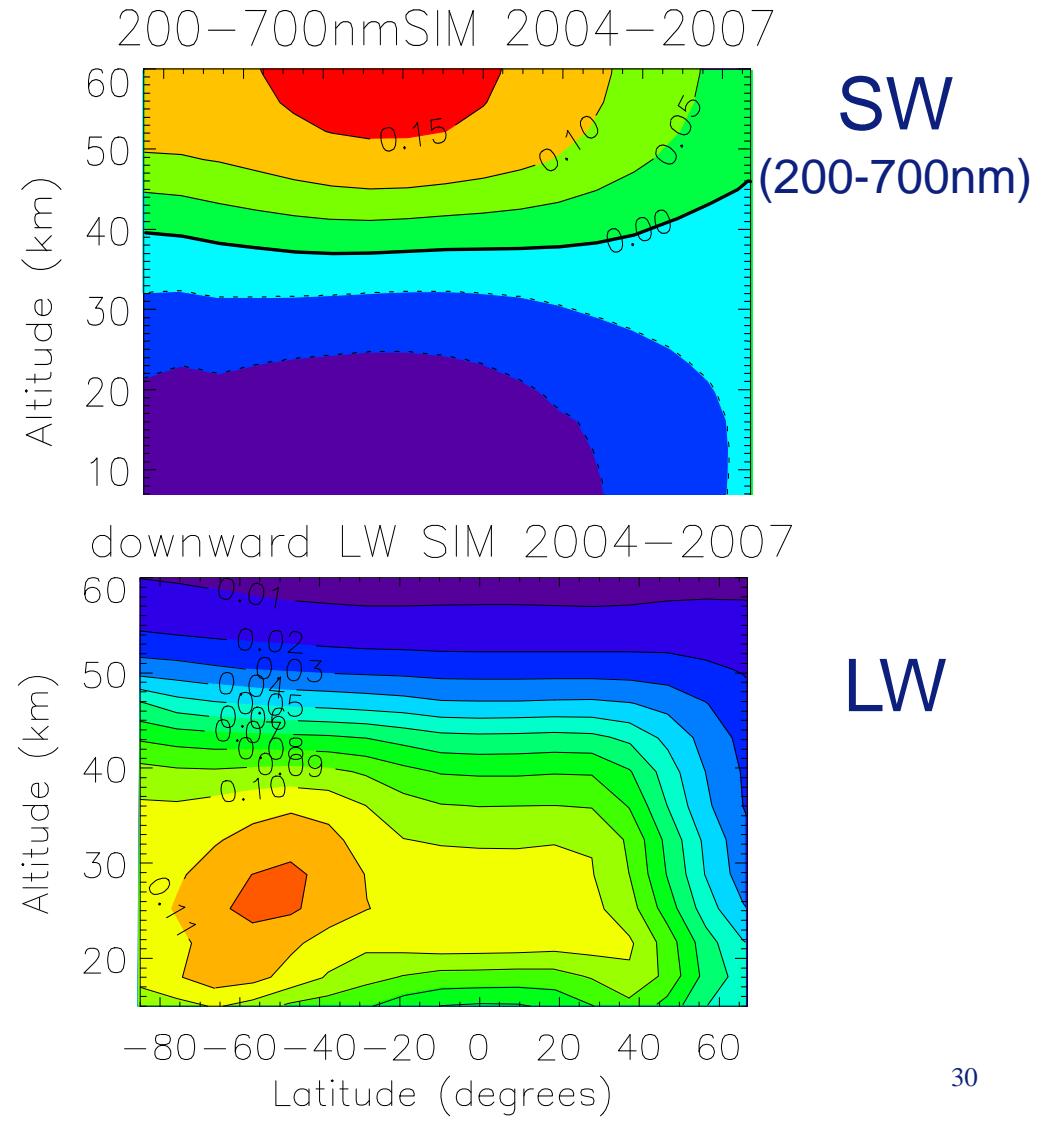
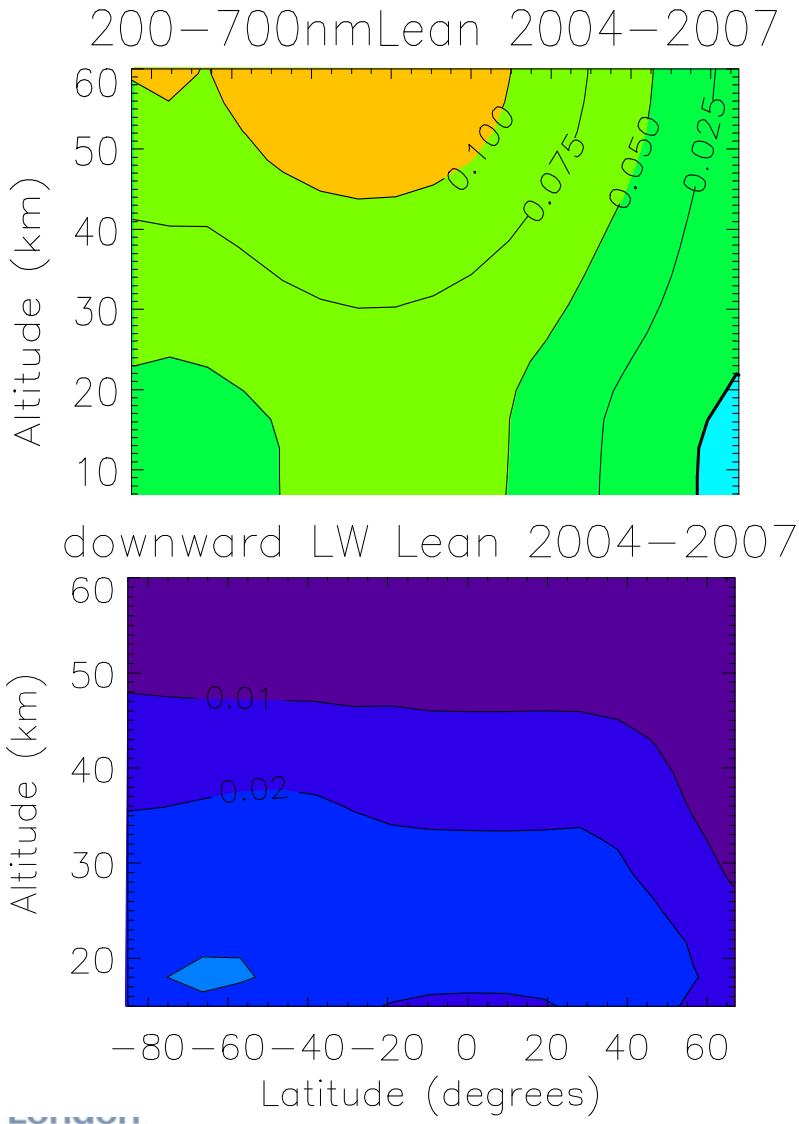
310-500 nm



500-700 nm



# Integrated radiative fluxes 2004-2007



# Solar Radiative Forcing of Climate\*

## 2004-2007 (mW m<sup>-2</sup>)

### Lean

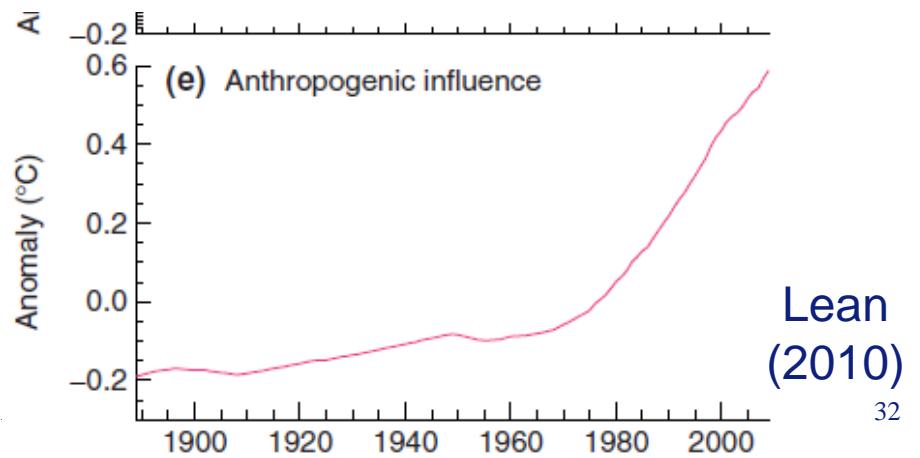
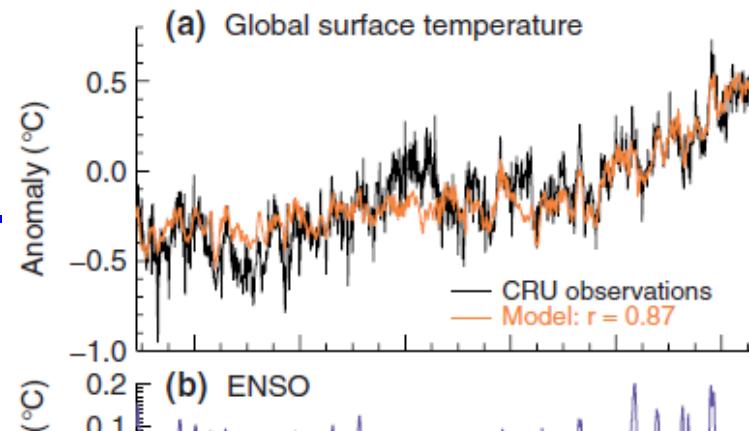
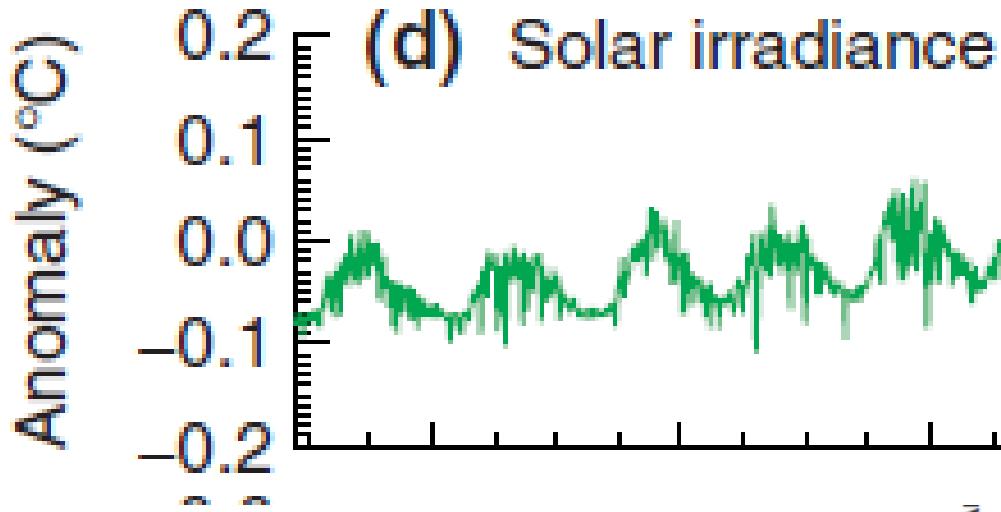
	200-310 nm	310-500 nm	500-700 nm	700-1600 nm	Total solar 200-1600 nm	Thermal (LW)	Net
TOA	20	40	30	20	110	0	110
105hPa	0	30	10	20	60	20	80

### SIM

	200-310 nm	310-500 nm	500-700 nm	700-1600 nm	Total solar 200-1600 nm	Thermal (LW)	Net
TOA	160	110	-130	-50	90	0	90
105hPa	0	60	-170	-50	-160	60	-100

# Global surface temperature

## Influences of various



# Solar forcing of climate: current understanding

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“Top-down” via UV heating the stratosphere

New solar data would make much larger

and/or

“Bottom-up” via visible radiation warming surface ?

New solar data would invert

# Summary: Part 2

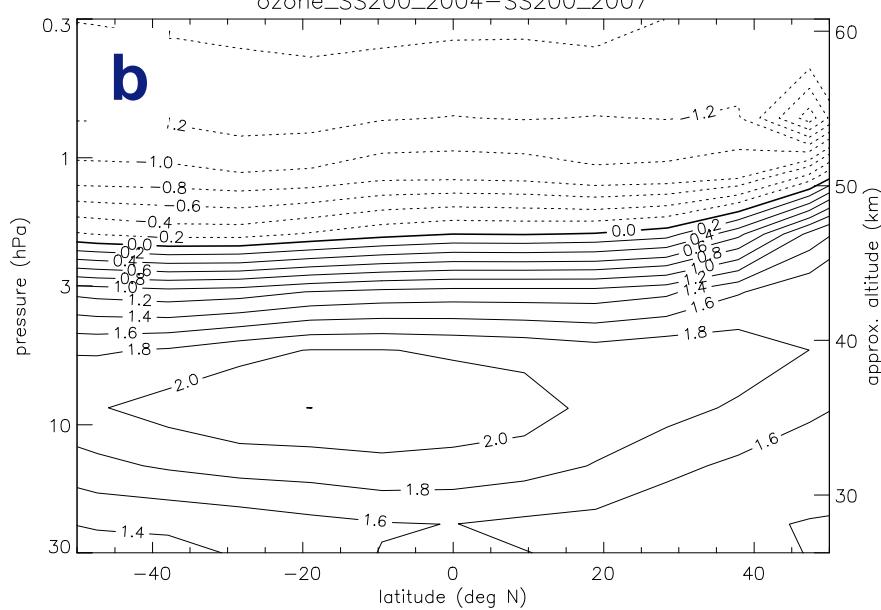
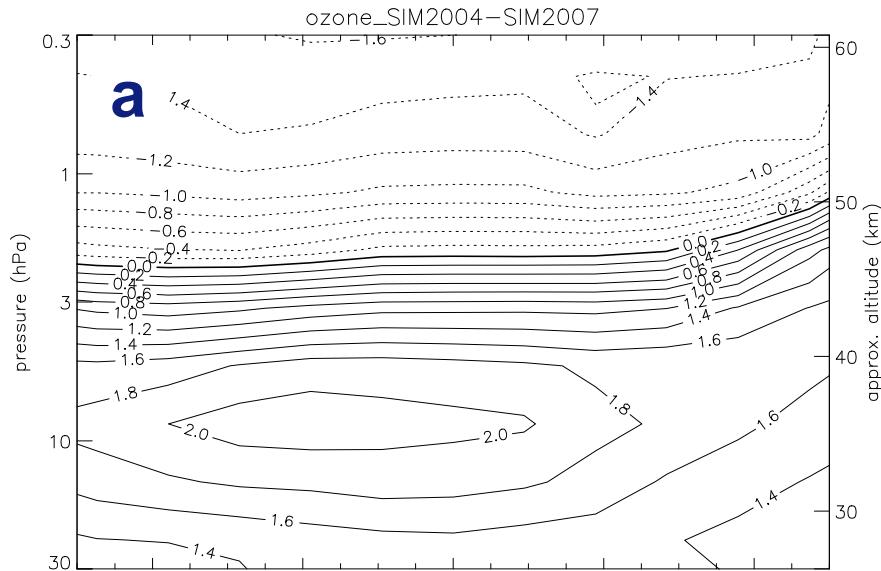
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- Input to 2D model SIM (& SOLSTICE) spectra produce a very different response in  $O_3$  from semi-empirical models of SSI: a *reduction* in lower mesosphere at higher solar activity and a large increase in mid- to upper stratosphere.
- This structure can be explained by enhanced production of  $HO_x$ , and by a shift of  $O_x$  from  $O_3$  to O.
- This structure is not inconsistent with contemporaneous measurements of  $O_3$  from AURA-MLS.
- SIM data would suggest that solar radiative forcing of climate produced a warming from 2004 to 2007, despite declining TSI.
- Is the Sun behaving oddly at present? Has this happened before?  
| What will the Sun do over the next few years...?

# SPARES

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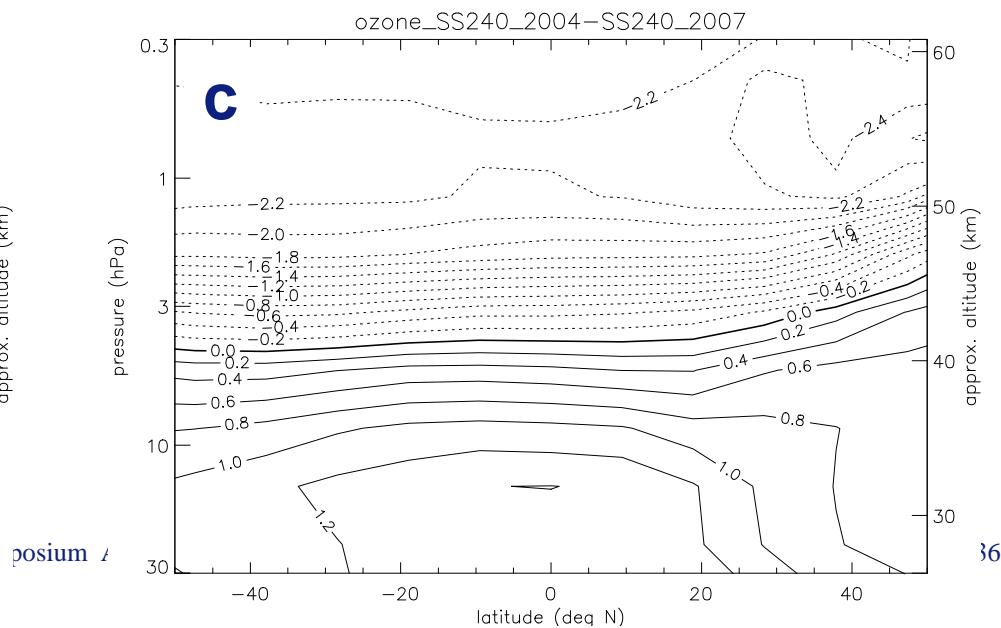
# Choice of spectra $\lambda < 240$ nm



$\lambda < 200$     $200 < \lambda < 240$     $\lambda > 240$

<b>a</b> Lean	<b>SIM</b>	<b>SIM</b>
<b>b</b> SOLSTICE	<b>SIM</b>	<b>SIM</b>
<b>c</b> SOLSTICE	<b>SOLSTICE</b>	<b>SIM</b>

Ozone difference (%) 2004-2007



possum  $\lambda$

36